





FASER: ForwArd Search ExpeRiment at the LHC

Sebastian Trojanowski University of Sheffield



University of Birmingham, May 15, 2019



arXiv: 1708.09389; 1710.09387; 1801.08947; 1806.02348 (PRD,with J.L.Feng, I.Galon, F.Kling)

FASER Collaboration: arXiv:1811:10243 Letter of Intent (CERN-LHCC-2018-030)

arXiv:1811.12522 Physics case

arXiv:1812.09139 Technical Proposal (CERN-LHCC-2018-036) arXiv:1901.04468 Input to the European Particle Physics Strategy

FASER APPROVAL

Voir en français

FASER: CERN approves new experiment to look for long-lived, exotic particles

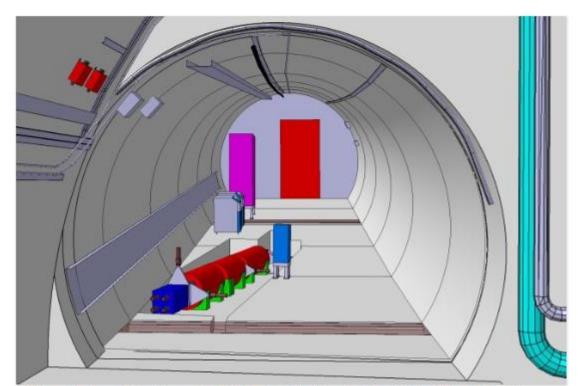
The experiment, which will complement existing searches for dark matter at the LHC, will be operational in 2021

5 MARCH, 2019 | By Cristina Agrigoroae

related article







A 3D picture of the planned FASER detector as seen in the Ti12 tunnel. The detector is precisely aligned with the collision axis in ATLAS, 480 m away from the collision point (image: FASER/CERN) (image: CERN)

Geneval Today the CERN Research Board approved a new experiment designed to look for light and weakly

(FASER group see https://twiki.cern.ch/twiki/bin/view/FASER)

FASER COLLABORATION

Spokespersons: J. Boyd, J. L. Feng

The FASER Collaboration: ~40 collaborators, 16 institutions, 8 countries
Henso Abreu (Technion), Claire Antel (Geneva), Akitaka Ariga (Bern), Tomoko Ariga (Kyushu/Bern), Jamie
Boyd (CERN), Dave Casper (UC Irvine), Franck Cadoux (Geneva), Xin Chen (Tsinghua), Andrea Coccaro
(Genova), Candan Dozen (Tsinghua China), Yannick Favre (Geneva), Jonathan Feng (UC Irvine), Didier
Ferrere (Geneva), Iftah Galon (Rutgers), Sergio Gonzalez-Sevilla (Geneva), Shih-Chieh Hsu (Washington),
Zhen Hu (Tsinghua), Peppe Iacobucci (Geneva), Sune Jakobsen (CERN), Roland Jansky (Geneva), Enrique
Kajomovitz (Technion), Felix Kling (UC Irvine), Susanne Kuehn (CERN), Lorne Levinson (Weizmann), Sam
Meehan (CERN), Josh McFayden (CERN), Friedemann Neuhaus (Mainz), Hidetoshi Otono (Kyushu),
Lorenzo Paolozzi (Geneva), Brian Petersen (CERN), Osamu Sato (Nagoya), Matthias Schott (Mainz), Anna
Sfyrla (Geneva), Savannah Shively (UC Irvine), Jordan Smolinsky (UC Irvine), Aaron Soffa (UC Irvine),
Yosuke Takubo (KEK), Eric Torrence (Oregon), Sebastian Trojanowski (Sheffield), Gang Zhang (Tsinghua
China



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KYUSHU

















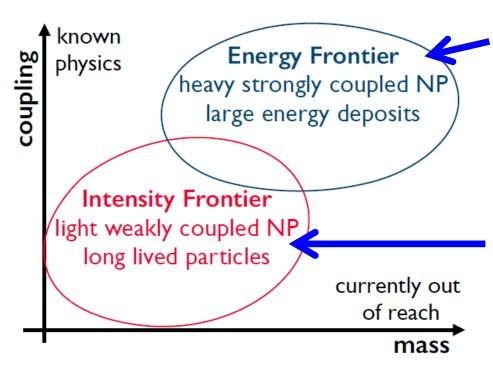




OUTLINE

- Motivation behind the intensity frontier searches for light long-lived particles (LLPs)
- FASER: ForwArd Search ExpeRiment at the LHC (idea, basic detector design)
- Remarks about FASER physics program
 - -- dark photons,
 - -- axion-like particles,
 - -- light scalars,
 - -- possible measurements for SM neutrinos,
 - -- ... and many other models
- Background: simulations & in-situ measurements
- Concluding remarks

MOTIVATION



<u>heavy</u> and <u>strongly-coupled new physics</u> e.g. SUSY, extra dimensions, ... here also missing energy searches for heavy WIMP DM, magnetic monopoles,...

Light and very weakly coupled new physics:

- -- requires large "luminosities" (statistics)
- -- new particles decay back to SM, but with highly displaced vertices
- -- SM BG needs to be highly suppressed

Exciting physics:

-- cosmology (dark matter, inflation, bariogenesis,...)

-- neutrino masses (GeV-scale heavy neutral leptons)

 $-(g-2)_{\mu}$

Standard Model

Dark sector

Dark Matter Light mediators:

dark photon, dark scalars, ...

Generalized WIMP miracle: $\Omega_{DM}h^2 \sim m^2/g^4 \sim 0.1$ g « $g_{weak} => m$ « m_{weak}

"The WIMPless Miracle..." J.L. Feng, J. Kumar, Phys.Rev.Lett. 101 (2008) 231301

HIDDEN SECTOR PORTALS

- new "hidden" particles are SM singlets
- interactions between the SM and "hidden" sector arise due to

mixing through some SM portal

$$\mathcal{L}_{\text{portal}} = \sum O_{\text{SM}} \times O_{\text{DS}}$$

B. Patt, F. Wilczek, 0605188

B. Batell, M. Pospelov, A. Ritz, 0906.5614

Renormalizable portals

Portal

Coupling

Dark Photon,
$$A_{\mu}$$

Dark Higgs,
$$S$$

Axion,
$$a$$

Sterile Neutrino,
$$N$$

$$-\frac{\epsilon}{2\cos\theta_W}F'_{\mu\nu}B^{\mu\nu}$$

$$(\mu S + \lambda S^2)H^{\dagger}H$$

$$\frac{a}{f_a}F_{\mu\nu}\tilde{F}^{\mu\nu}, \ \frac{a}{f_a}G_{i,\mu\nu}\tilde{G}_i^{\mu\nu}, \ \frac{\partial_{\mu}a}{f_a}\overline{\psi}\gamma^{\mu}\gamma^5\psi$$

$$y_N LHN$$

FASER

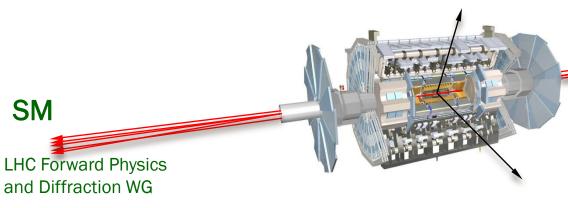
FASER - IDEA

FASER – newly proposed, small ($\sim 0.05 \text{ m}^3$) and inexpensive ($\sim 1\text{M}$ \$) experiment detector to be placed few hundred meters downstream away from the ATLAS IP

to harness large, currently "wasted" forward LHC cross section

 $\sigma_{\rm inel} \sim 75$ mb, e.g., $N_{\pi} \sim 10^{17}$ at 3 ab⁻¹

(for comparison $\sigma \sim fb - pb$, e.g., $N_{H} \sim 10^{7}$ at 300 fb⁻¹ in high-p_⊤ searches)



physics (LLP decays) π, K, D, B, ...

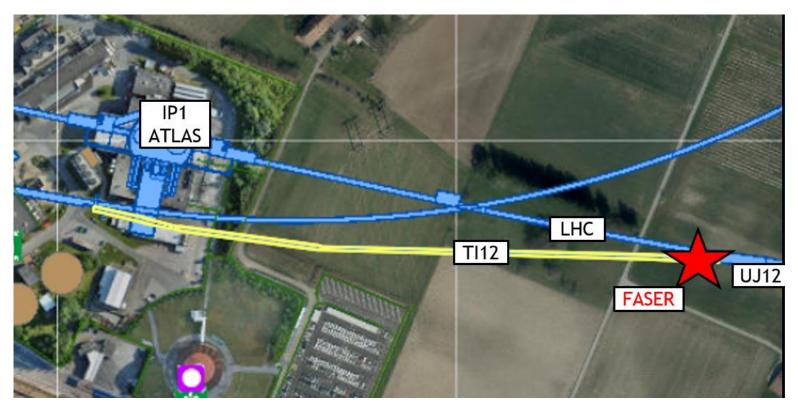
p-p collision axis

VERY SCHEMATICALLY **FASER** ATLAS IP

FASER will complement ATLAS/CMS by searching for highly-displaced decays of new Light Long-Lived Particles

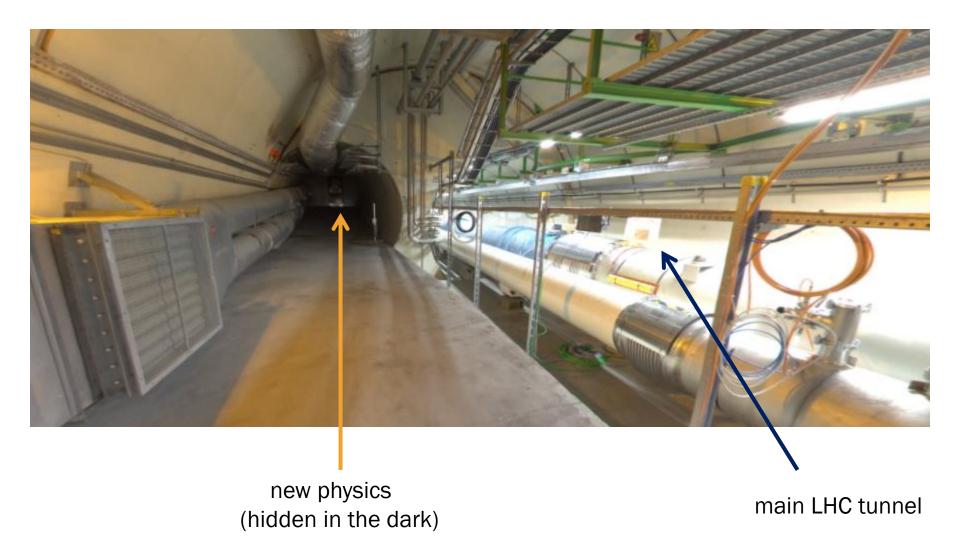
> (part of Physics Beyond Colliders Study Group at CERN)

FASER LOCATION – TUNNEL TI12



- location in a side tunnel TI12 (former service tunnel connecting SPS to LEP)
- L ~ 485m away from the IP along the beam axis
- space for a **5-meter-long** detector
- precise position of the beam axis in the tunnel up to mm precision (CERN Engineering Dep)
- corrections due to beam crossing angle (for $\sim 300 \mu rad$ the displacement is $\sim 7-8$ cm) 9

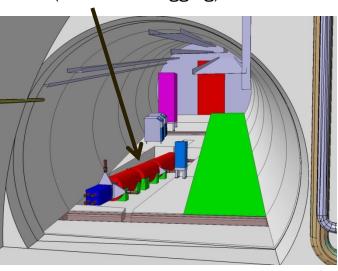
TUNNEL TI12



BASIC DETECTOR LAYOUT

1100.00 mm Tracking stations 3 planes of silicon strip detector per station Scintillator/Pb Veto to veto incoming charged 100.00 mm particles and protons Trigger/preshower scintillator station Electromagnetic 0.6 Tesla permanent calorimeter Trigger/timing new physics dipole magnets (Lead/scintillator) scintillator station with 20 cm aperture particle cylindrical decay volume beam axis

small civil engineering (max 50cm digging)



Thank you !!!

Recycling existing spare modules:

- ATLAS SCT modules (Tracker)
- LHCb ECAL modules (Calorimeter)

• 2 stages of the project:

FASER 1: L = 1.5 m, R = 10 cm, $V = 0.05 \text{ m}^3$, 150 fb⁻¹ (Run 3) (above layout)

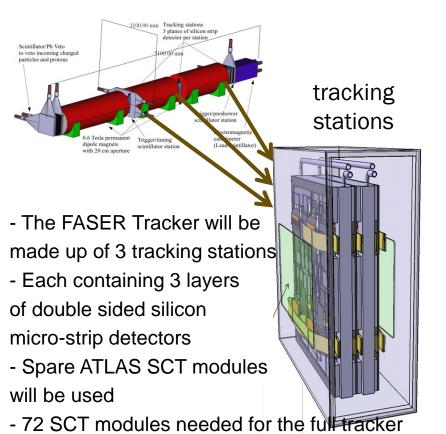
FASER 2: L = 5 m, R = 1 m, V = 16 m³, 3 ab⁻¹ (HL-LHC)

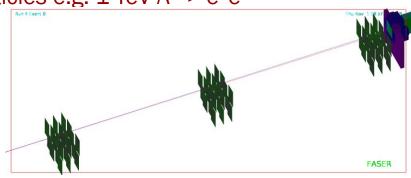
SIGNAL DETECTION IN THE TRACKER

Signal is a pair of oppositely charged high-energy particles e.g. 1 TeV A' -> e⁺e⁻

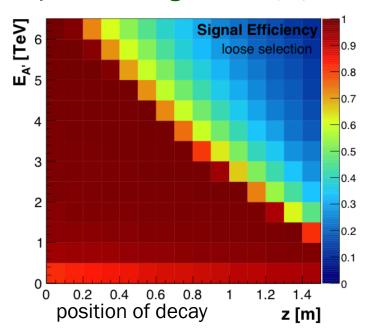
In the following we assume 100% detection efficiency for a better comparison with other experiments

Ongoing work on full detector simulations



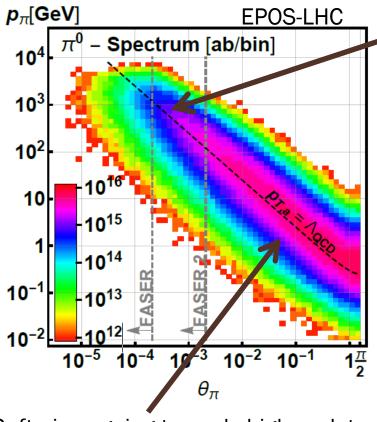


CHARGED TRACK SEPARATION EFFICIENCY **2nd/3rd tracking station** (separation > 0.3mm)



FASER PHYSICS

EXAMPLE OF LHC/FASER KINEMATICS LLP FROM PION PRODUCTION AT THE IP



Hard pions highly collimated along the beam axis since their p_T ~ $\Lambda_{\rm QCD}~$ e.g. for E_{\pi 0} $\geq 10~GeV$

- $\sim 1.7\%$ of π_0 s go towards FASER
- ~ 24% of $\pi_0 s$ go towards FASER 2

This can be compared to the angular size of both detectors with respect to the total solid angle of the forward hemisphere (2π) :

- $\sim (2 \times 10^{-6})\%$ for FASER
- $\sim (2 \times 10^{-4})\%$ for FASER 2

ATLAS π_0 new particle FASEF

Soft pions going towards high-p_T detectors:

- produced LLPs would be too soft for triggers
- large SM backgrounds

LLPs produced from B mesons in FASER 2

 $p_T \sim m_B$ \longrightarrow larger angular spread \longrightarrow target for FASER 2

at FASER energies: $N_B/N_{\pi} \sim 10^{-2}$

(10⁻⁷ for typical beam dumps) 14

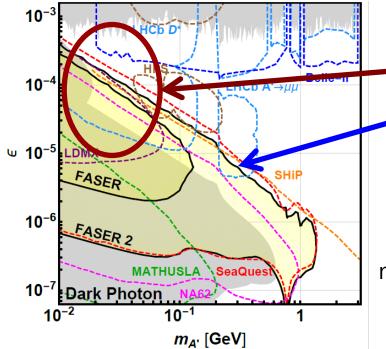
1708.09389, PRD 97 (2018) no.3, 035001

DARK PHOTON

- (broken) dark U(1) gauge group,
- kinetic mixing with the SM photon: $\epsilon F^{\mu\nu} F'_{\mu\nu}$,
- after field redefinition:

$$\mathcal{L} \supset -\frac{1}{4} F^{\mu\nu} F_{\mu\nu} - \frac{1}{4} F'^{\mu\nu} F'_{\mu\nu} + \frac{1}{2} \frac{m_{A'}^2}{M_{A'}^2} A'_{\mu} A'^{\mu} + \sum \bar{f}(i \partial \!\!\!/ - \epsilon \, e q_f \, A') f$$

- production: π^0 and η decays, bremsstrahlung, direct production in $q\bar{q}$ scatterings
- decays: dominantly into e^+e^- and $\mu^+\mu^-$ up to ~ 500 MeV,



then various hadronic decay modes

$$\bar{d} = c \frac{1}{\Gamma_{A'}} \gamma_{A'} \beta_{A'} \approx (80 \text{ m}) B_e \left[\frac{10^{-5}}{\epsilon} \right]^2 \left[\frac{E_{A'}}{\text{TeV}} \right] \left[\frac{100 \text{ MeV}}{m_{A'}} \right]^2$$

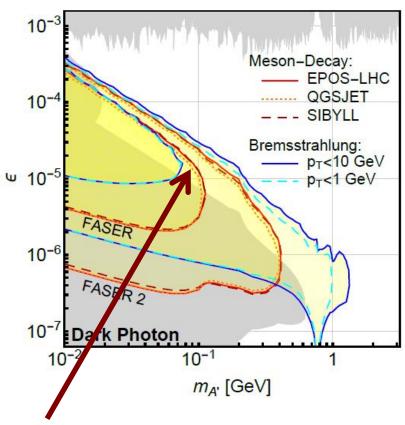
A' as a DM-SM mediator

FASER 2 comparable to proposed large SHiP detector

$$N_{\rm sig} \propto \mathcal{L}^{\rm int} \, \epsilon^2 \, e^{-L_{\rm min}/\bar{d}} \quad {\rm for} \ \ \bar{d} \ll L_{\rm min}$$
 $\bar{d} \sim \epsilon^2$

no of events grows exponentially with a small shift in $\boldsymbol{\epsilon}$

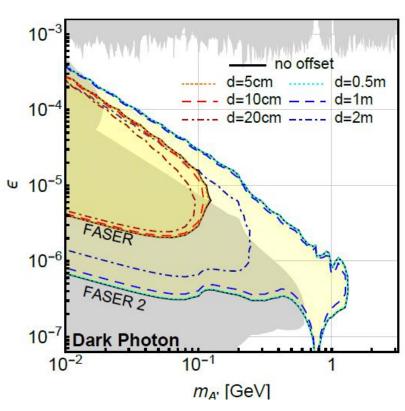
DARK PHOTON REACH – VARIOUS MC TOOLS & OFFSET



Almost impreceptible differences in reach for various MC tools

$$N_{
m sig} \propto {\cal L}^{
m int} \, \epsilon^2 \, e^{-L_{
m min}/ar d} \quad {
m for} \quad ar d \ll L_{
m min}$$

no of events grows exponentially with a small shift in ε



FASER reach unaffected by a small offset as long as the beam collision axis goes through the detector

1806.02348, PRD 98 (2018) no.5, 055021

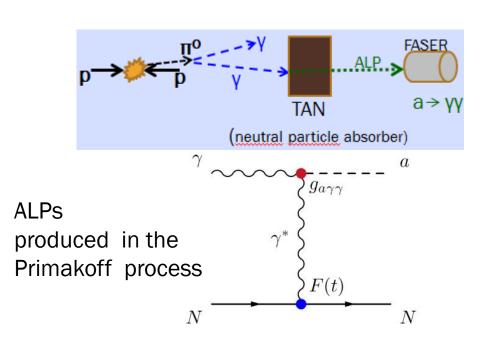
ALPS AT FASER -

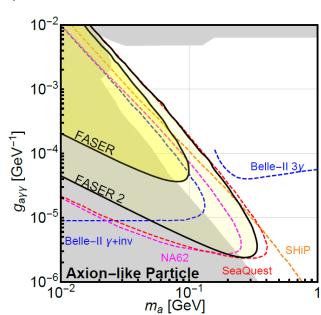
LHC AS A PHOTON BEAM DUMP

- similarly to the QCD axion, they can appear as pseudo-Nambu-Goldstone bosons in theories with broken global symmetries
- suppressed dim-5 couplings to gauge bosons $(1/\Lambda)aV^{\mu\nu}\tilde{V}_{\mu\nu}$,
- dim-5 couplings to fermions also allowed $(\partial_{\mu}a/\Lambda)f\gamma_{\mu}\gamma_{5}f$,
- interesting pheno scenario dominant $a\gamma\gamma$ coupling

B. Döbrich et al, JHEP 1602 (2016) 018

Photon beam dump (also "light shining through a wall")





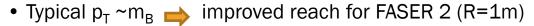
1710.09387, PRD 97 (2018) no.5, 055034

DARK HIGGS BOSONS

- Dark Higgs boson: additional hidden real scalar field ϕ ,
- often adopted phenomenological parametrization:

$$\mathcal{L} \supset -m_{\phi}^{2} \phi^{2} - \sin \theta \frac{m_{f}}{v} \phi \bar{f} f - \lambda v h \phi \phi$$

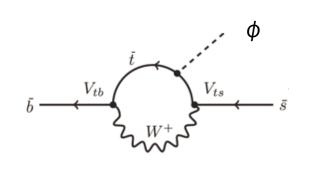
- Higgs-like couplings suppressed by θ^2 ,
- production: B and K decays, $h \to \phi \phi$,
- decays: into the heaviest kinematically allowed states: $\mu^+\mu^-$, $\pi\pi$, KK, . . .
- at FASER energies: $N_B/N_{\pi} \sim 10^{-2}$ (10⁻⁷ for typical beam+dumps)

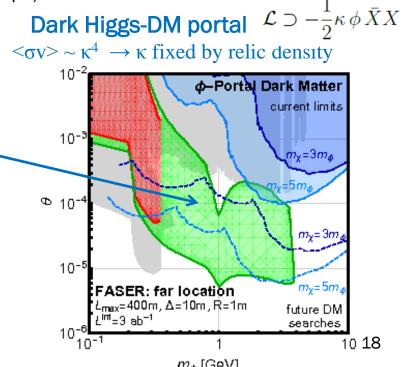


complementarity
between FASER
and other proposed
experiments
(large boost,
probing lower τ)

 m_{ϕ} [GeV]

Dark Higgs



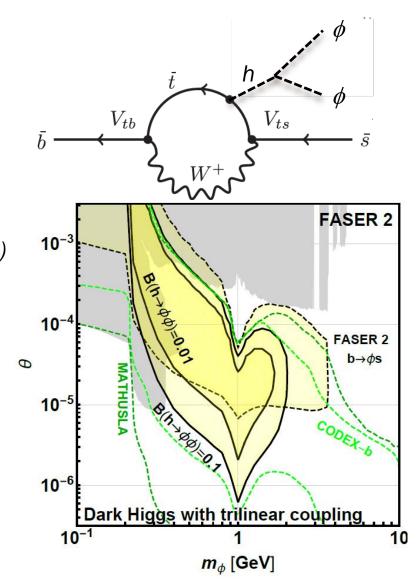


1710.09387, PRD 97 (2018) no.5, 055034

PROBING INVISIBLE DECAYS OF THE SM HIGGS

 $\mathcal{L} \supset -\frac{\lambda}{v}h\phi\phi$

- trilinear coupling
 - \implies invisible Higgs decays $h \rightarrow \phi \phi$
- far-forward region: efficient production via off-shell Higgs, $B \to X_s h^*(\to \phi \phi)$
- can extend the reach in θ up to 10^{-6} for B(h $\rightarrow \phi \phi$)~0.1
- up to ~100s of events



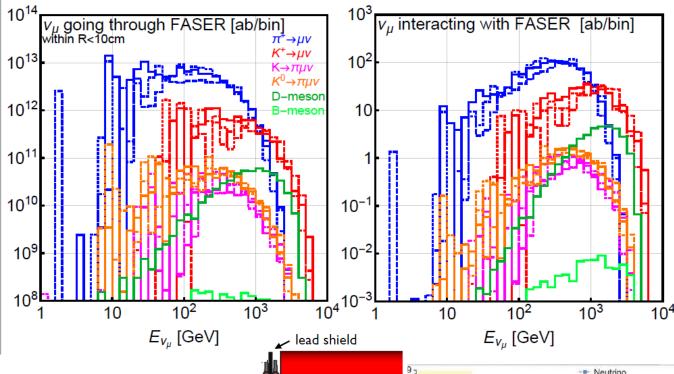
A. Ariga, T. Ariga, D. Casper, P. Denton, J. Feng, F. Kling, H. Otono, O. Sato, J. Smolinsky, ... (further work in progress)

SM NEUTRINOS IN FASER

Ideas currently explored:

1) Few cm thick lead plate will be put between several front veto layers for BG veto purposes (in front of FASER)

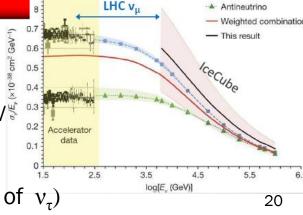
Incoming neutrinos can
CC interact inside the
lead plate producing muon
with no counterpart in
layers in front of the plate



Potentially hundreds of events in FASER

Measurement of the neutrino CC scattering cross section for E_{ν} ~TeV $^{\$}$

- 2) Employing larger neutrino detector in front of FASER
- additional information about kinematics (e.g. measurements) of v_{τ})



SM BACKGROUNDS

BACKGROUNDS - SIMULATIONS (FLUKA)

Spectacular signal:

- A' e+ e-

- -- two opposite-sign, high energy (few hundred GeV) charged tracks,
- -- that originate from a common vertex inside the decay volume,
- -- and point back to the IP (+no associated signal in a veto layer in front of FASER),
- -- and are consistent with bunch crossing timing.
- Neutrino-induced events: low rate
- The radiation level in TI18 is low (<10⁻² Gy/year), encouraging for detector electronics
- Proton showers in a nearby
 Disperssion Suppresor lead to negligible BG after ~90m of rocks in front of FASER

Other particles: detailed simulations, highly reduced rate (shielding + LHC magnets)

study by the members of the CERN FLUKA team:

	Cut	T > 100 GeV	Cut T > 500 GeV		Cut T > 1 TeV	
Part. type	fluence rate (cm ⁻² s ⁻¹)	fluence per bunch crossing per cm ²	fluence rate (cm ⁻² s ⁻¹)	fluence per bunch crossing per cm ²	fluence rate (cm ⁻² s ⁻¹)	fluence per bunch crossing per cm ²
μ+	0.18	6.1·10 ⁻⁹	0.02	5.8·10 ⁻¹⁰	0.002	6.8-10 ⁻¹¹
μ-	0.40	1.3.10-8	0.22	7.4-10-9	0.14	4.6·10 ⁻⁹
n _o	~ 10-7	~ 10-14	0	0	0	0
γ	~ 10-4	~ 10 ⁻¹²	~ 10-6	~ 10 ⁻¹³	~ 10-6	~ 10 ⁻¹³
π	~ 10-5	~ 10 ⁻¹²	~ 10-7	~ 10-14	0	0

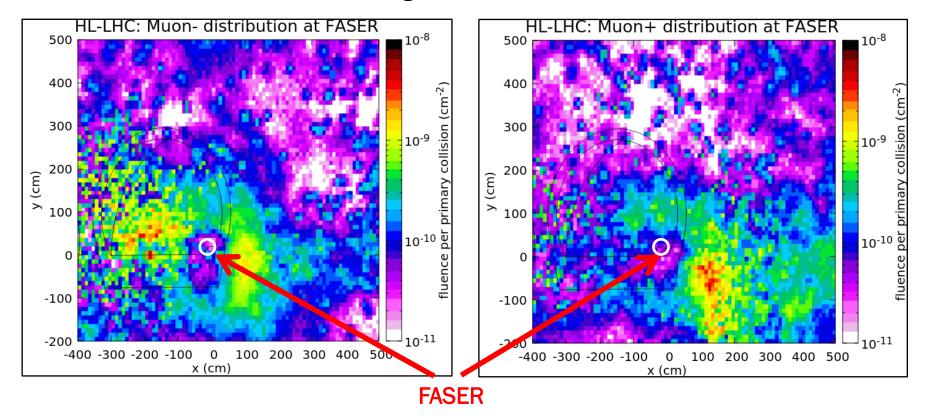
Muons coming from the IP – front veto layers

Expected trigger rate ~650 Hz

Process	Expected Number of Events
μ	540M
$\mu + \gamma_{\text{brem}}$	41K
$[\mu + (\gamma_{\rm brem} \to e^+e^-)]$	[7.4K]
$\mu + \text{EM shower}$	22K
μ + hadronic shower	21K

BACKGROUNDS - SIMULATIONS (2)

Cross section of the tunnel containing FASER



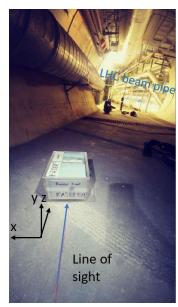
At FASER location:

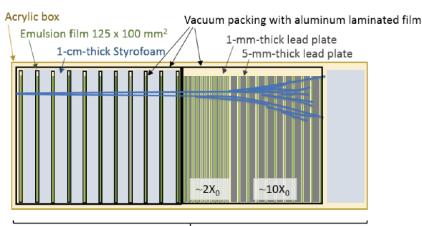
muon flux reduced along the beam collision axis (helpful role of the LHC magnets)

BACKGROUNDS - IN-SITU MEASUREMENTS

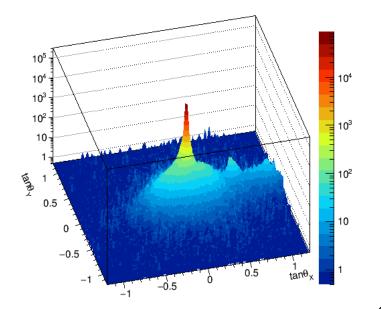
- Emulsion detectors –
 focusing on a small region around the beam axis (FASER location)
- TimePix Beam Lumi Monitors
- BatMons (battery-operated radiation monitors)

Analyses show that results are consistent with FLUKA simulations





200 mm



PRACTICALLY ZERO BG SEARCH

SUMMARY

FASER TIMELINE

Sep 2017: First paper, J. Feng, I. Galon, F. Kling, ST, PRD 97 035001 (2018)

...within ~1.5 year FASER grew to an international collaboration recognized at CERN

Currently: ~40 active members from ~16 institutions in ~8 countries,

Spokespersons: Jamie Boyd (CERN), Jonathan L. Feng (UC Irvine)

During LHC Run 2 (2018): detailed BG simulations (CERN Eng Dep) + in-situ measurements

Sep 2018: FASER Letter of Intent – accepted by the LHC Committee

Dec 2018: Technical Proposal recommended by the LHC Committee for a full approval

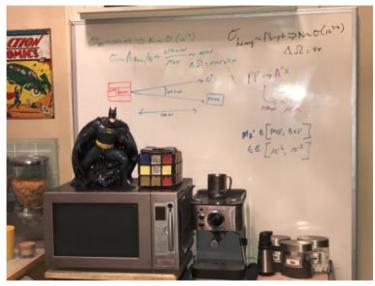
Dec 2018/Jan 2019: fundings granted for the detector (Heisig-Simons and Simons foundations)

Mar 2019: FASER fully approved by the CERN Research Board

PLANS:

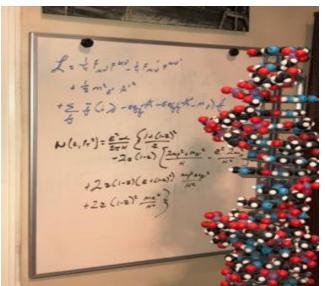
- -- Final detector design, manufacture, installation and commisioning during Long Shutdown 2 (ongoing work)
- Data taking during LHC Run 3 (2021-23)
- FASER 2 (major upgrade for HL-LHC)

FASER IN POPULAR CULTURE









related article



CONCLUSIONS

• Light Long-lived Particles (LLPs) – exciting new physics !!!

New physics reach even after first 10fb-1 (end of 2021?)

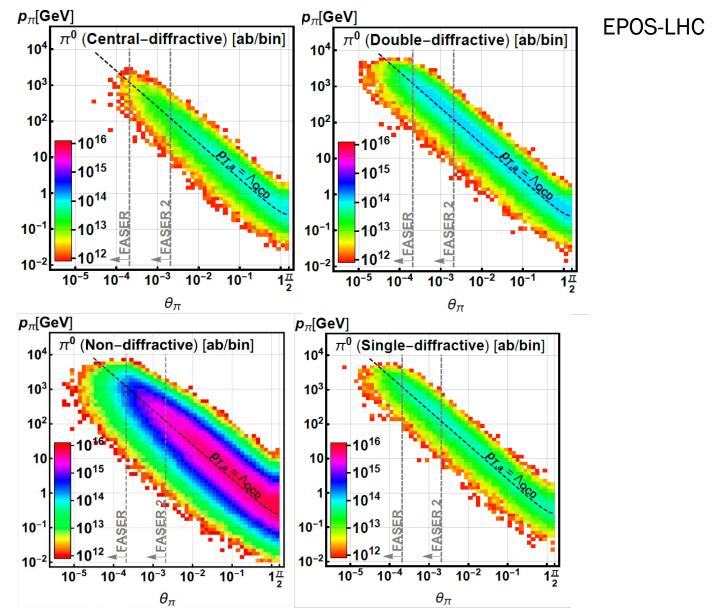
- FASER is a newly proposed, <u>small and inexpensive</u> experiment to be placed at the LHC to search for light long-lived particles to complement the existing experimental programs at the LHC, as well as other proposed experiments,
- FASER is fully approved by the CERN Research Board
- FASER would not affect any of the existing LHC programs and do not have to compete with them for the beam time etc.
- Rich physics prospects:
- popular LLP models (dark photon, dark Higgs boson, GeV-scale HNLs, ALPs...),
- Many connections to DM and cosmology
- Invisible decays of the SM Higgs,
- Measurments of SM neutrinos
- Possible timeline:

Install FASER 1 in LS2 (2019-20) for Run 3 (150 fb⁻¹)

- R = 10 cm, L = 1.5 m, Target dark photons, B-L gauge bosons, ALPs, HNLs(τ)... Install FASER 2 in LS3 (2023-25) for HL-LHC (3 ab⁻¹)
- R = 1 m, L = 5 m, Full physics program: dark vectors, ALPs, dark Higgs, HNLs...

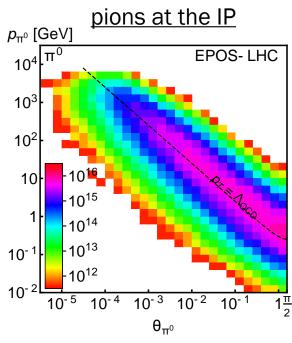
BACKUP

INELASTIC P-P COLLISIONS

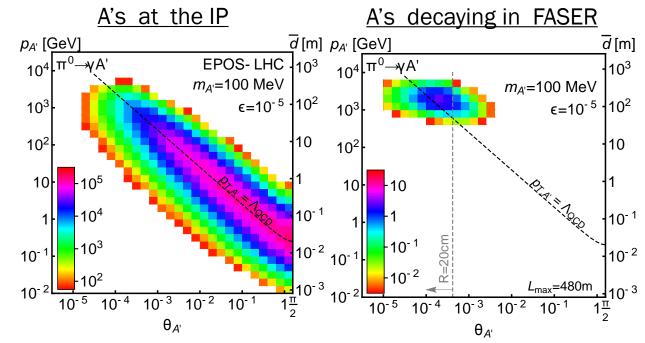


1708.09389, PRD 97 (2018) no.3, 035001

DARK PHOTONS AT FASER - KINEMATICS



- Monte Carlo fitted to experimental data (LHCf, ALFA)
- typically $p_T \sim \Lambda_{OCD}$
- for E~TeV \implies p_T/E ~0.1 mrad
- even ~ 10^{15} pions per (θ ,p) bin



- $\pi^0 \longrightarrow A'\gamma$
- high-energy π⁰
 ⇒ collimated A's
- $\epsilon^2 \sim 10^{-10}$ suppression but still up to 10^5 A's per bin

 only highly boosted A's survive until FASER

E_{∆′} ~TeV

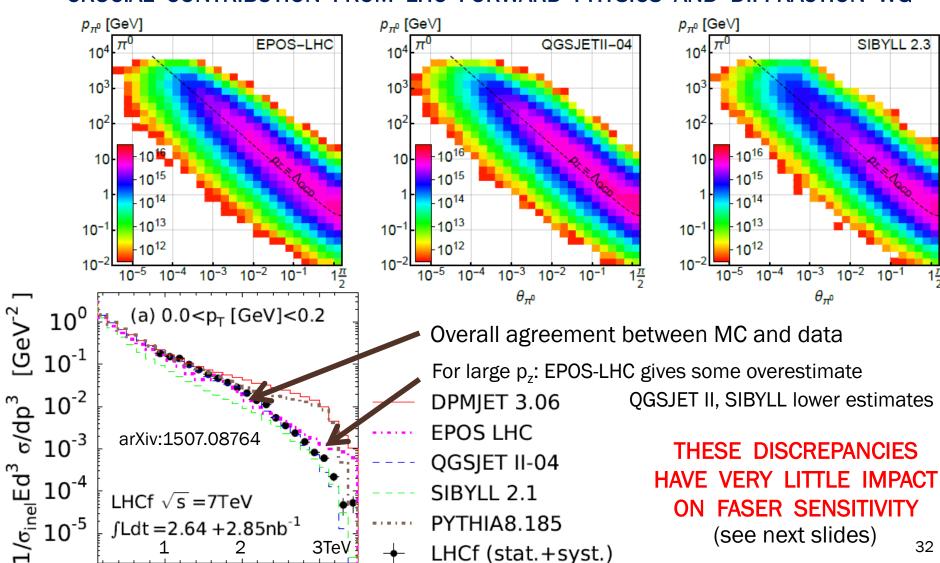
- further suppression from decay in volume probability
- still up to $N_{A'} \sim 100$ events in FASER,

mostly within **r<20cm**

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COMPARISON - VARIOUS MC TOOLS

CRUCIAL CONTRIBUTION FROM LHC FORWARD PHYSICS AND DIFFRACTION WG



HEAVY NEUTRAL LEPTONS AT FASER

1801.08947

Typical simplified approach:

- we focus on only one HNL leaving a signature in FASER
- we vary as free parameters

$$m_N$$
, U_{eN} , $U_{\mu N}$, $U_{\tau N}$, where only one $U_{\ell N} \neq 0$ at a time.

B and D meson decays – we consider about ~ 20 production channels, dominant ones dictated by the CKM suppression, kinematics and fragmentation fractions

$$D^{0,\pm} \to N e^{\pm} K^{\mp,0,(*)}, D_s^{\pm} \to N e^{\pm}, \dots$$

 $B^{0,\pm} \to N e^{\pm} D^{\mp,0,(*)}, B^{\pm} \to N e^{\pm}, \dots$
 $B_c^{\pm} \to N e^{\pm}, \dots$

Decay modes:

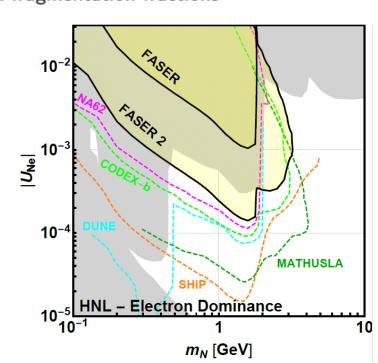
$$BR(N \rightarrow 3\nu) \sim 10\% - 20\%$$
 invisible

$${\rm BR}(N \to \nu \, l_1^+ \, l_2^-) \sim 20\% \ \ ({\rm BR}(N \to \nu \, e^+ \, e^-) \sim {\rm few \; percent})$$

 $BR(N \to hadrons) \sim 60\% - 70\%$, various final states

FASER 2

- \Rightarrow up to $\sim 10^3$ events for $m_N \gtrsim m_D$
- \Rightarrow for $m_N \lesssim m_D$ possible $\sim 10^1 \text{-} 10^2$ events



1801.08947, PRD 97 (2018) no.9, 095016

HEAVY NEUTRAL LEPTONS

• seesaw mechanism, e.g., for type-I seesaw

$$\mathcal{L} = \mathcal{L}_{\rm SM} + i \, \bar{\widetilde{N}}_I \partial \!\!\!/ \widetilde{N}_I - F_{\alpha I} \bar{L}_{\alpha} \, \widetilde{N}_I \, \tilde{\Phi} - \frac{1}{2} \bar{\widetilde{N}}_I^c \, M_I \, \widetilde{N}_I + \text{h.c.}$$

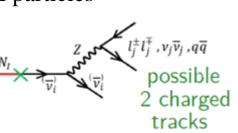
• once Higgs gets vev, they mix with active (SM) neutrinos Mixing angles: U_{eI} , U_{uI} , $U_{\tau I}$

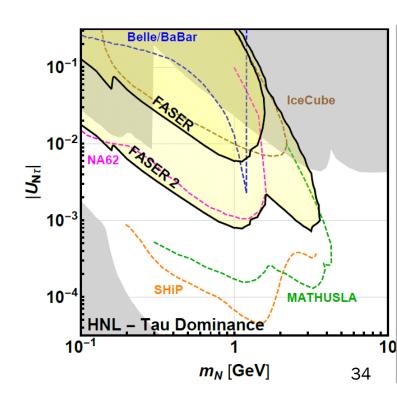
production in B and D meson decays

$$D^{0,\pm} \to N \ e^{\pm} \ K^{\mp,0,(*)}, \ D_s^{\pm} \to N \ e^{\pm}, \dots$$

 $B^{0,\pm} \to N \ e^{\pm} \ D^{\mp,0,(*)}, \ B^{\pm} \to N \ e^{\pm}, \dots$
 $B_c^{\pm} \to N \ e^{\pm}, \dots$

• decay back into lighter SM particles (visible BR often 80-90%)





1811.12522, (physics case)

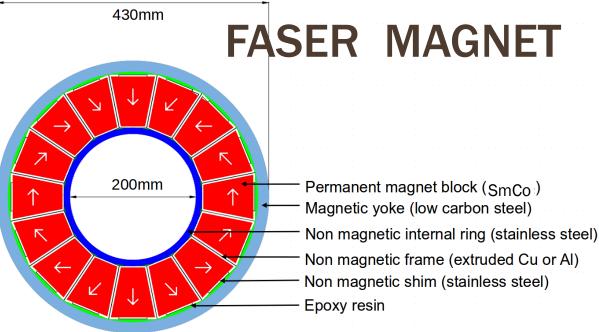
MORE MODELS OF NEW PHYSICS

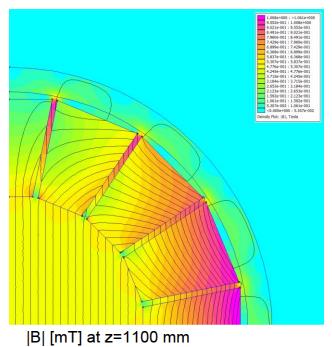
(table refers to the benchmark scenarios of the Physics Beyond Colliders CERN study group)

Benchmark Model	Label	Section	PBC	Refs	FASER	FASER 2
Dark Photons	V1	IV A	BC1	[7]		
B-L Gauge Bosons	V2	IVB		[30]	\checkmark	
$L_i - L_j$ Gauge Bosons	V3	IVC		[30]		
Dark Higgs Bosons	S1	VA	BC4	[26, 27]		
Dark Higgs Bosons with hSS	S2	VВ	BC5	[26]		$\sqrt{}$
HNLs with e	F1	VI	BC6	[28, 29]		
HNLs with μ	F2	VI	BC7	[28, 29]		$\sqrt{}$
HNLs with τ	F3	VI	BC8	[28, 29]	$\sqrt{}$	$\sqrt{}$
ALPs with Photon	A1	VIIA	BC9	[32]		
ALPs with Fermion	A2	VIIB	BC10		\checkmark	$\sqrt{}$
ALPs with Gluon	A3	VIIC	BC11		$\sqrt{}$	\checkmark

Other models & FASER sensitivity studies e.g.:

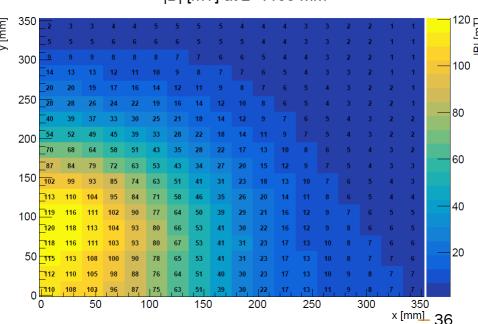
- RPV SUSY (D. Drecks, J. de Vries, H.K. Dreiner, Z.S. Wang, 1810.03617)
- Inelastic dark matter (A. Berlin, F. Kling, 1810.01879)





- 0.55T permanent dipole magnets based on the Halbach array design
 - LOS to pass through the magnet center
 - minimum digging to the floor in TI12
 - minimized needed services (power, cooling)
- manufacture: CERN magnet group
- stray field around scintillator PMTs ~5mT

shielding (mu-metal)



calorimeter

(Lead/scintillator)

00.00 mm Tracking stations 3 planes of silicon strip

Scintillator/Pb Veto to veto incoming charged

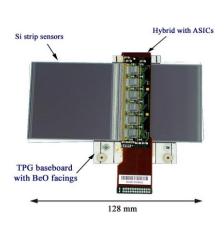
dipole magnets

FASER TRACKING STATIONS

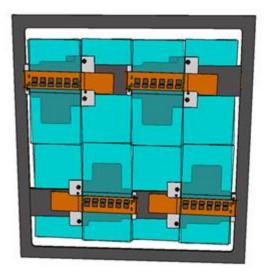
- The FASER Tracker will be made up of 3 tracking stations
- Each containing 3 layers of double sided silicon micro-strip detectors.
- Spare ATLAS SCT modules will be used
 - 80µm strip pitch, 40mrad stereo angle
 - Many thanks to the ATLAS SCT collaboration!
- 72 SCT modules needed for the full tracker
- Due to the low radiation in TI12 the silicon can be operated at room temperature, but the detector needs to be cooled to remove heat from the on-detector ASICs

Tracker readout using FPGA based board from University of Geneva (already used in

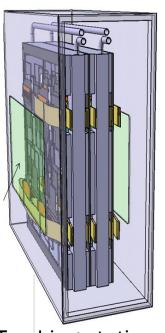
Baby MIND neutrino experiment)



SCT module



Tracking layer



Tracking station

CALORIMETER & SCINTILLATORS Stintliner Th V20 to twen benoring charged particles and protons Trigger/inning stilliner and protons Trigger/inning stilliner and protons Outter Module Outter Module

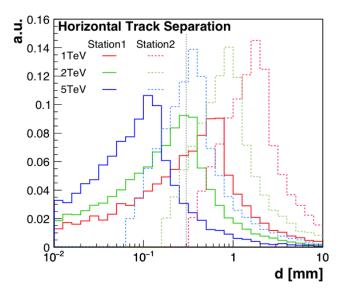
FASER will have an ECAL:

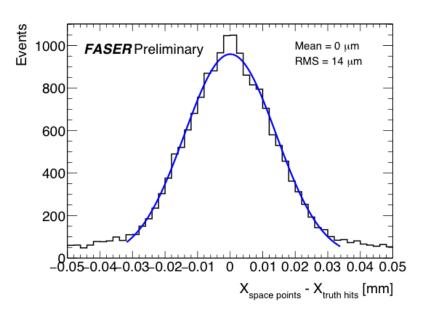
measuring the EM energy in the event (up to 1% accuracy in energy ~1 TeV)

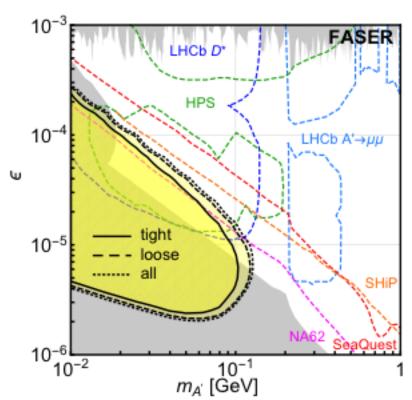
- Will use 4 spare LHCb outer ECAL modules
 - Many thanks to LHCb Collaboration for allowing us to use these!
 - 66 layers of lead/scintillator (2mm lead, 4mm plastic scintillator)
 - 25 radiation lengths long
 - no longitudinal shower information
 - Resolution will degrade at higher energy due to not containing full shower in calorimeter
- Scintillators used for vetoing charged particles entering the decay volume, for triggering and as a preshower
 - To be produced at CERN scintillator lab
 - Vetoing: achievable extremely efficient charged particle veto (eff>99.99%)
 - Trigger: also timing the signal with respect to timing of the \$pp\$ interactions
 - Preshower: thin radiator in front, photon showering (disentangling from v interactions in ECAL)⁸

MORE ABOUT TRACK SEPARATION

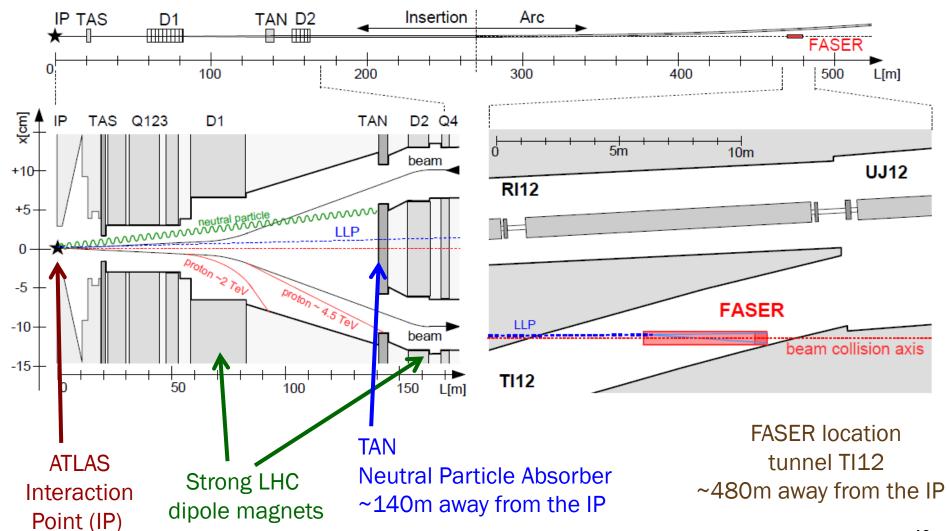








FASER AND SURROUNDING LHC INFRASTRUCTURE



POSSIBLE LOCATIONS (TI12 vs TI18)

- When designing the detector 2 main possible locations were considered:
- tunnels TI12 and TI18 on two sides of the ATLAS IP (~480m away from the IP)
- Both are former service tunnels connecting SPS and the main LHC tunnel
- Both are currently unused
- Both slope steeply upwards when leaving the main LHC tunnel (SPS is shallower than LHC)
- In both cases the line-of-sight (along the beam collision axis)
 is below the tunnel floor as it enters the tunnel, and then emerges from the floor
- Lowering of the floor up to 460mm is possible to maximize the detector length

(CERN survey team)

- The tunnels do have identical geometry: about 5m long detector can be fit in tunnel TI12 about 3m long detector can be fit in tunnel TI18
- Based on this the preferred location is the tunnel TI12
- BG measurements have been performed in both locations (below fluxes within 10 mrad)

	beam	observed tracks	efficiency	normalized flux, all	normalized flux, main peak	
	$[fb^{-1}]$	$[{\rm cm}^{-2}]$		$[\mathrm{fb}\ \mathrm{cm}^{-2}]$	$[\mathrm{fb}\ \mathrm{cm}^{-2}]$	
TI18	2.86	18407	0.25	$(2.6 \pm 0.7) \times 10^4$	$(1.2 \pm 0.4) \times 10^4$	
TI12	7.07	174208	0.80	$(3.0 \pm 0.3) \times 10^4$	$(1.9 \pm 0.2) \times 10^4$	
FLUKA simulation, E>100 GeV			GeV	1×10^{4}		