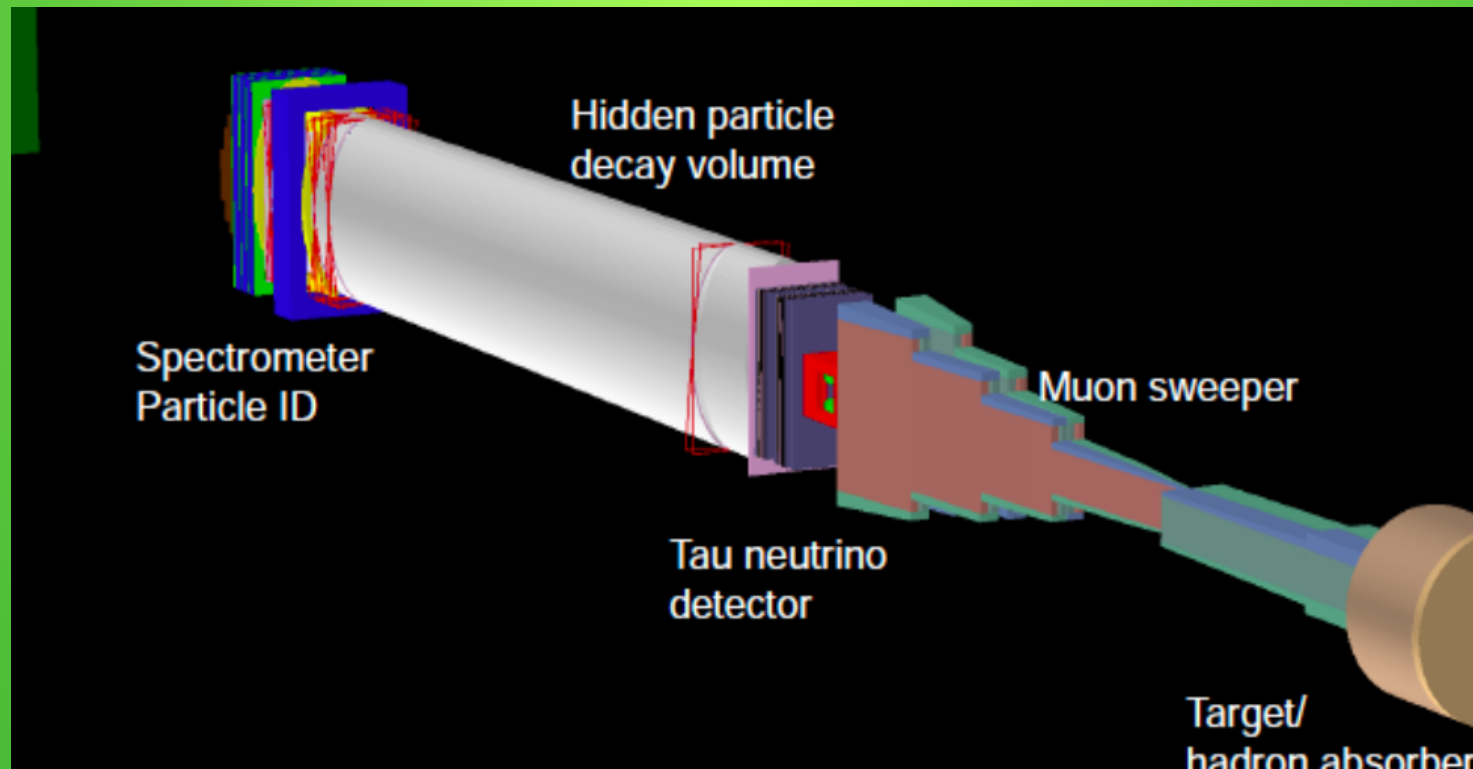


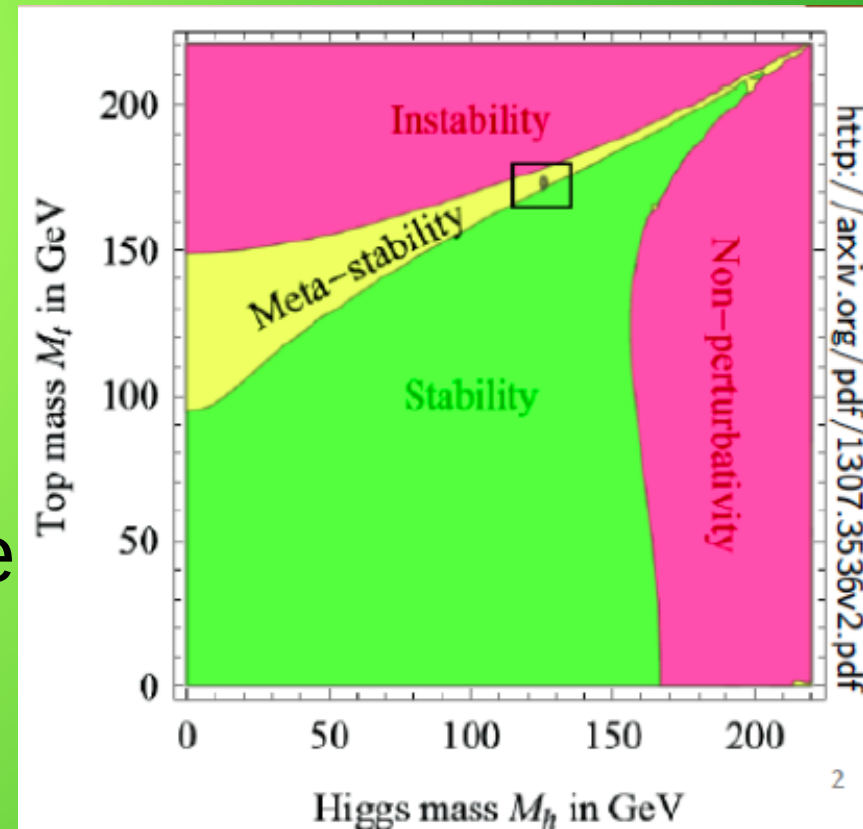
Search for Hidden Particles (ShiP): an experimental proposal at the SPS

ship.web.cern.ch/ship
Mario Campanelli
University College London



The Standard Model and beyond

- All SM particles have been discovered so far (apart from anti- ν_τ)
- Despite some anomalies, no compelling evidence of new physics found so far
- The Higgs mass points to a (meta-) stable universe
- The SM could be valid to the Planck scale
- Naturalness only a problem if we assume new particles between the EW and Planck scales



What we know we do not know

- Apart from naturalness, we do not understand:
 - Barion Asymmetry of the Universe
 - Dark Matter (indications are for cold, non-barionic)
 - The pattern of masses and mixings
 - Inflation
- Limits to masses of new particles being pushed in the TeV scale by the LHC.
 - “protection” against a small Higgs mass getting weaker

ATLAS limits for SUSY

ATLAS SUSY Searches* - 95% CL Lower Limits

October 2019

ATLAS Preliminary

$\sqrt{s} = 13$ TeV

Model	Signature	$\int \mathcal{L} dt$ [fb ⁻¹]	Mass limit	Reference					
Inclusive Searches	$q\bar{q}, \bar{q} \rightarrow q\bar{\chi}_1^0$	0 e, μ mono-jet	2-6 jets 1-3 jets	E_T^{miss} 139 E_T^{miss} 36.1	\bar{q} [10x Degen.] \bar{q} [1x, 8x Degen.]	1.9 0.43 0.71	$m(\tilde{\chi}_1^0) < 400$ GeV $m(\tilde{q}) - m(\tilde{\chi}_1^0) = 5$ GeV	ATLAS-CONF-2019-040 1711.03301	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0$	0 e, μ	2-6 jets	E_T^{miss} 139	\tilde{g} \tilde{g}	2.35 Forbidden 1.15-1.95	$m(\tilde{\chi}_1^0) = 0$ GeV $m(\tilde{\chi}_1^0) = 1000$ GeV	ATLAS-CONF-2019-040 ATLAS-CONF-2019-040	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}(\ell\ell)\tilde{\chi}_1^0$	3 e, μ $ee, \mu\mu$	4 jets 2 jets	E_T^{miss} 36.1 E_T^{miss} 36.1	\tilde{g} \tilde{g}	1.85 1.2	$m(\tilde{\chi}_1^0) < 800$ GeV $m(\tilde{g}) - m(\tilde{\chi}_1^0) = 50$ GeV	1706.03731 1805.11381	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}WZ\tilde{\chi}_1^0$	0 e, μ SS e, μ	7-11 jets 6 jets	E_T^{miss} 36.1 E_T^{miss} 139	\tilde{g} \tilde{g}	1.8 1.15	$m(\tilde{\chi}_1^0) < 400$ GeV $m(\tilde{g}) - m(\tilde{\chi}_1^0) = 200$ GeV	1708.02794 1909.08457	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$	0-1 e, μ SS e, μ	3 b 6 jets	E_T^{miss} 79.8 E_T^{miss} 139	\tilde{g} \tilde{g}	2.25 1.25	$m(\tilde{\chi}_1^0) < 200$ GeV $m(\tilde{g}) - m(\tilde{\chi}_1^0) = 300$ GeV	ATLAS-CONF-2018-041 ATLAS-CONF-2019-015	
	3 rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\bar{\chi}_1^0/\tilde{\chi}_1^\pm$	Multiple Multiple Multiple	36.1 36.1 139	E_T^{miss} E_T^{miss} E_T^{miss}	\tilde{b}_1 \tilde{b}_1 \tilde{b}_1	Forbidden Forbidden 0.58-0.82 Forbidden 0.74	$m(\tilde{\chi}_1^0) = 300$ GeV, BR($b\tilde{\chi}_1^0$)=1 $m(\tilde{\chi}_1^0) = 300$ GeV, BR($b\tilde{\chi}_1^\pm$)=BR($t\tilde{\chi}_1^\pm$)=0.5 $m(\tilde{\chi}_1^0) = 200$ GeV, $m(\tilde{\chi}_1^\pm) = 300$ GeV, BR($t\tilde{\chi}_1^\pm$)=1	1708.09266, 1711.03301 1708.09266 ATLAS-CONF-2019-015
$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\bar{\chi}_2^0 \rightarrow b\tilde{h}\tilde{\chi}_1^0$		0 e, μ	6 b	E_T^{miss} 139	\tilde{b}_1 \tilde{b}_1	Forbidden 0.23-0.48	$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 100$ GeV $\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 0$ GeV	1908.03122 1908.03122	
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\bar{\chi}_1^0$ or $t\bar{\chi}_1^0$		0-2 e, μ	0-2 jets/1-2 b	E_T^{miss} 36.1	\tilde{t}_1	1.0	$m(\tilde{\chi}_1^0) = 1$ GeV	1506.08616, 1709.04183, 1711.11520	
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$		1 e, μ	3 jets/1 b	E_T^{miss} 139	\tilde{t}_1	0.44-0.59	$m(\tilde{\chi}_1^0) = 400$ GeV	ATLAS-CONF-2019-017	
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{\tau}b\nu, \tilde{\tau}_1 \rightarrow \tau\tilde{G}$		1 $\tau + 1 e, \mu, \tau$	2 jets/1 b	E_T^{miss} 36.1	\tilde{t}_1	1.16	$m(\tilde{\tau}_1) = 800$ GeV	1803.10178	
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\bar{\chi}_1^0/\tilde{c}\tilde{c}, \tilde{c} \rightarrow c\bar{\chi}_1^0$		0 e, μ	2 c	E_T^{miss} 36.1	\tilde{c} \tilde{t}_1 \tilde{t}_1	0.85 0.46 0.43	$m(\tilde{\chi}_1^0) = 0$ GeV $m(\tilde{t}_1, \tilde{c}) - m(\tilde{\chi}_1^0) = 50$ GeV $m(\tilde{t}_1, \tilde{c}) - m(\tilde{\chi}_1^0) = 5$ GeV	1805.01649 1805.01649 1711.03301	
$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + h$		0 e, μ	mono-jet	E_T^{miss} 36.1	\tilde{t}_2	0.32-0.88	$m(\tilde{\chi}_1^0) = 0$ GeV, $m(\tilde{t}_1) - m(\tilde{\chi}_1^0) = 180$ GeV	1706.03986	
$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$		1-2 e, μ 3 e, μ	4 b 1 b	E_T^{miss} 36.1 E_T^{miss} 139	\tilde{t}_2	0.86	$m(\tilde{\chi}_1^0) = 360$ GeV, $m(\tilde{t}_1) - m(\tilde{\chi}_1^0) = 40$ GeV	ATLAS-CONF-2019-016	
EW direct		$\tilde{\chi}_1^\pm\tilde{\chi}_2^0$ via WZ	2-3 e, μ $ee, \mu\mu$	≥ 1	E_T^{miss} 36.1 E_T^{miss} 139	$\tilde{\chi}_1^\pm/\tilde{\chi}_2^0$ $\tilde{\chi}_1^\pm/\tilde{\chi}_2^0$	0.6 0.205	$m(\tilde{\chi}_1^0) = 0$ $m(\tilde{\chi}_1^\pm) - m(\tilde{\chi}_1^0) = 5$ GeV	1403.5294, 1806.02293 ATLAS-CONF-2019-014
		$\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp$ via WW	2 e, μ		E_T^{miss} 139	$\tilde{\chi}_1^\pm$	0.42	$m(\tilde{\chi}_1^0) = 0$	1908.08215
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0$ via Wh	0-1 e, μ	2 $b/2 \gamma$	E_T^{miss} 139	$\tilde{\chi}_1^\pm/\tilde{\chi}_2^0$	0.74	$m(\tilde{\chi}_1^0) = 70$ GeV	ATLAS-CONF-2019-019, 1909.09226	
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp$ via $\tilde{\ell}_L\tilde{\nu}$	2 e, μ		E_T^{miss} 139	$\tilde{\chi}_1^\pm$	1.0	$m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^\pm) + m(\tilde{\chi}_1^0))$	ATLAS-CONF-2019-008	
	$\tilde{\tau}\tilde{\tau}, \tilde{\tau} \rightarrow \tau\tilde{\chi}_1^0$	2 τ		E_T^{miss} 139	$\tilde{\tau}$	0.16-0.3 0.12-0.39	$m(\tilde{\chi}_1^0) = 0$	ATLAS-CONF-2019-018	
	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\bar{\chi}_1^0$	2 e, μ	0 jets	E_T^{miss} 139	$\tilde{\ell}$	0.7	$m(\tilde{\chi}_1^0) = 0$	ATLAS-CONF-2019-008	
		2 e, μ	≥ 1	E_T^{miss} 139	$\tilde{\ell}$	0.256	$m(\tilde{\ell}) - m(\tilde{\chi}_1^0) = 10$ GeV	ATLAS-CONF-2019-014	
	$\tilde{H}\tilde{H}, \tilde{H} \rightarrow h\tilde{G}/Z\tilde{G}$	0 e, μ 4 e, μ	$\geq 3 b$ 0 jets	E_T^{miss} 36.1 E_T^{miss} 36.1	\tilde{H} \tilde{H}	0.13-0.23 0.3	BR($\tilde{H} \rightarrow h\tilde{G}$)=1 BR($\tilde{H} \rightarrow Z\tilde{G}$)=1	1806.04030 1804.03602	
Long-lived particles	Direct $\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet	E_T^{miss} 36.1	$\tilde{\chi}_1^\pm$ $\tilde{\chi}_1^\pm$	0.46 0.15	Pure Wino Pure Higgsino	1712.02118 ATL-PHYS-PUB-2017-019	
	Stable \tilde{g} R-hadron	Multiple		36.1	\tilde{g}	2.0		1902.01636, 1808.04095	
	Metastable \tilde{g} R-hadron, $\tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0$	Multiple		36.1	\tilde{g}	2.05 2.4	$m(\tilde{\chi}_1^0) = 100$ GeV	1710.04901, 1808.04095	
RPV	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e\mu/\tau\mu/\mu\tau$	$e\mu, e\tau, \mu\tau$		3.2	$\tilde{\nu}_\tau$	1.9	$\lambda'_{511} = 0.11, \lambda'_{132/133/233} = 0.07$	1607.08079	
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp/\tilde{\chi}_2^0 \rightarrow WW/Z\ell\ell\nu\nu$	4 e, μ	0 jets	E_T^{miss} 36.1	$\tilde{\chi}_1^\pm/\tilde{\chi}_2^0$	0.82 1.33	$m(\tilde{\chi}_1^0) = 100$ GeV	1804.03602	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow q\bar{q}q$	4-5 large-R jets	Multiple	36.1 36.1	\tilde{g} \tilde{g}	1.3 1.05 2.0	Large λ'_{12} $\lambda'_{112} = 2e-4, 2e-5$	1804.03568 ATLAS-CONF-2018-003	
	$\tilde{u}\tilde{u}, \tilde{t} \rightarrow t\bar{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow t\bar{b}s$	Multiple		36.1	\tilde{g}	0.55 1.05	$m(\tilde{\chi}_1^0) = 200$ GeV, bino-like	ATLAS-CONF-2018-003	
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\bar{s}$	2 jets + 2 b		36.7	\tilde{t}_1	0.61	$m(\tilde{\chi}_1^0) = 200$ GeV, bino-like	1710.07171	
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow q\bar{\ell}$	2 e, μ 1 μ	2 b DV	36.1 136	\tilde{t}_1 \tilde{t}_1	0.4-1.45 1.0 1.6	BR($\tilde{t}_1 \rightarrow b\bar{\ell}$) > 20% BR($\tilde{t}_1 \rightarrow q\bar{\mu}$) = 100%, $\cos\theta = 1$	1710.05544 ATLAS-CONF-2019-006	

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

10⁻¹

1

Mass scale [TeV]

"Exotics" limits

ATLAS Exotics Searches* - 95% CL Upper Exclusion Limits

Status: May 2019

$\int L dt = (3.2 - 139) \text{ fb}^{-1}$

ATLAS Preliminary

$\sqrt{s} = 8, 13 \text{ TeV}$

Model	ℓ, γ	Jets [†]	Emiss [†]	$\int L dt [\text{fb}^{-1}]$	Limit	Reference		
Extra dimensions	ADD $G_{KK} \rightarrow g/g$	$0 e, \mu$	1-4	Yes	36.1	M_{Pl} 7.7 TeV	1711.03301	
	ADD non-resonant $\gamma\gamma$	-	-	-	36.1	M_{Pl} 2.9	1707.04147	
	ADD QBH	-	2	-	37.0	M_{Pl} 8.8 TeV	1703.09127	
	ADD BH High $\rightarrow \mu\mu$	$\geq 1 e, \mu$	≥ 2	-	3.2	M_{Pl} 8.2 TeV	1608.02265	
	ADD BH multiple	-	≥ 3	-	3.6	M_{Pl} 9.95 TeV	1512.02586	
	RS1 $G_{KK} \rightarrow \gamma\gamma$	2γ	-	-	36.7	G_{KK} mass 4.1 TeV	1707.04147	
	Bulk RS $G_{KK} \rightarrow WW/ZZ$	multi-channel	-	-	36.1	G_{KK} mass 2.3 TeV	1608.02265	
	Bulk RS $G_{KK} \rightarrow WW \rightarrow e\mu\mu$	$0 e, \mu$	2,3	-	139	G_{KK} mass 1.6 TeV	ATLAS-CONF-2019-003	
	Bulk RS $G_{KK} \rightarrow t\bar{t}$	$1 e, \mu$	$\geq 1 b, \geq 1 \bar{b}$	Yes	36.1	G_{KK} mass 3.8 TeV	1804.10823	
	ZUED / HPP	$2 e, \mu$	$\geq 2 b, \geq 3 \bar{b}$	Yes	36.1	HK mass 1.8 TeV	1803.09678	
Gauge bosons	SSM $Z' \rightarrow \ell\ell$	2ℓ	-	-	139	Z' mass 2.42 TeV	1903.06248	
	SSM $Z' \rightarrow \tau\tau$	-	-	-	36.1	Z' mass 2.1 TeV	1709.07242	
	Leptophobic $Z' \rightarrow b\bar{b}$	$1 e, \mu$	$\geq 1 b, \geq 1 \bar{b}$	Yes	36.1	Z' mass 3.0 TeV	1805.06299	
	Leptophobic $Z' \rightarrow \tau\tau$	$1 e, \mu$	$\geq 1 b, \geq 1 \bar{b}$	Yes	36.1	Z' mass 3.0 TeV	1804.10323	
	SSM $W' \rightarrow \ell\nu$	$1 e, \mu$	-	-	139	W' mass 6.0 TeV	CERN-EP-2019-100	
	SSM $W' \rightarrow \tau\nu$	1τ	-	-	Yes	36.1	W' mass 3.7 TeV	1801.06992
	HVT $V' \rightarrow WZ \rightarrow e\mu\mu$ model B	$0 e, \mu$	2,3	-	139	V' mass 3.9 TeV	ATLAS-CONF-2019-003	
	HVT $V' \rightarrow WW/ZH$ model B	multi-channel	-	-	36.1	V' mass 3.25 TeV	1712.06518	
	LRSM $W_{\mu} \rightarrow t\bar{b}$	multi-channel	-	-	36.1	W_{μ} mass 3.25 TeV	1807.10473	
	LRSM $W_{\mu} \rightarrow \mu Nb$	2μ	1,3	-	80	W_{μ} mass 5.0 TeV	1904.18079	
CI	CI $e\mu\mu$	-	2	-	37.0	A 21.8 TeV κ_{11}	1703.04127	
	CI $\ell\ell\mu\mu$	$\geq 1 e, \mu$	$\geq 1 b, \geq 1 \bar{b}$	Yes	36.1	A 40.0 TeV κ_{11}	1707.02484	
	CI $\tau\tau\tau\tau$	$\geq 1 e, \mu$	$\geq 1 b, \geq 1 \bar{b}$	Yes	36.1	A 2.57 TeV	1811.02305	
DM	Axial-vector mediator (Dirac DM)	$0 e, \mu$	1-4	Yes	36.1	m_{DM} 1.55 TeV	1711.03301	
	Colored scalar mediator (Dirac DM)	$0 e, \mu$	1-4	Yes	36.1	m_{DM} 1.67 TeV	1711.03301	
	VV_{μ} EFT (Dirac DM)	$0 e, \mu$	1,3,4	Yes	3.2	M_{Pl} 700 GeV	1608.02372	
Scalar reson. $\phi \rightarrow t\bar{t}$ (Dirac DM)	$0.1 e, \mu$	1 b, 0-1 \bar{b}	Yes	36.1	m_{ϕ} 3.4 TeV	1812.00743		
LQ	Scalar LQ 1 st gen	$1.2 e$	≥ 2	Yes	36.1	LQ mass 1.4 TeV	1902.00377	
	Scalar LQ 2 nd gen	1.2μ	≥ 2	Yes	36.1	LQ mass 1.56 TeV	1902.00377	
	Scalar LQ 3 rd gen	2τ	2 b	-	36.1	LQ mass 1.03 TeV	1902.08103	
	Scalar LQ 3 rd gen	$0.1 e, \mu$	2 b	Yes	36.1	LQ mass 870 GeV	1902.08103	
Heavy quarks	VLO $T\bar{T} \rightarrow Ht/Zt/Wb+X$	multi-channel	-	-	36.1	T mass 1.37 TeV	1808.02343	
	VLO $SB \rightarrow Wt/Zb+X$	multi-channel	-	-	36.1	S mass 1.34 TeV	1808.02343	
	VLO $T_{\text{SU}(2)} T_{\text{SU}(2)} \rightarrow Wt+X$	$2(SS) \geq 2 e, \mu \geq 1 b, \geq 1 \bar{b}$	Yes	36.1	$T_{\text{SU}(2)}$ mass 1.64 TeV	1807.11883		
	VLO $V \rightarrow Wb+X$	$1 e, \mu$	$\geq 1 b, \geq 1 \bar{b}$	Yes	36.1	V mass 1.85 TeV	1812.07343	
	VLO $Q \rightarrow Hb+X$	$0 e, \mu, \tau$	$\geq 1 b, \geq 1 \bar{b}$	Yes	79.8	Q mass 1.21 TeV	ATLAS-CONF-2019-024	
	VLO $QQ \rightarrow WbWq$	$1 e, \mu$	≥ 4	Yes	20.3	Q mass 690 GeV	1509.04261	
Excited quarks	Excited quark $q^* \rightarrow q\gamma$	-	2	-	139	q^* mass 6.7 TeV	ATLAS-CONF-2019-007	
	Excited quark $q^* \rightarrow q\gamma$	1γ	1	-	36.7	q^* mass 5.3 TeV	1709.10440	
	Excited quark $q^* \rightarrow b\gamma$	-	1 b, 1 \bar{b}	-	36.1	q^* mass 2.5 TeV	1805.06299	
	Excited lepton $\ell^* \rightarrow \ell\gamma$	$3 e, \mu, \tau$	-	-	20.3	ℓ^* mass 3.0 TeV	1411.2921	
Other	Type III Seesaw	$1 e, \mu$	≥ 2	Yes	79.8	N^c mass 560 GeV	ATLAS-CONF-2019-004	
	LRSM Majorana ν	2μ	2	-	36.1	N_{μ} mass 870 GeV	1809.1110	
	Higgs triplet $H^{\pm\pm} \rightarrow \ell\ell$	$2.3, 4 e, \mu$ (SS)	-	-	36.1	$H^{\pm\pm}$ mass 400 GeV	1710.07474	
	Higgs triplet $H^{\pm\pm} \rightarrow \ell\tau$	$3 e, \mu, \tau$	-	-	20.3	$H^{\pm\pm}$ mass 400 GeV	1411.2921	
Multi-charged particles	-	-	-	36.1	multi-charged particle mass 1.22 TeV	1812.02467		
Magnetic monopoles	-	-	-	34.4	monopole mass 2.37 TeV	1905.1013		

*Only a selection of the available mass limits on new states or phenomena is shown.

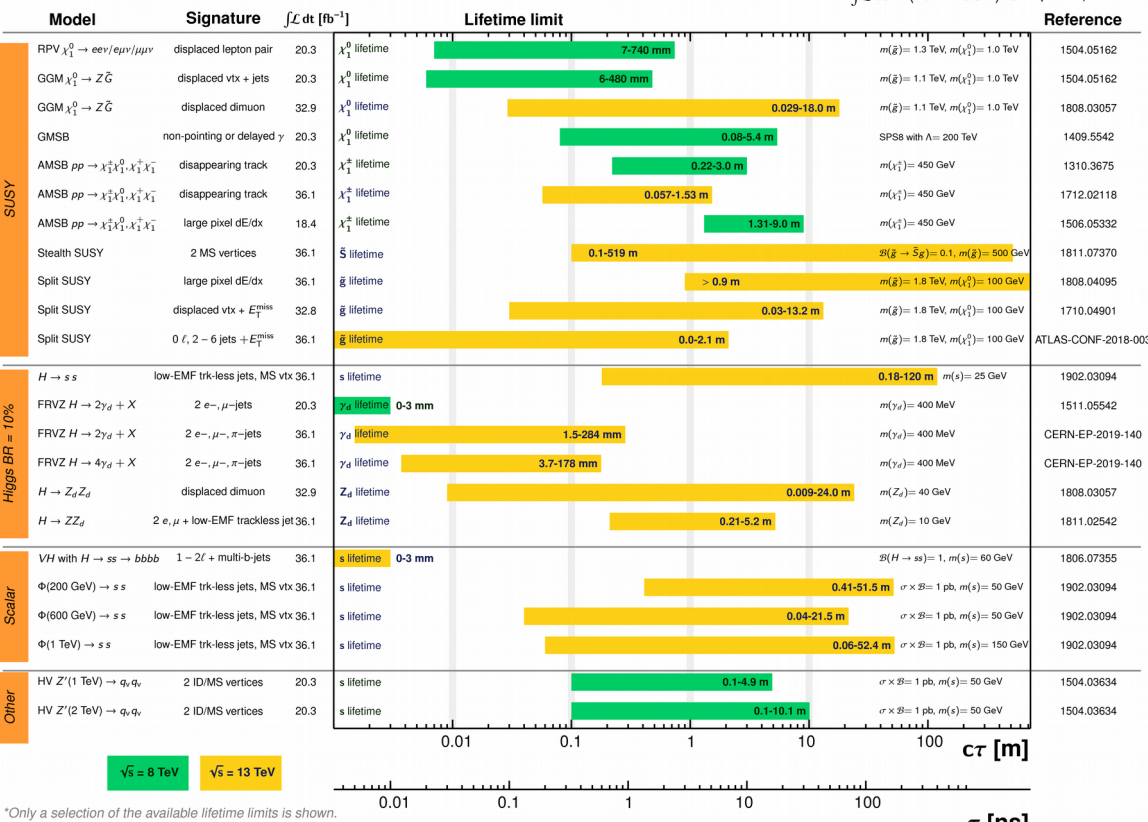
†Small-radius (large-radius) jets are denoted by the letter j (\bar{j}).

ATLAS Long-lived Particle Searches* - 95% CL Exclusion

Status: July 2019

ATLAS Preliminary

$\int L dt = (18.4 - 36.1) \text{ fb}^{-1}$ $\sqrt{s} = 8, 13 \text{ TeV}$

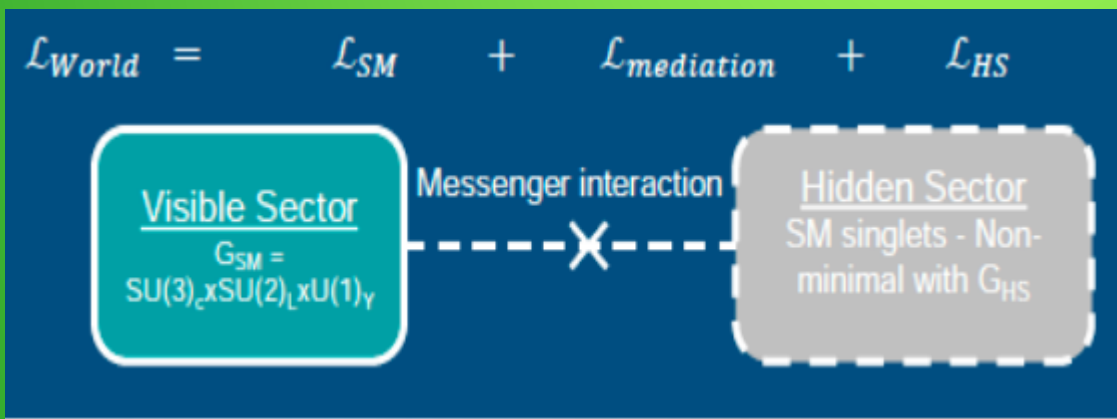


*Only a selection of the available lifetime limits is shown.

- Keep in mind: limits on particle lifetimes limited by size of LHC detectors

The “hidden sector” approach to new physics

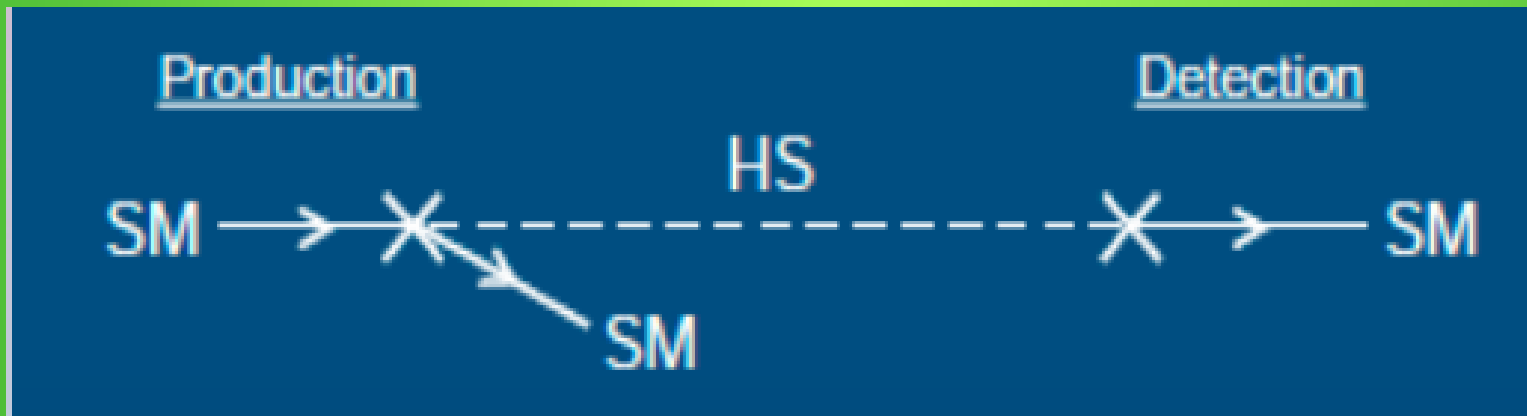
- Maybe new particles have not been yet found not because they are heavy, but because their coupling is very small, or null
- If an additional term to the Lagrangian is not interacting with SM, there could be invisible particles contributing to dark matter, and no naturalness issues
- However, an interference term between the Lagrangians would allow a very small coupling:



$$\mathcal{L}_{mediation} = \sum_{k,l,n} \frac{\mathcal{O}_{HS}^{(k)} \mathcal{O}_{SM}^{(l)}}{\Lambda^n}$$

“Portals”

- Indications for a Hidden Sector may come from “ordinary” particles (SM, SUSY, axions etc.) acting as mediators with the HS Lagrangian
- The experimental signature is either missing energy or the appearance of SM particles very far away from its production, indicating an “oscillation” into the HS (and back)



Models	Final states
Neutrino portal, SUSY neutralino	$\ell^\pm \pi^\mp, \ell^\pm K^\mp, \ell^\pm \rho^\mp, \rho^\pm \rightarrow \pi^\pm \pi^0$
Vector, scalar, axion portals, SUSY sgoldstino	$\ell^+ \ell^-$
Vector, scalar, axion portals, SUSY sgoldstino	$\pi^+ \pi^-, K^+ K^-$
Neutrino portal, SUSY neutralino, axino	$\ell^+ \ell^- \nu$
Axion portal, SUSY sgoldstino	$\gamma \gamma$
SUSY sgoldstino	$\pi^0 \pi^0$

Standard Model portals:

D = 2: Vector portal

- Kinetic mixing with massive dark/secluded/paraphoton V : $\frac{1}{2} \epsilon F_{\mu\nu}^{SM} F_{HS}^{\mu\nu}$

→ Interaction with 'mirror world' constituting dark matter

D = 2: Higgs portal

- Mass mixing with dark singlet scalar χ : $(\mu\chi + \lambda\chi^2)H^\dagger H$

$$\begin{pmatrix} H \\ h \end{pmatrix} = \begin{pmatrix} \cos \rho & -\sin \rho \\ \sin \rho & \cos \rho \end{pmatrix} \begin{pmatrix} \phi'_0 \\ S' \end{pmatrix}$$

→ Mass to Higgs boson and right-handed neutrino, and function as inflaton in accordance with Planck and BICEP measurements

D = 5/2: Neutrino portal

- Mixing with right-handed neutrino N (Heavy Neutral Lepton): $YH^\dagger \bar{N}L$

→ Neutrino oscillation, baryon asymmetry, dark matter

D = 4: Axion portal

- Mixing with Axion Like Particles, pseudo-scalars pNGB, axial vectors: $\frac{a}{F} G_{\mu\nu} \tilde{G}^{\mu\nu}, \frac{\partial_\mu a}{F} \bar{\psi} \gamma_\mu \gamma_5 \psi$, etc

→ Solve strong CP problem, Inflaton

- And possibly higher dimensional operator portals and **Super-Symmetric portals** (light neutralino, light sgoldstino,...)

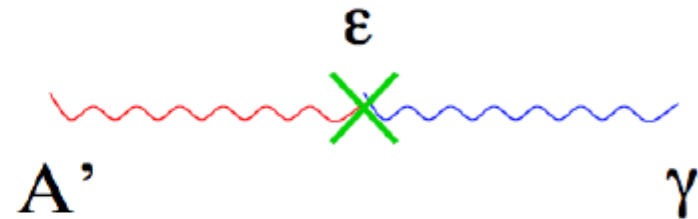
→ SUSY parameter space explored by LHC

→ Some of SUSY low-energy parameter space open to complementary searches

Vector and scalar portals

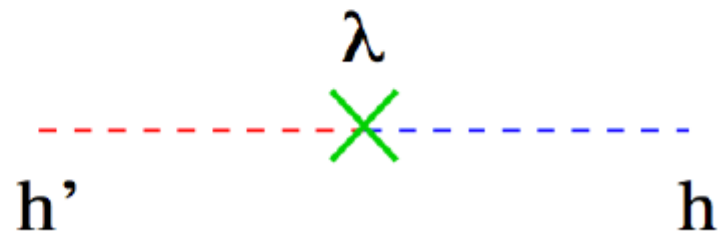
- Vector Portal:
(A' = “hidden photon”)

$$\epsilon F'_{\mu\nu} F^{\mu\nu}$$



- Higgs Portal:
(H' = “hidden Higgs”)

$$\lambda |H'|^2 |H|^2$$



Sterile neutrinos

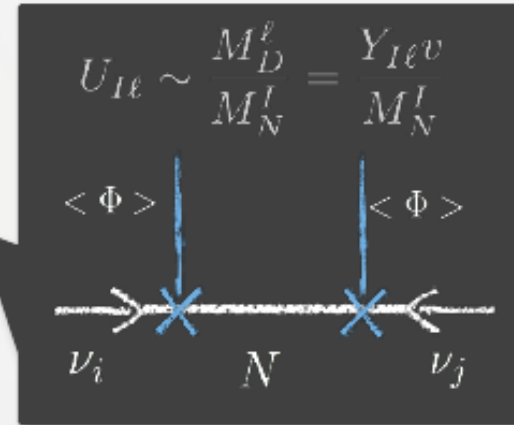
Fermions get mass via the Yukawa couplings:

$$-\mathcal{L}_{\text{Yukawa}} = Y_{ij}^d \overline{Q}_{Li} \phi D_{Rj} + Y_{ij}^u \overline{Q}_{Li} \tilde{\phi} U_{Rj} + Y_{ij}^\ell \overline{L}_{Li} \phi E_{Rj} + \text{h.c.}$$

If we want the same coupling for neutrinos, we need right-handed (sterile) neutrinos... the most generic Lagrangian is

$$\mathcal{L}_N = i \overline{N}_i \partial_\mu \gamma^\mu N_i - \frac{1}{2} M_{ij} \overline{N}_i^c N_j - Y_{ij}^\nu \overline{L}_{Li} \tilde{\phi} N_j$$

Kinetic term
Majorana mass term
Yukawa coupling



Seesaw mechanism:

$$\mathcal{V} = (\nu_{Li}, N_j)$$

$$-\mathcal{L}_{M_\nu} = \frac{1}{2} \overline{\mathcal{V}} M_\nu \mathcal{V} + \text{h.c.}$$

if $M_N \gg M_D$:

$$M_\nu = \begin{pmatrix} 0 & M_D \\ M_D^T & M_N \end{pmatrix}$$

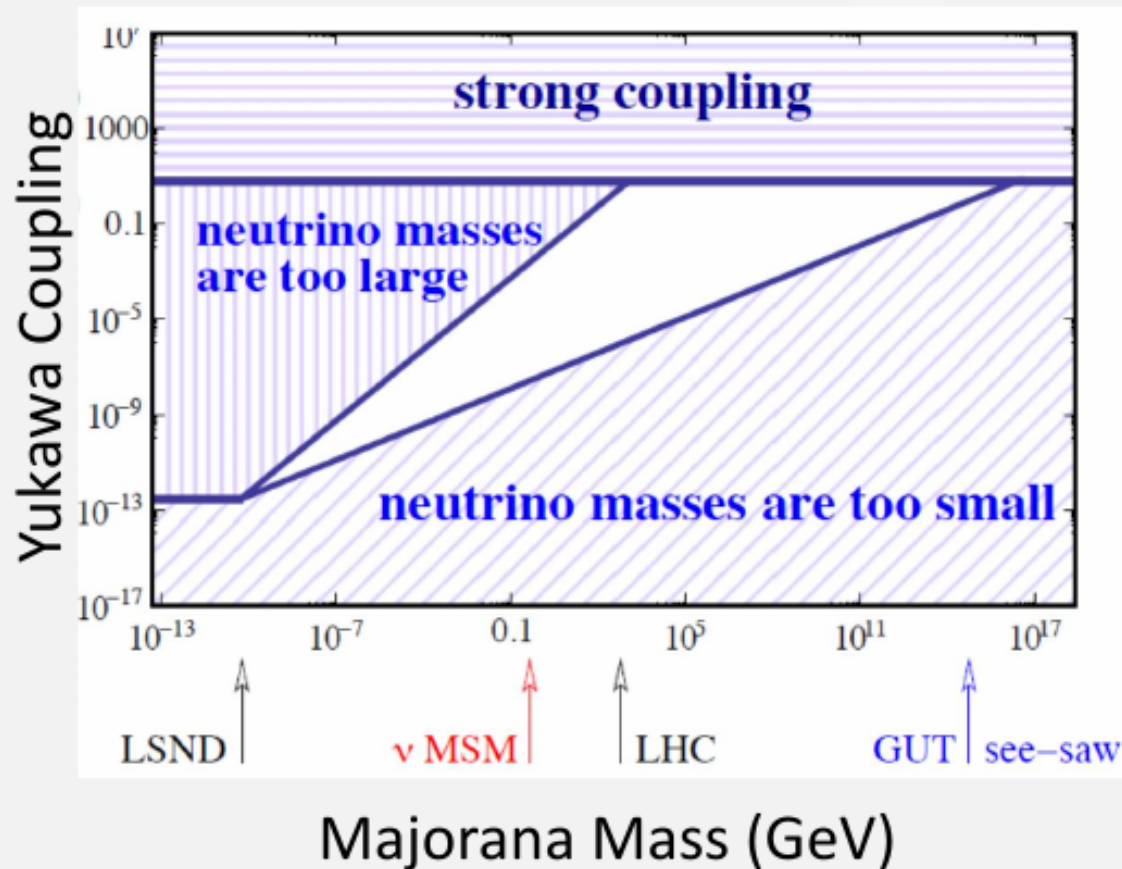
$$\lambda_\pm = \frac{M_N \pm \sqrt{M_N^2 + 4M_D^2}}{2}$$

$$\lambda_- \sim \frac{M_D^2}{M_N}$$

$$\lambda_+ \sim M_N$$

The see-saw mechanism

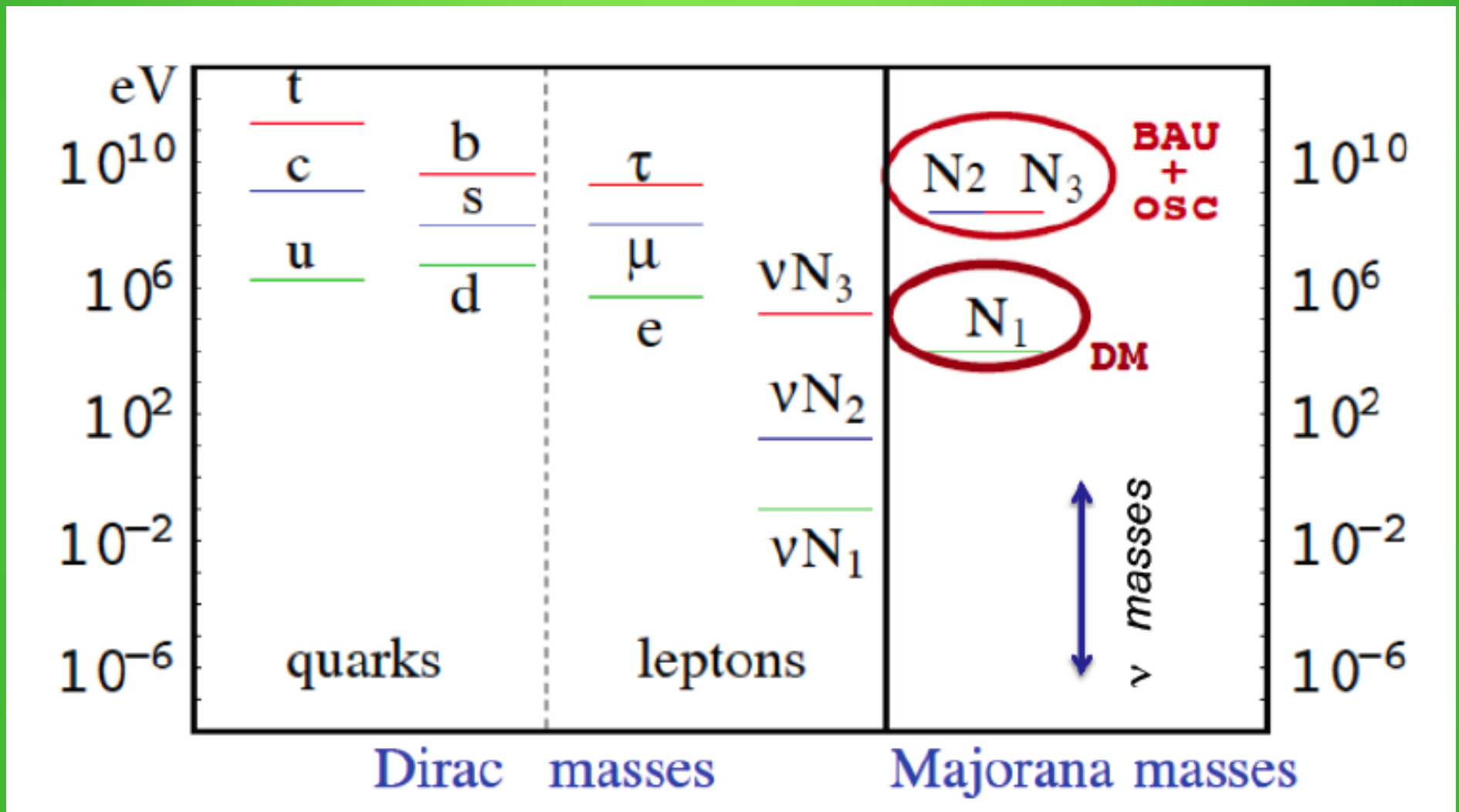
Seesaw formula $m_D \sim Y_{I\alpha} \langle \phi \rangle$ and $m_\nu = \frac{m_D^2}{M}$



- Assuming $m_\nu = 0.1\text{eV}$
 - if $Y \sim 1$ implies $M \sim 10^{14}\text{GeV}$
 - if $M_N \sim 1\text{GeV}$ implies $Y_\nu \sim 10^{-7}$
- remember $Y_{top} \sim 1$. and $Y_e \sim 10^{-6}$

If we want to explain the smallness of neutrino masses (in a natural way) the mass of sterile neutrinos should be at least at the GeV scale

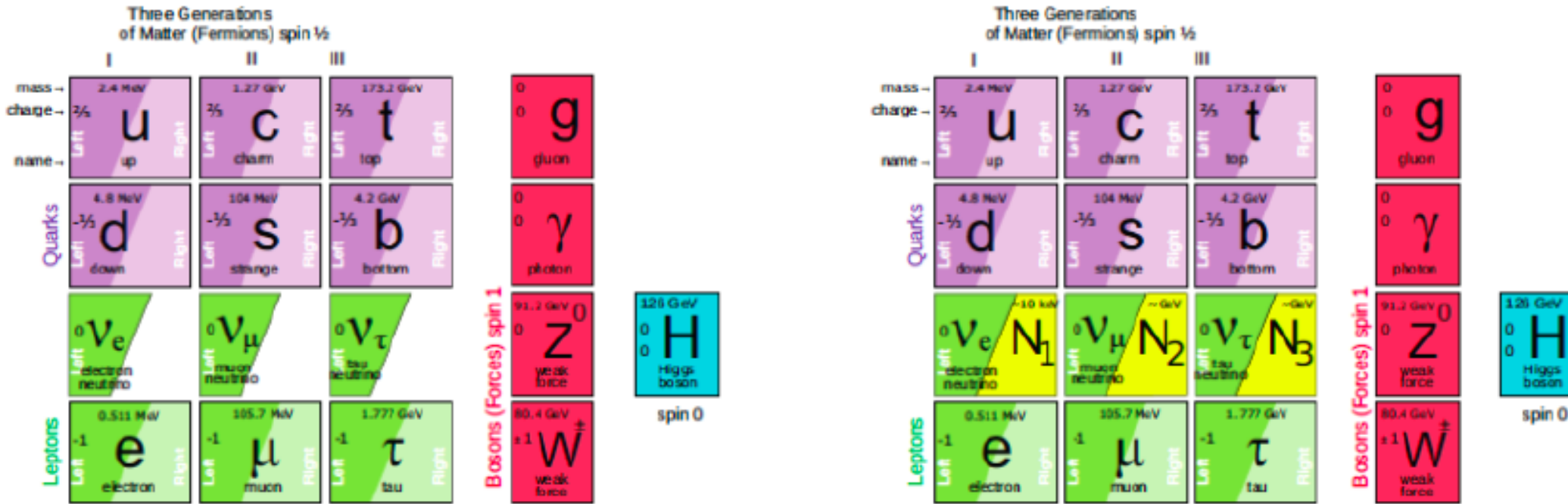
Resulting mass ranges



- Sterile neutrinos could have masses and couplings similar to those of the ordinary charged leptons

The ν MSSM

T.Asaka, M.Shaposhnikov, PL B620 (2005) 17
M.Shaposhnikov Nucl. Phys. B763 (2007) 49



Particle content of SM made symmetric by adding 3 HNL: N_1, N_2, N_3

With $M(N) \sim$ few KeV, it is a good DM candidate (or DM can be generated outside of this model through decay of inflaton)

With $M(N, N) \sim$ GeV, could explain Barion Asymmetry of Universe (via leptogenesis), and generate neutrino masses through see-saw.

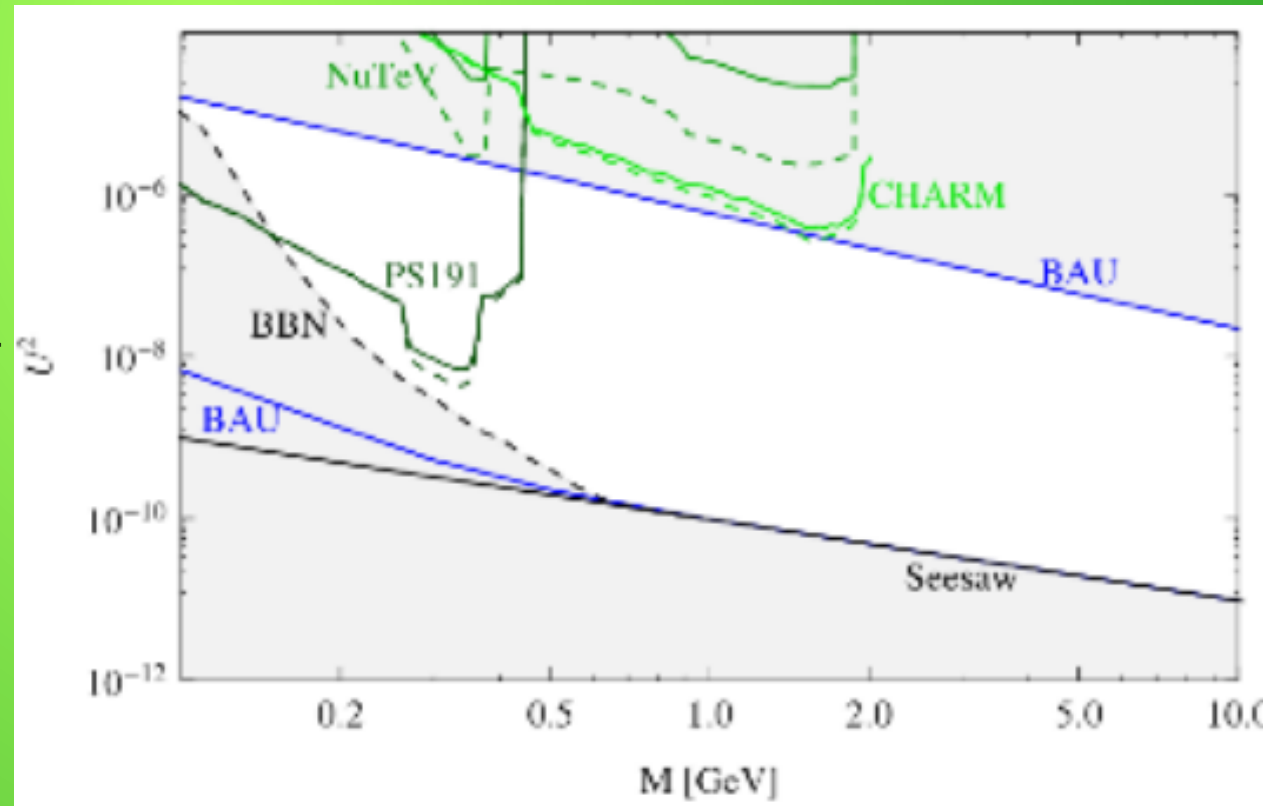
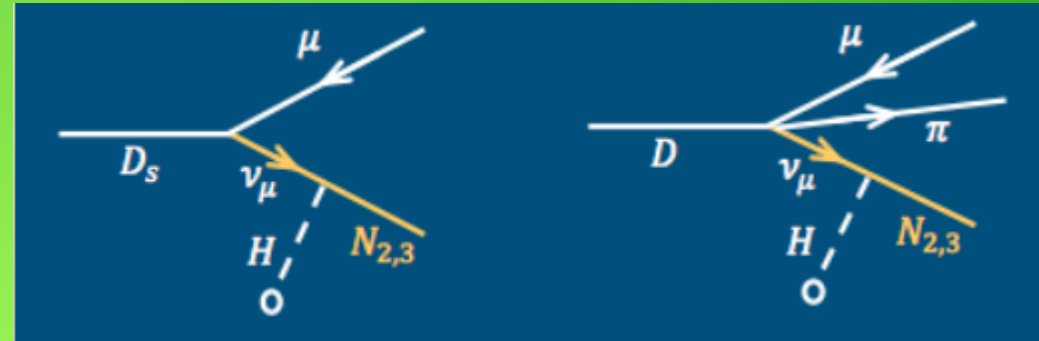
HNL production mechanism

Interaction with Higgs vev leads to a mixing with active neutrinos

Several past searches; PS191 used neutrinos from K decays, while other experiments not sensitive to mixings of cosmological interest.

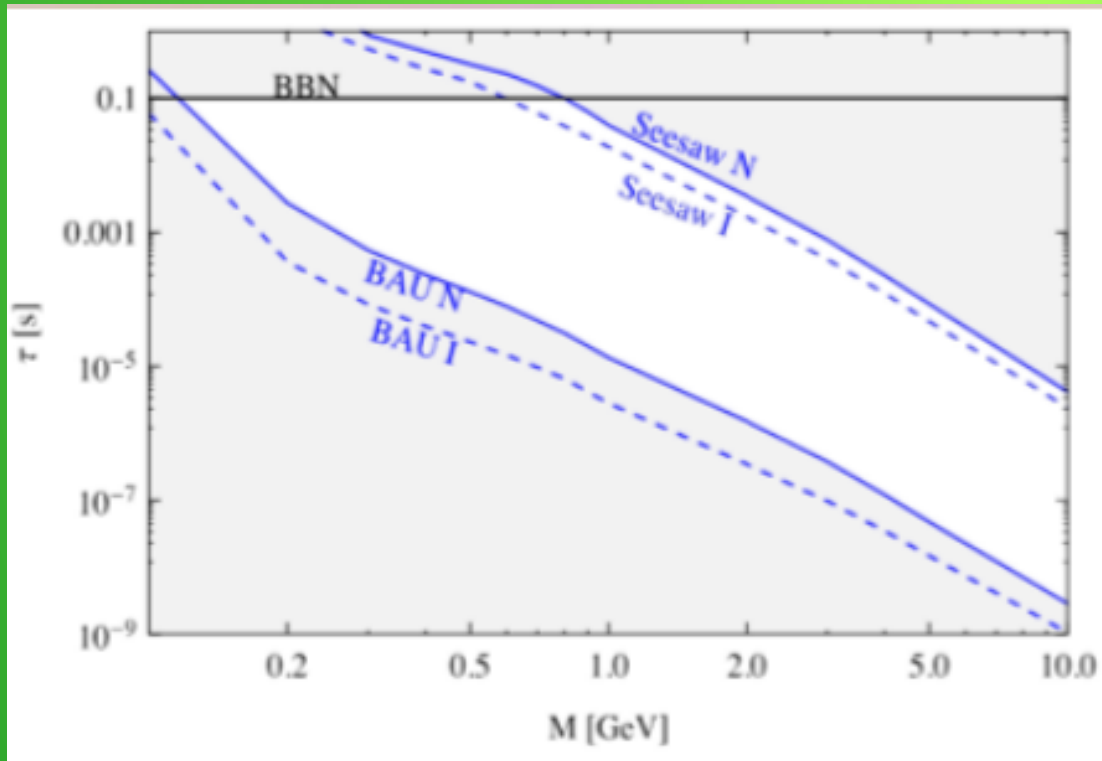
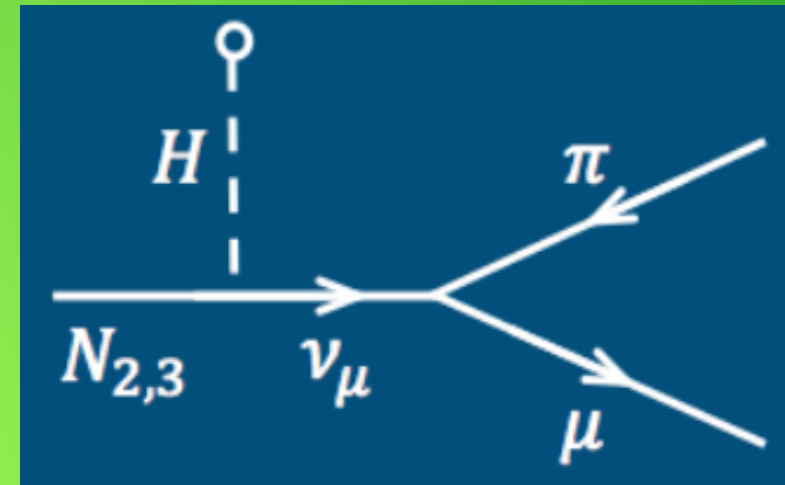
Latest result: LHCb with B decays obtained $U_{21} \approx 10^{-4}$, arXiv:1401.5361

Further exploration needed of the region with higher masses and smaller mixings



HNL decay modes

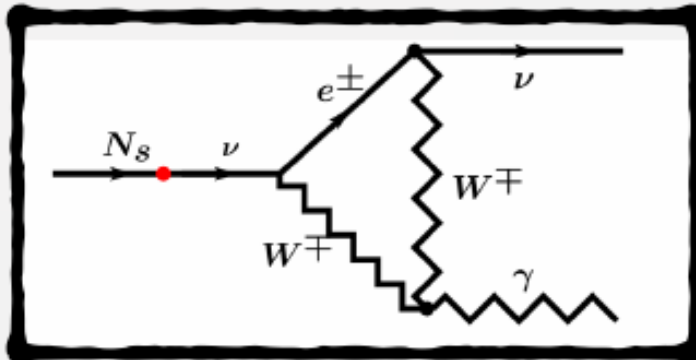
Interaction with Higgs vev would make it oscillate back into a virtual neutrino, that produces a muon and a W (\rightarrow hadrons, eg pions)
 Exact branching fractions depend n flavor mixing
 Due to small couplings, ms lifetimes, decay paths
 O(km)



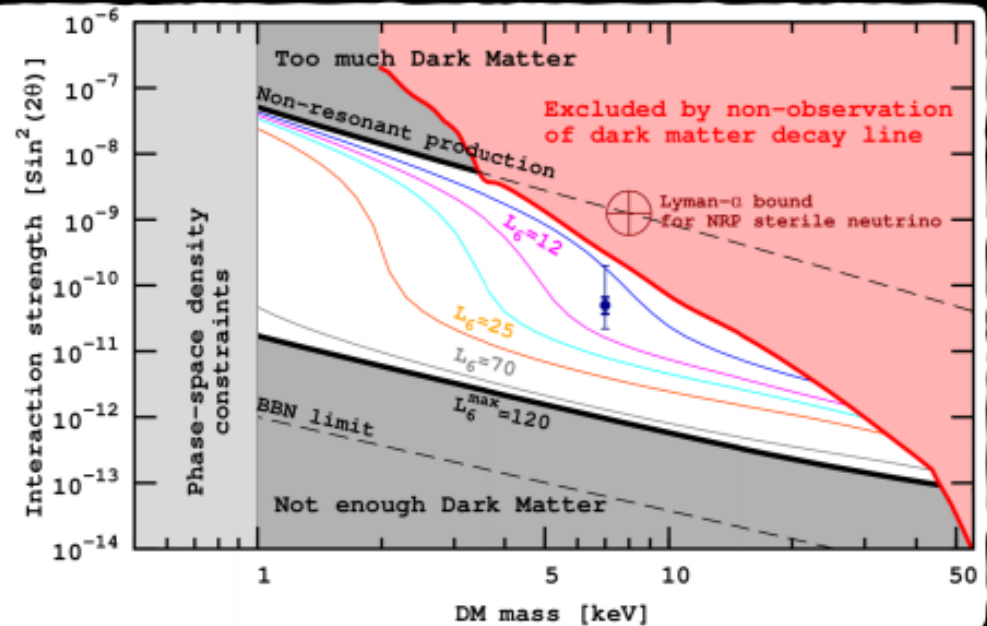
Decay mode	Branching ratio
$N_{2,3} \rightarrow \mu/e + \pi$	0.1 - 50 %
$N_{2,3} \rightarrow \mu^-/e^- + \rho^+$	0.5 - 20 %
$N_{2,3} \rightarrow \nu + \mu + e$	1 - 10 %

Constraints on N_1 mass

DM sterile neutrinos decay subdominantly as $N_1 \rightarrow \nu\gamma$ with a branching ratio $\mathcal{B}(N_1 \rightarrow \gamma\nu) \sim \frac{1}{123}$



Discussion in the community, not yet clear if this is a “good” signal, needs confirmation



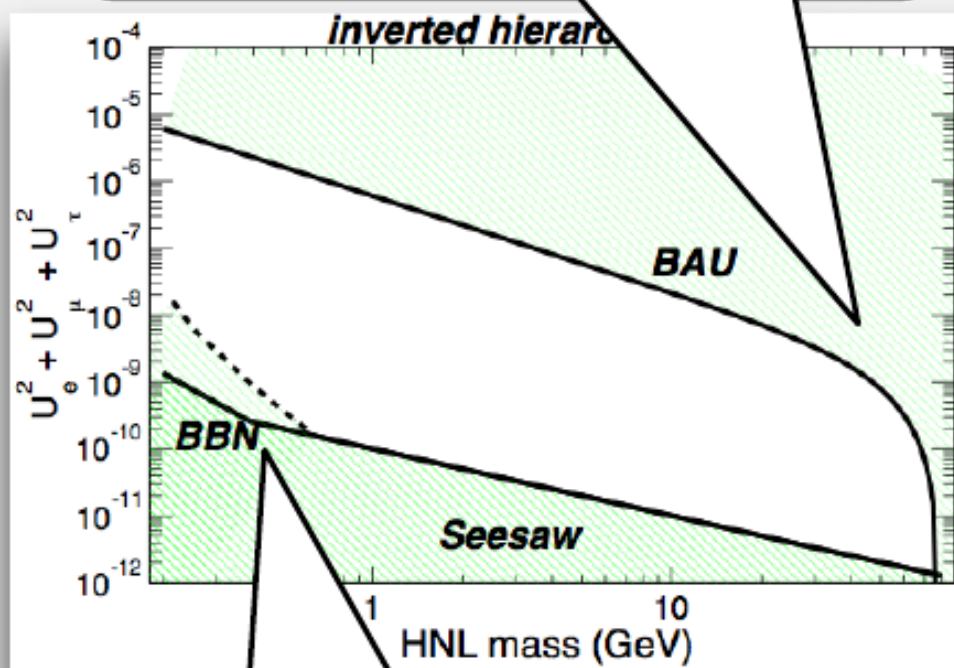
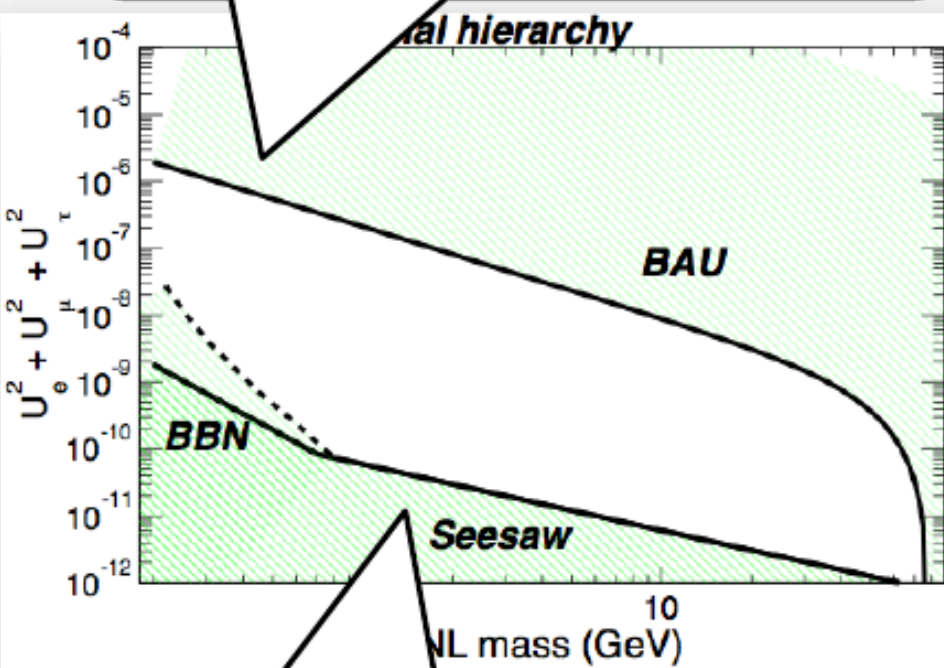
Bulbul et al. 2014 (arXiv:1402.2301)

Boyarsky et al. 2014 (arXiv:1402.4119)

Constraints on N_2, N_3 masses

If U^2 is too large, $N_{2,3}$ are in **thermal equilibrium** during the expansion of the Universe

At $M_N \geq M_W$ the rate is **enhanced** by $N \rightarrow Wl$ leading to stronger constraints on U^2

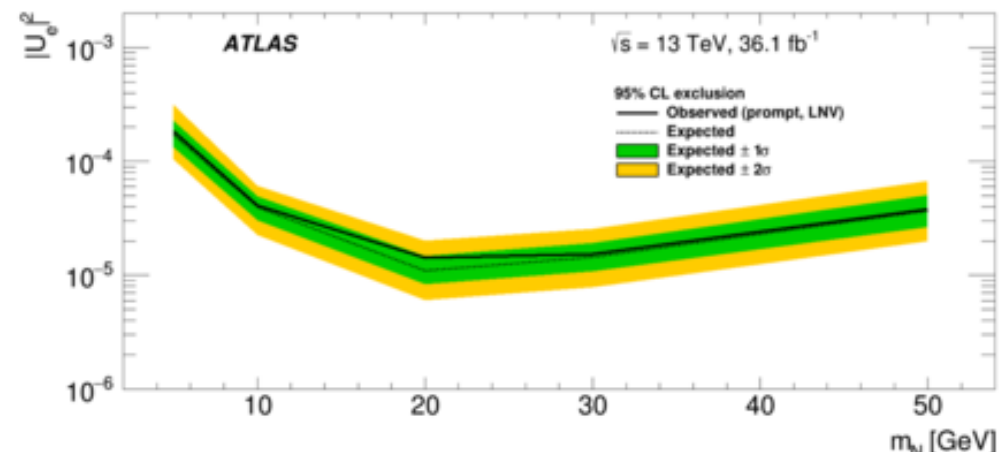
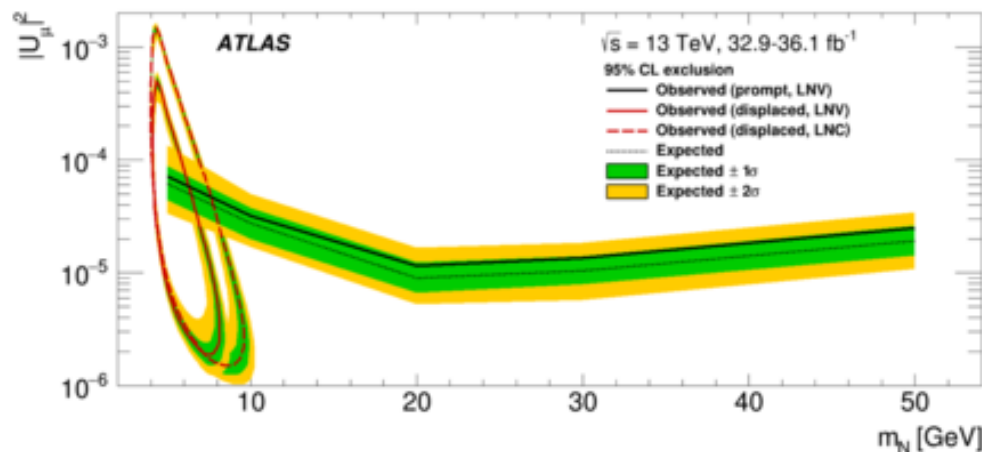
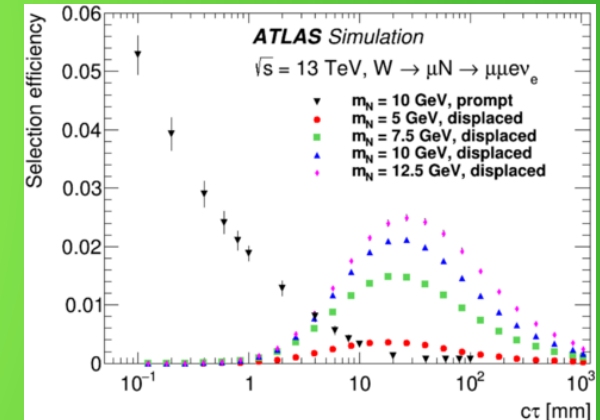
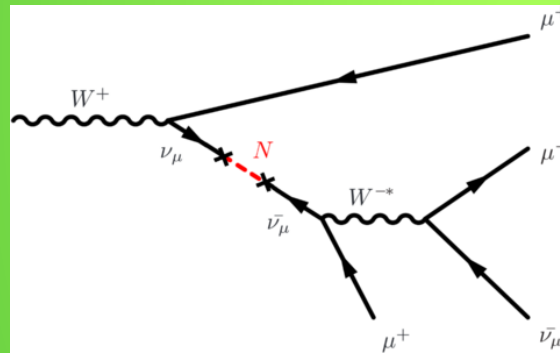


The **seesaw** limit defines the region where $N_{2,3}$ can explain the observed active neutrino Δm^2

If $\tau(N_2, N_3) < 0.1$ s, they cannot affect the **Big Bang nucleosynthesis**

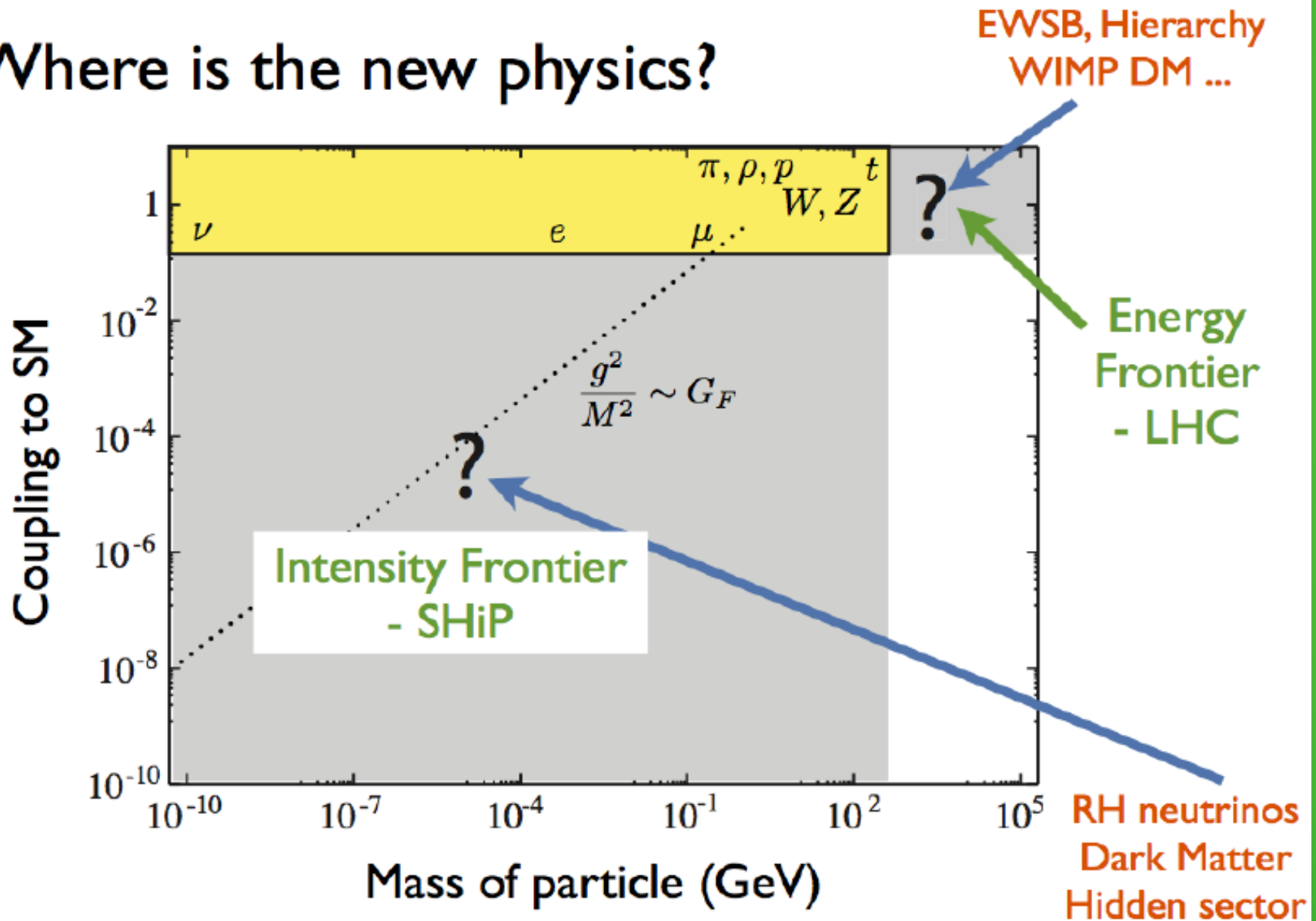
High-mass searches at the LHC

- Explore HNL mass range above 10 GeV
- Search for two same-sign leptons and no MET
- ATLAS paper JHEP 10(2019) 265 uses both prompt and displaced signatures



Searches in the cosmologically-interesting region

Where is the new physics?



Model-independent experimental considerations

We have to look for very weakly interacting particles:

- Production BR $O(1E-10)$
- Lifetimes $O(\text{km})$
- Can travel through ordinary matter

Cosmologically interesting masses $O(\text{GeV})$

- Produced through decays of mesons
- Can decay to mesons or charged leptons
- Full final-state reconstruction and particle ID

To have high intensities:

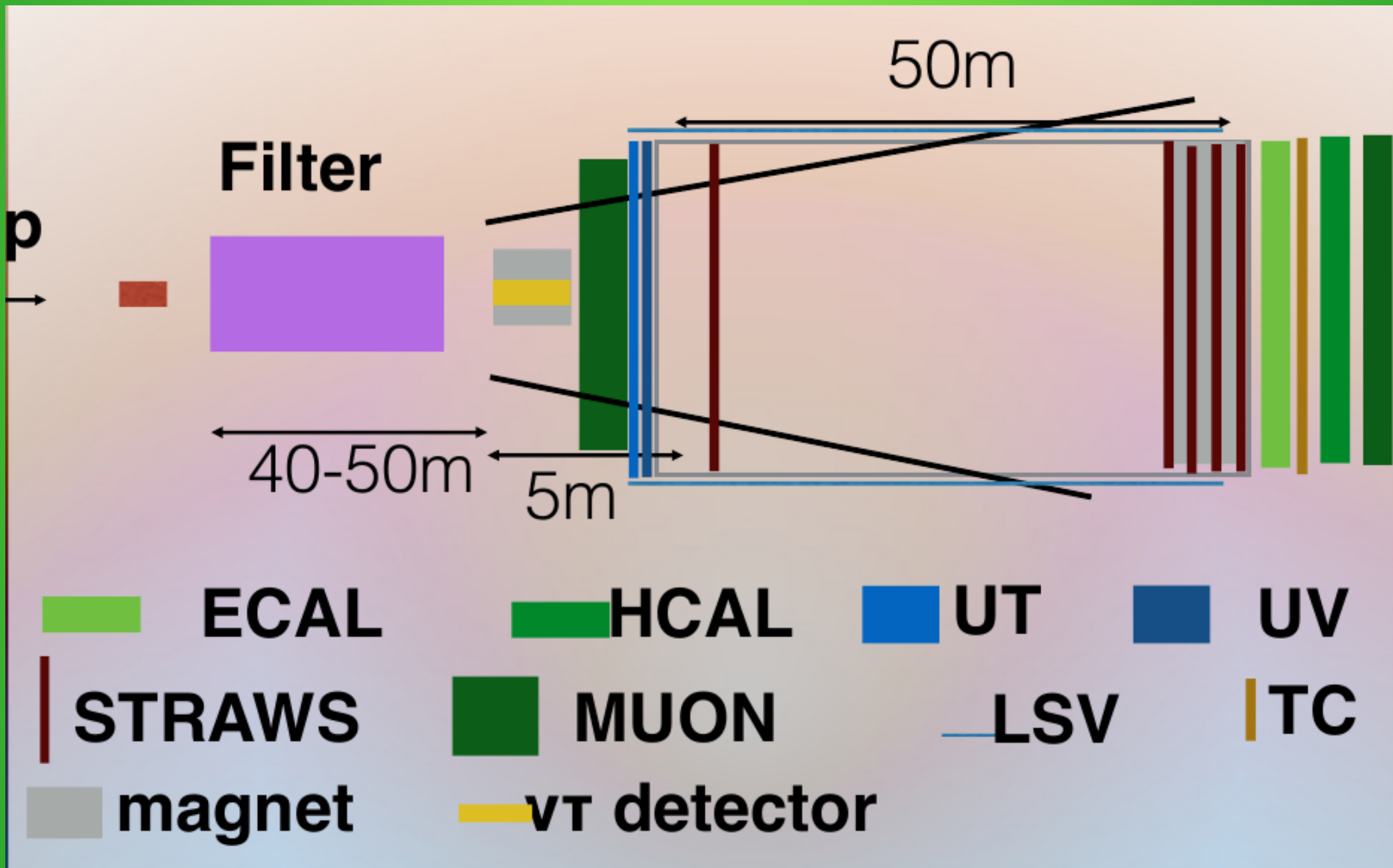
- fixed-target against a beam dump
- followed by a long decay tunnel and a spectrometer at the end

An experiment in practice

Use protons from CERN's SPS: 500 kW is 4×10^{13} protons/7 s $\rightarrow 2 \times 10^{20}$ in 5y

- Slow (ms \rightarrow 1s) and uniform extraction to reduce detector occupancy and combinatorics
- HS particles produced by mesons (mainly charm) decays; need to absorb all SM decay products to minimise BG \rightarrow heavy material thick target, with wide beam to dilute energy deposition (different from neutrino facility)
- Muons cannot be absorbed by target: muon shield, possibly magnetised
- Long decay tunnel away from external walls to minimise rescattering of muons and neutrons close to detector
- Vacuum in decay tunnel to reduce neutrino interactions
- Far-away detector with good PID and resolutions

Schematically...

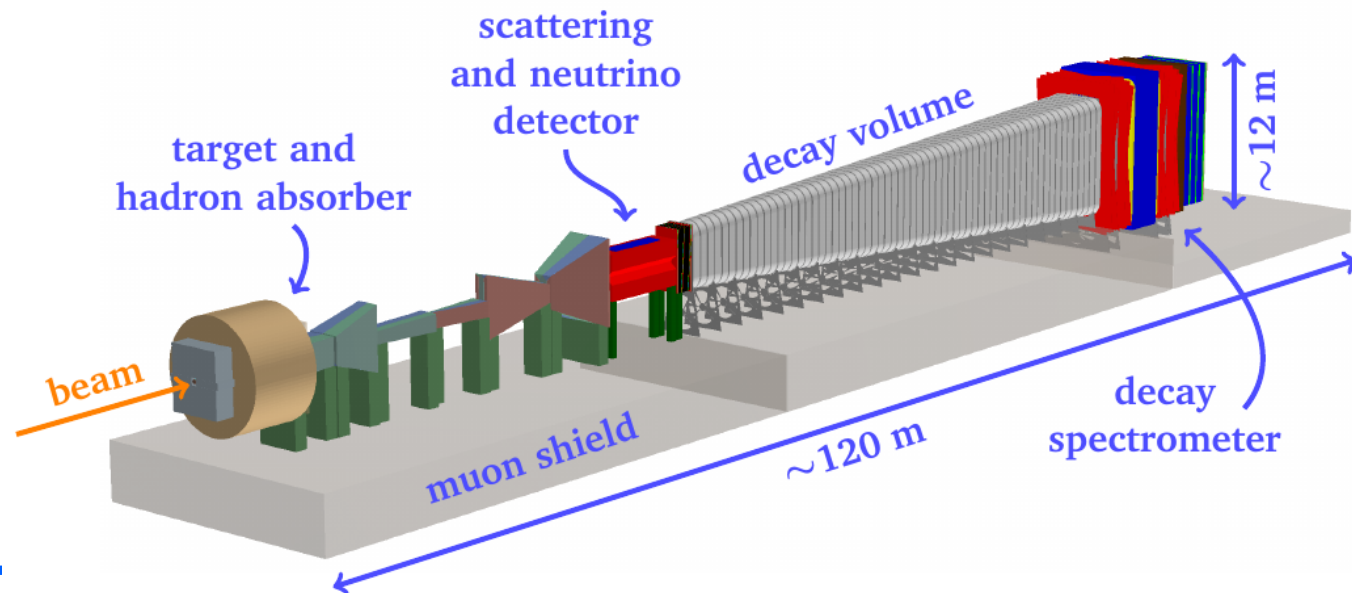
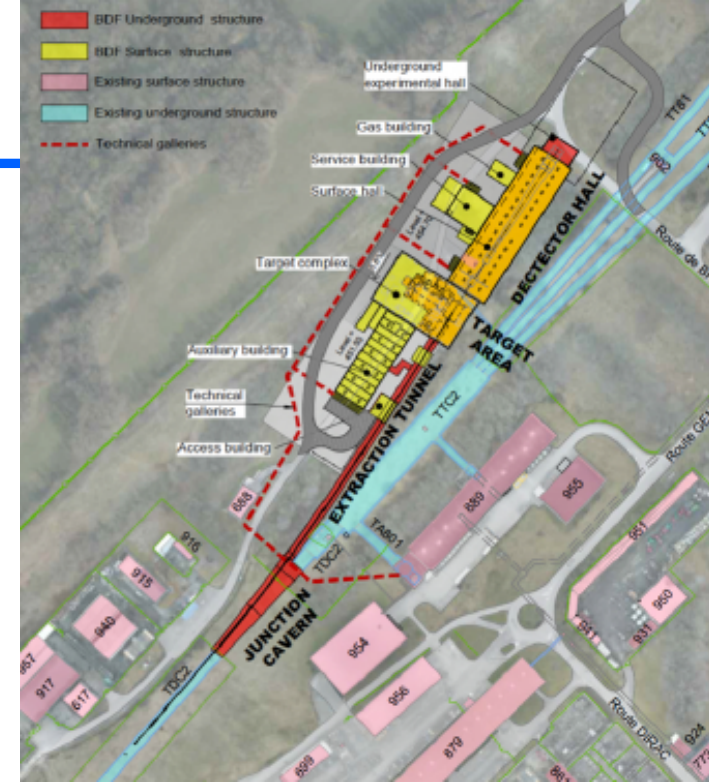


The SHiP experiment

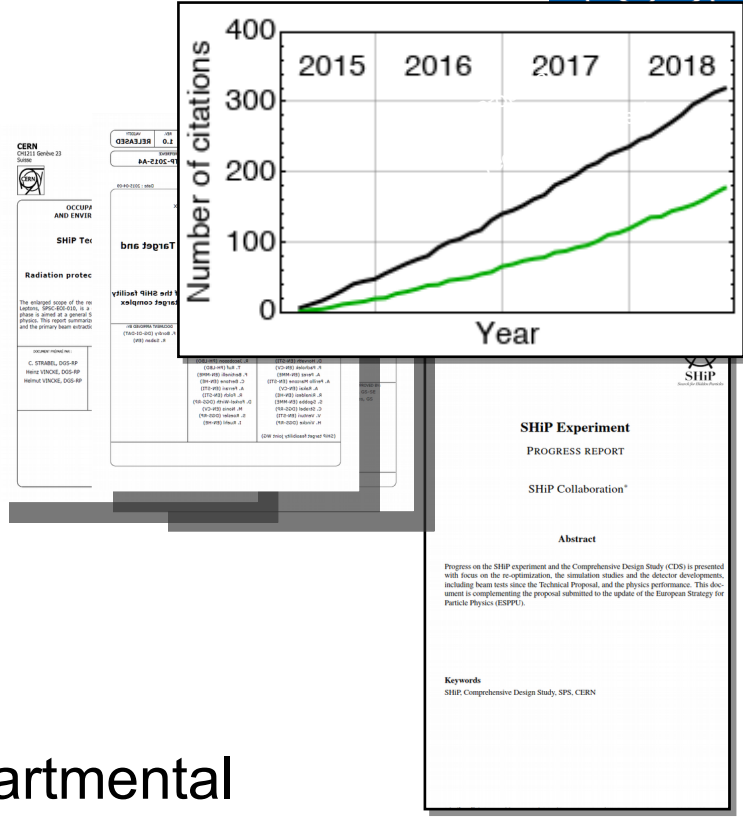
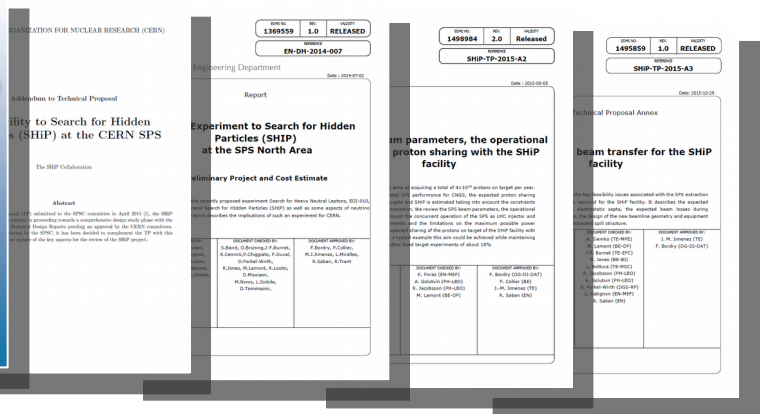
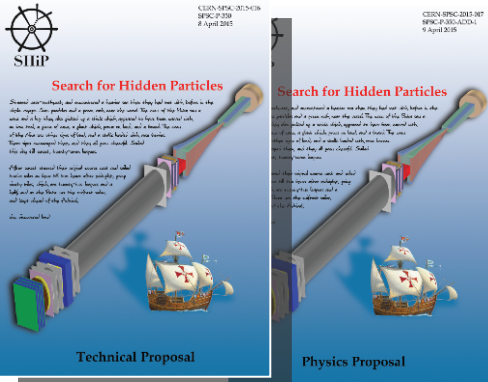
Dedicated detector for weakly coupled long-lived particles, plus tau neutrino and LDM scattering, to be run at future beam-dump facility at CERN.

The spectrometer is located $\sim 100\text{m}$ downstream of the target, after a magnetised muon shield, the scattering and neutrino detector and a long decay volume

Aim for a 0-BG experiment (2 events \rightarrow discovery)



SHiP history



SHiP Experiment
PROGRESS REPORT
 SHiP Collaboration*

Abstract

Progress on the SHiP experiment and the Comprehensive Design Study (CDS) is presented with focus on the re-optimization, the simulation studies and the detector developments, including beam tests since the Technical Proposal, and the physics performance. This document is complementing the proposal submitted to the update of the European Strategy for Particle Physics (ESPPU).

Keywords
 SHiP, Comprehensive Design Study, SPS, CERN

- 2013 Oct: EOI with SHiP@SPS NA
- 2014 Jan: Encouraged to produce TP and inter-departmental task force setup to study feasibility of proposed facility
- 2015 Apr: TP with ~700 pages by SHiP theorists, experimentalists, and CERN accelerator, engineering, and safety departments
- 2016 Jan: Recommendation by CERN SPSC to proceed to 3-year CDS
- 2016 Apr: CERN management launch of Beyond Collider Physics study group
- SHiP experimental facility included under PBC as Beam Dump Facility
- 2018: EPPSU contribution submitted by SHiP and BDF
- 2019 Dec: CDS submitted: CERN-SPSC-2019-049 ; SPSC-SR-263
- SHiP Collaboration: **290 authors, 52 Institutes, 17 countries**

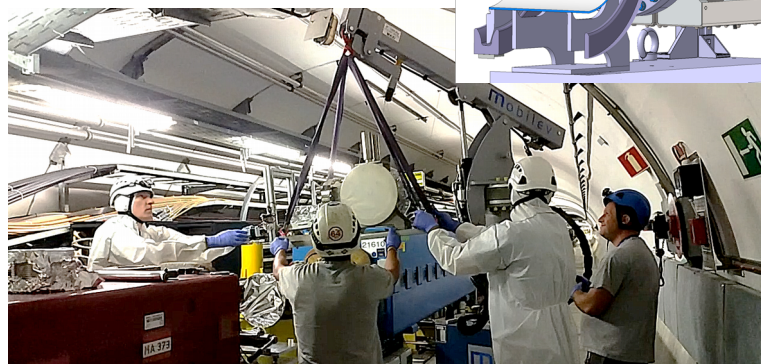
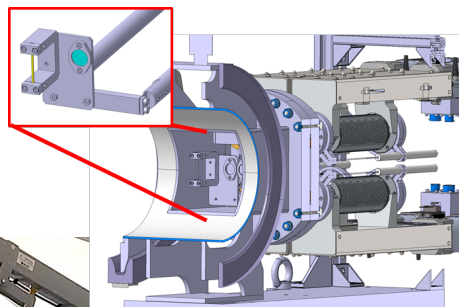
Status of Beam Dump Facility



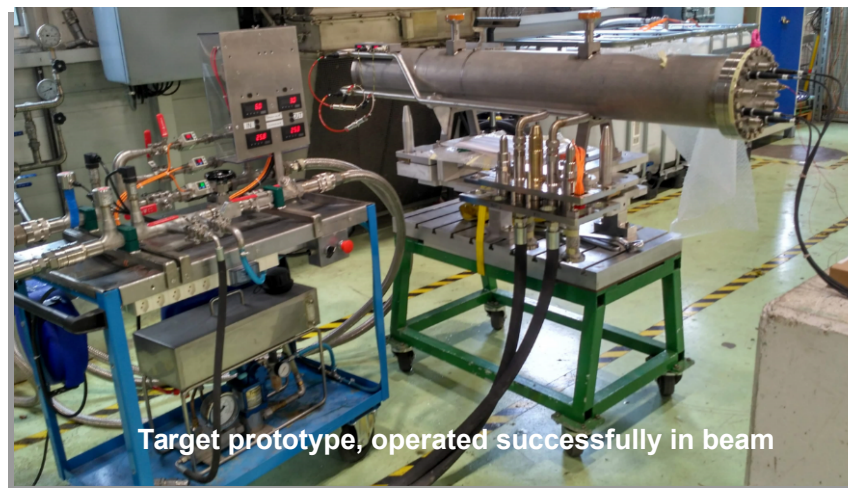
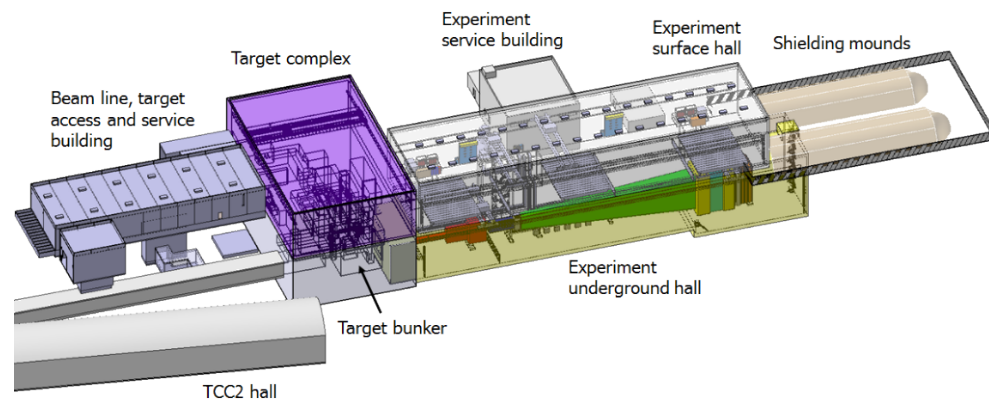
3-year Comprehensive Design Study completed by BDF team

- In-depth feasibility study with prototypes of key elements
 - SPS extraction and proton delivery
 - Target system and target complex, including remote handling
 - Underground experimental area, layout of surface buildings for construction/installation and operation
 - Evaluations of the radiological aspects and safety
 - First iteration of detailed integration and civil engineering studies
 - Updated realistic schedule and cost, detailed project plan and resources for TDR phase
- ➔ Documented in 580-page Yellow Report
- ➔ BDF ready for 3-year TDR phase

A few high-lights:



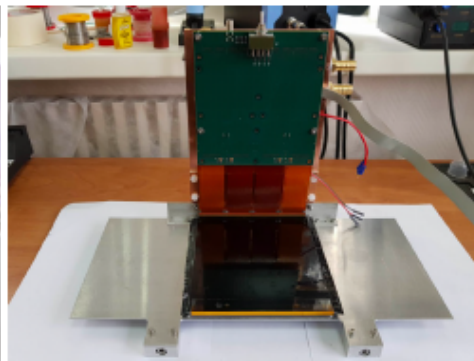
Crystal shadowing of extraction septum wires combined with improvements of beam dynamics and automated alignment achieved factor 3-4 less losses in SPS extraction, validating the SHiP requirements



Target prototype, operated successfully in beam

Current status of the experiment

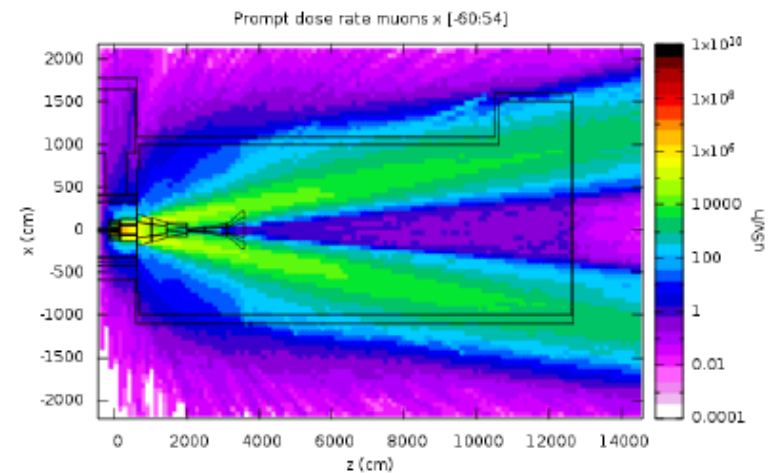
- Collaboration completed Comprehensive Design Study, then we expect to be requested a TDR
- Phase-1 prototypes for all sub-detectors built and tested on a beam in summer 2018
- From the summer 2019 ECFA newsletter:
 - Amongthem, the **SPS Beam Dump Facility with the SHiP** and (possibly) the TauFV experiment has been identified as having unique potential in the worldwide landscape for dark photon and heavy neutral lepton searches, as well as for third flavour physics ($\nu\tau$ interactions and τ rare decays). It is now mature and ready for an implementation decision pending the Strategy guidelines.
- Phase-2 prototypes under construction, to be tested on beam in 2019-21.



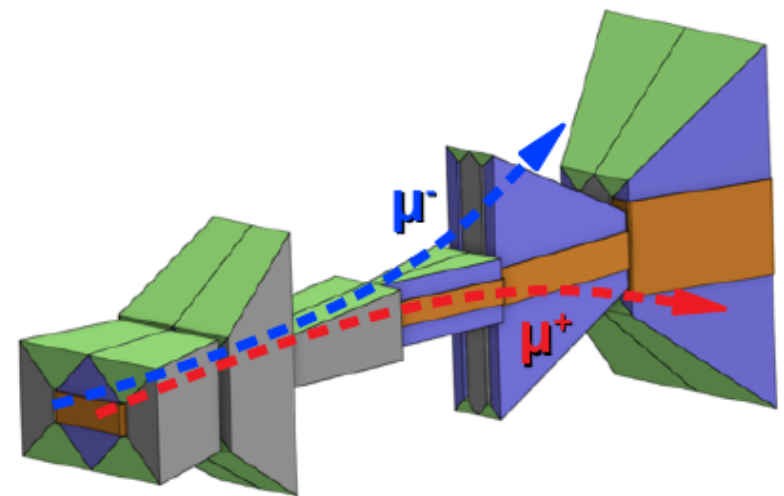
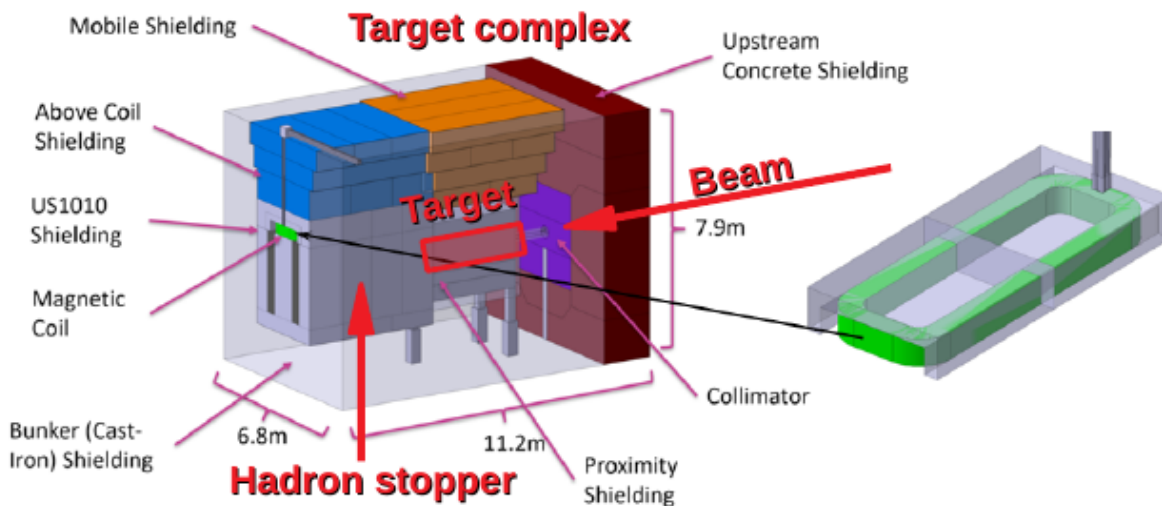
Target and shielding

[JINST 12(2017)05 P05011]

- ▶ heavy target to absorb π s before decay
- ▶ magnetized hadron stopper: immediately separate μ^\pm
- ▶ ideal muon shield configuration optimised with machine learning
 $\implies \mu$ rate reduced to ~ 25 kHz



- ▶ μ spectrum validated with dedicated experiment in 2018



Magnetisation of hadron stopper



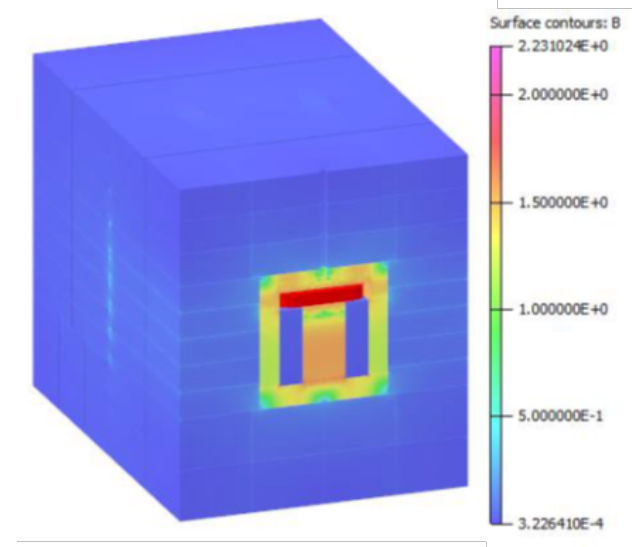
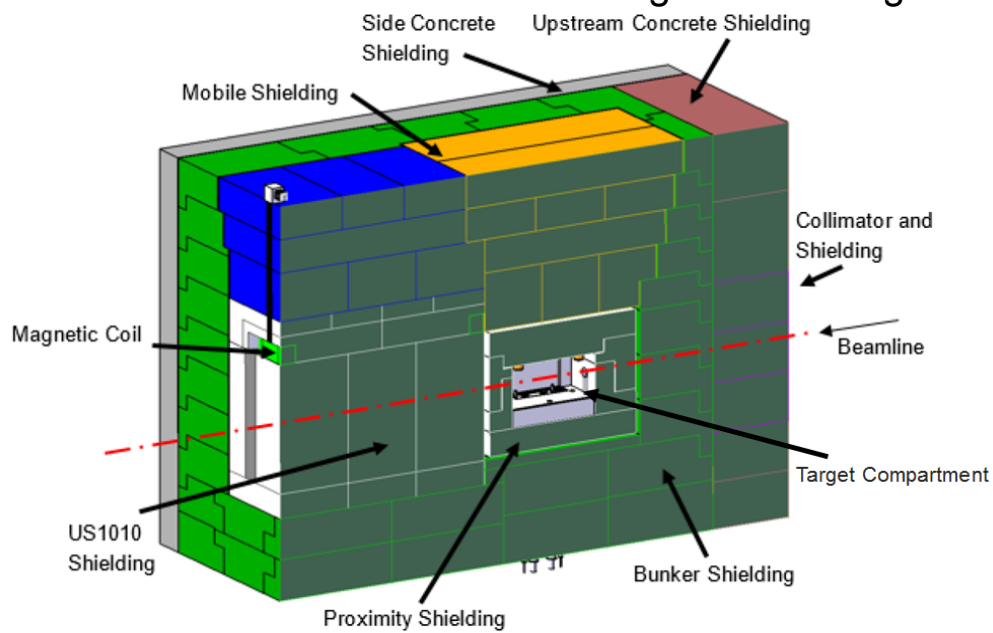
Detailed design study completed by RAL (V. Bayliss, J. Boehm, G. Gilley) through Collaboration Agreement with CERN

- Optimisation of the magnetic circuit

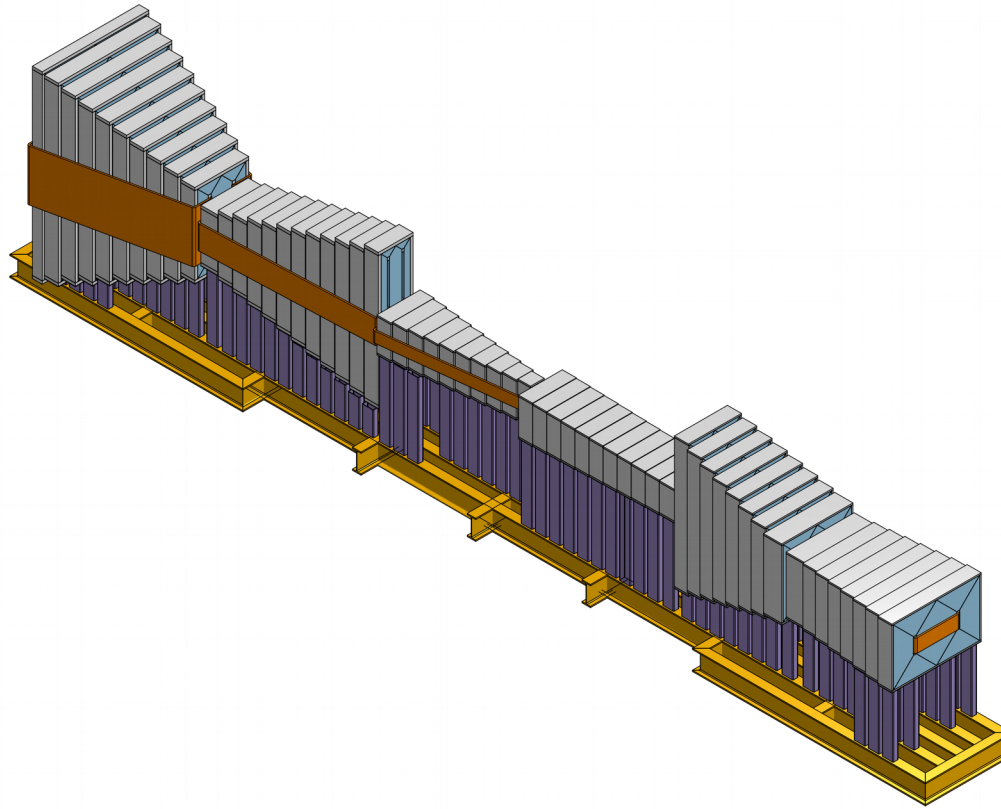
- Simulated field maps for use in physics simulations and for optimisation of the subsequent free-standing muon shield
- Hysteresis effects after multiple powering cycles;
- Magnetic forces of the entire magnetized assembly and target shielding
- Stray fields

- Preliminary engineering design compatible with the target complex and radiation environment

- Power requirements
- Thermal management (consideration of water and gas cooling)
- Technical solution for connections of power cables, cooling, sensors etc.
- Technical solution for the integration of magnetic iron blocks and remote handling of blocks and coils



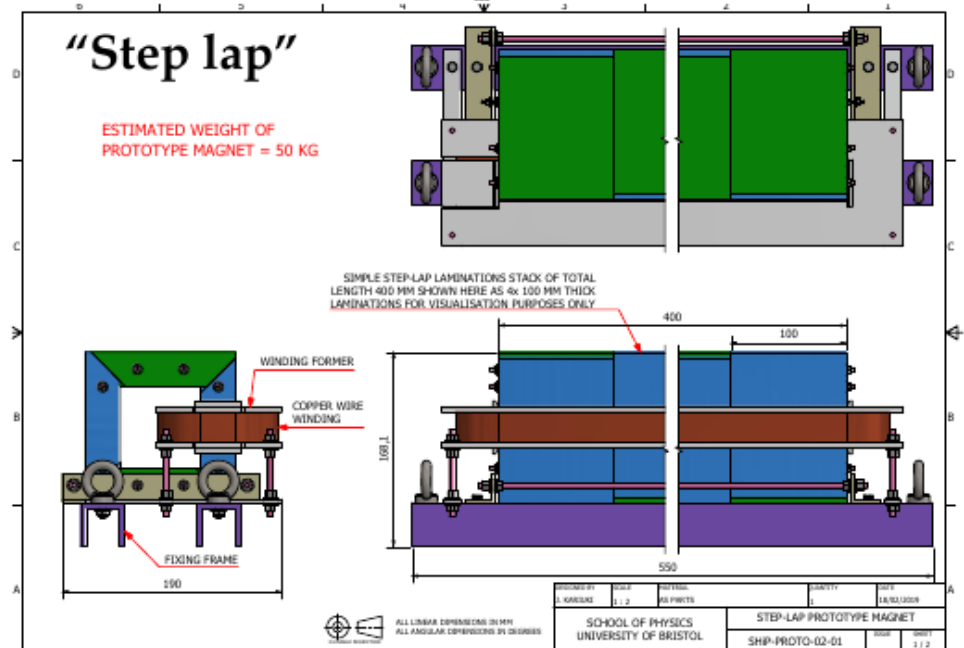
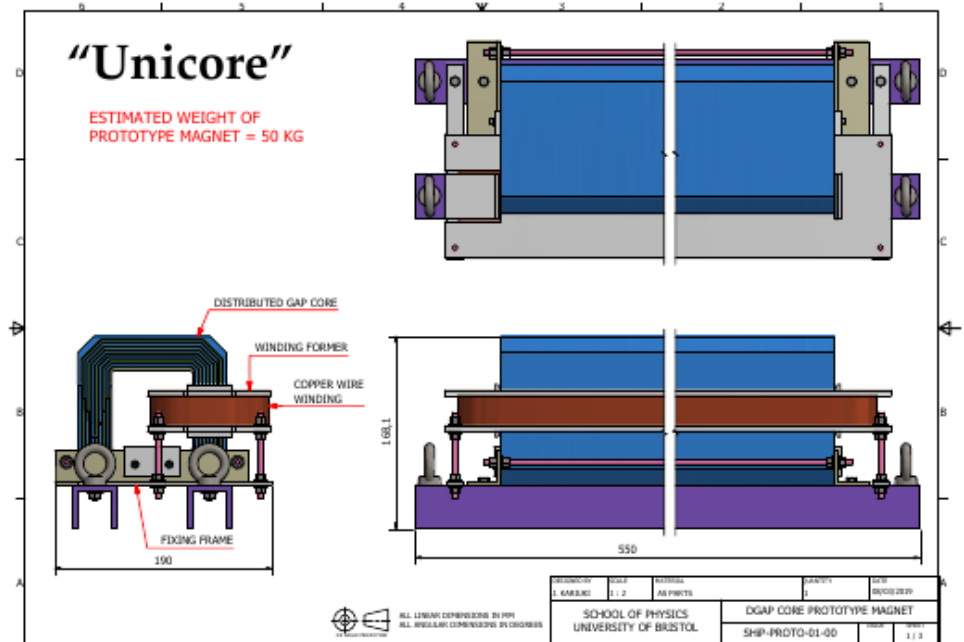
Magnetic Shield for SHiP (UK-Russia responsibility)



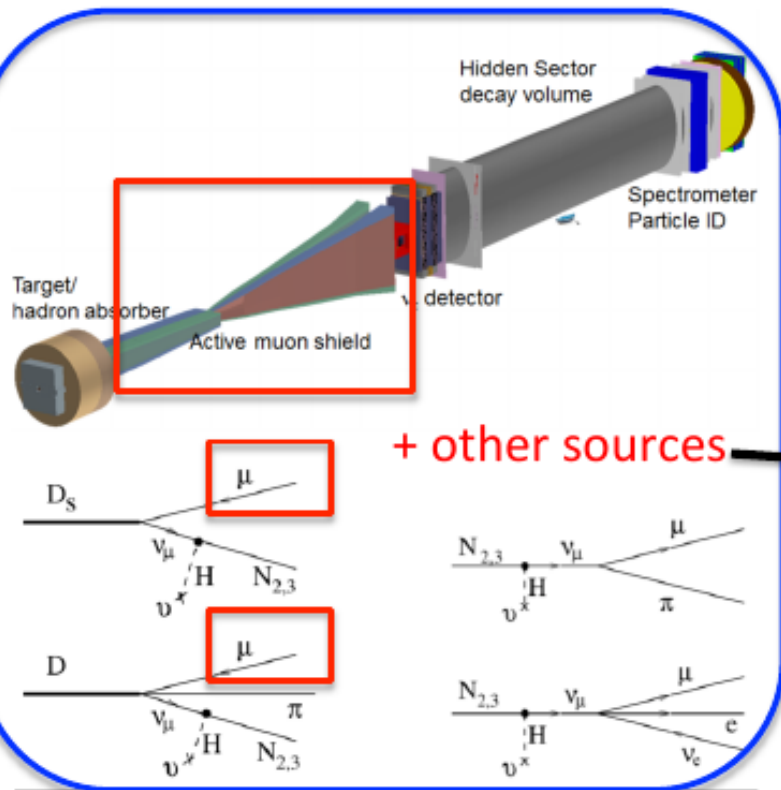
- ◇ about 600 individual modules (one block in the figure is 10 modules)
- ◇ total weight of about 10000 tons
- ◇ modules up to 6.5×4 m² in size
- ◇ about 2000 km of sheet cutting length

Magnet prototypes

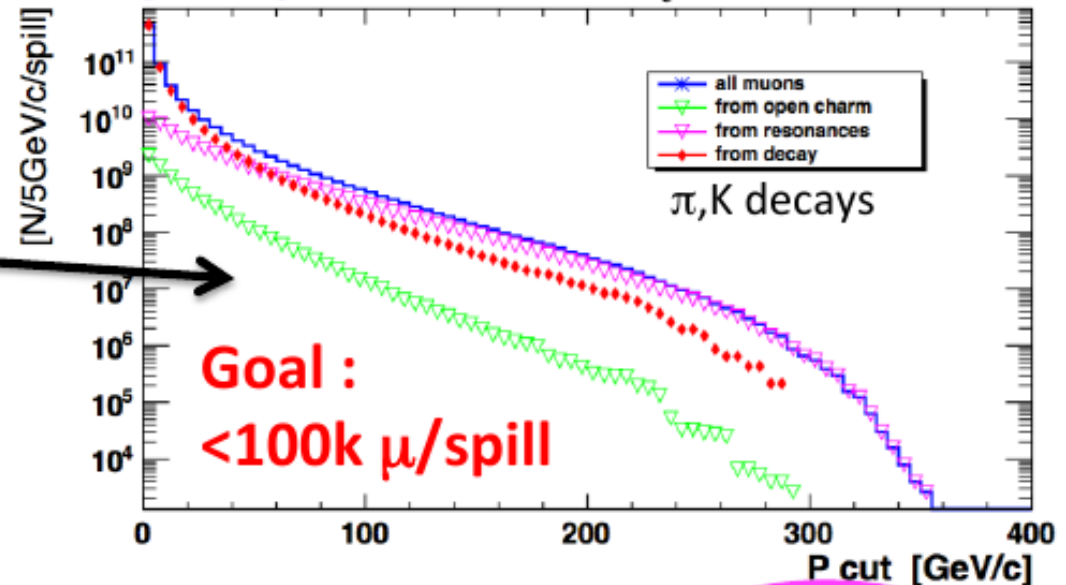
- Two 50kg magnet prototypes have been constructed with different assembly techniques by UK companies using UK Grain Oriented Steel
- Complementary approach to welding of laminations
- Target field of $\sim 1.8\text{T}$ with stacking factor >0.95
- Test for losses of magnetic flux around the loop due to the different types of connections between laminations
- Awaiting delivery to CERN for testing



The muon filter

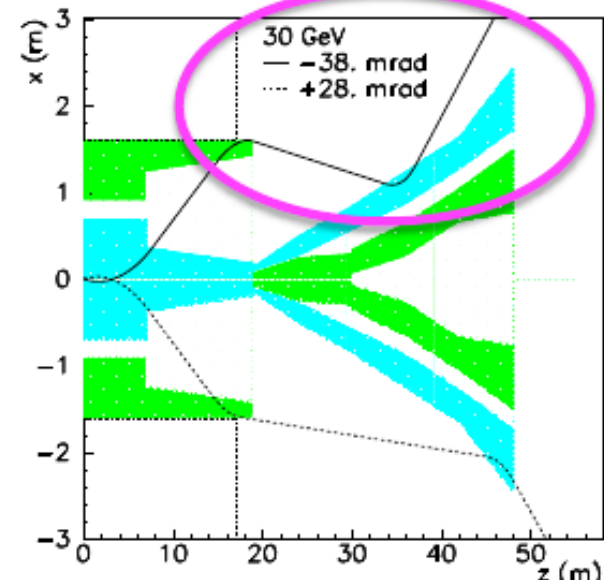
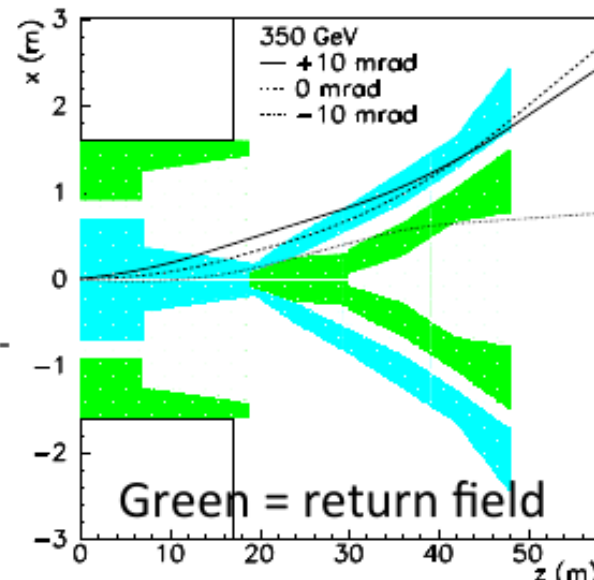


- Clear the μ flux to a level inferior to the ν -induced background and not to compromise the **occupancy limit** in the ν_τ detector

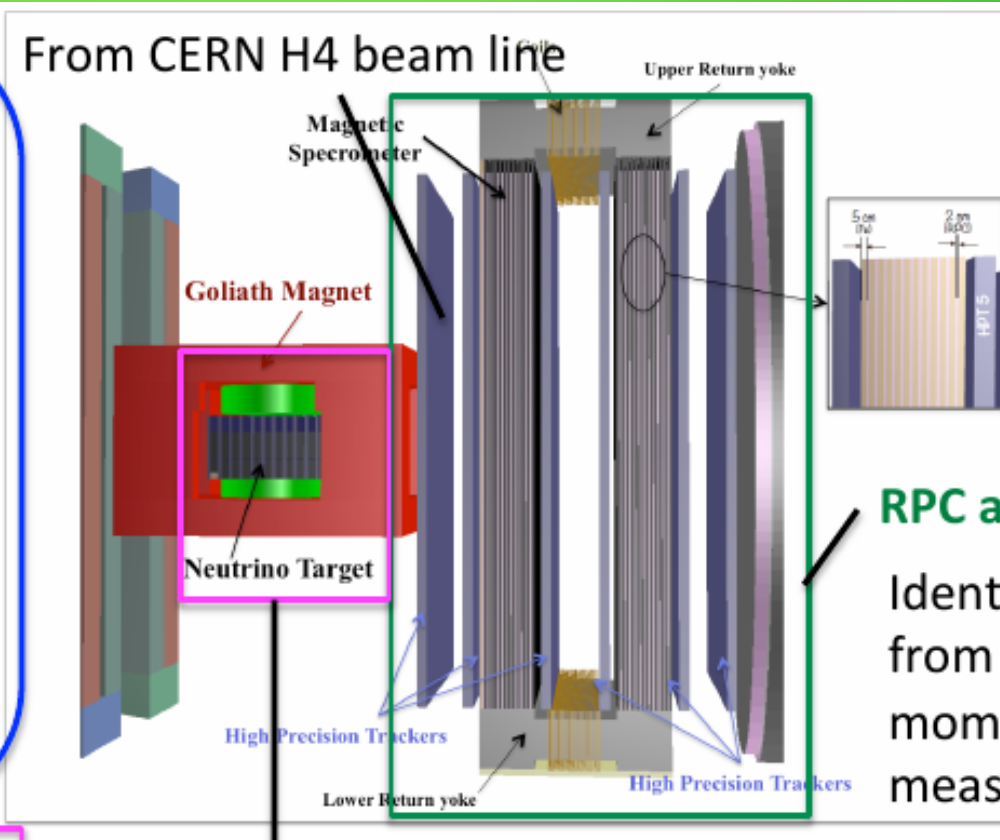
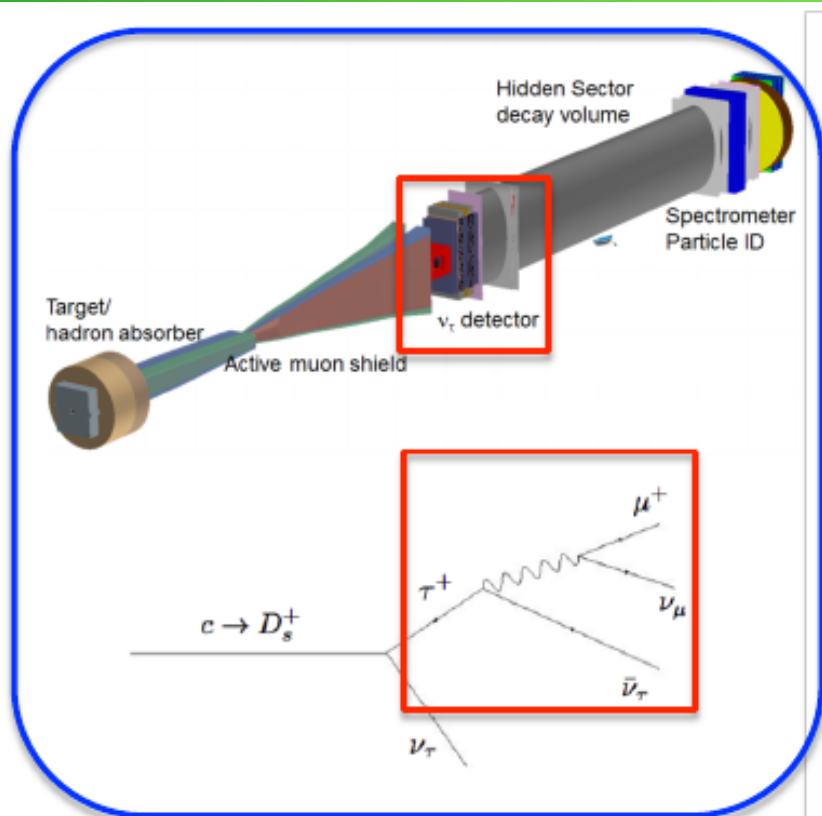


- **Active** muon filter made of two magnets ($B_y = 40 \text{ Tm}$ to bend out 350 GeV μ beyond the 5 m vacuum vessel aperture)

- **1st part** to separate μ^+/μ^-
- **2nd part** to bend the muons **further outward**

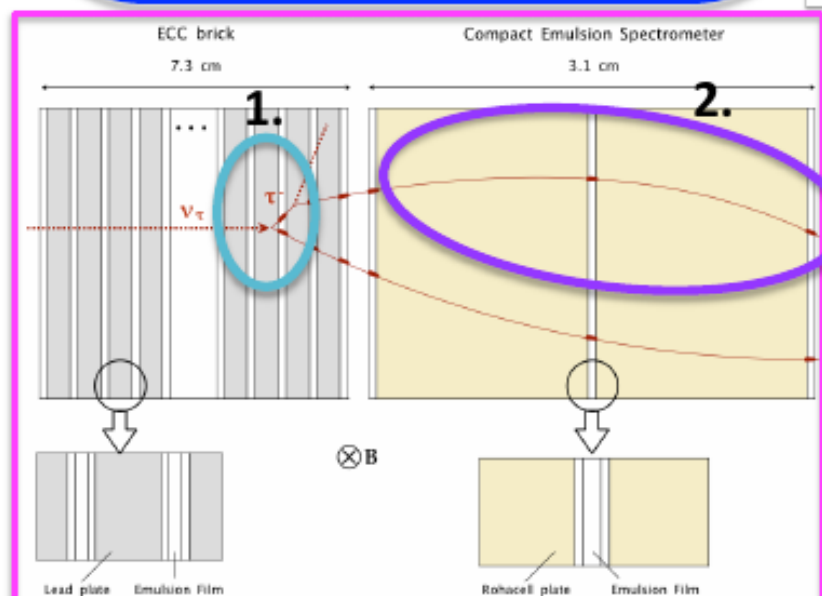


Bonus intermezzo: the ν_τ detector



μ flux :
 $10^3 \mu/\text{mm}^2$
 in 6 months

Identify μ coming from τ decays; μ momentum measurement

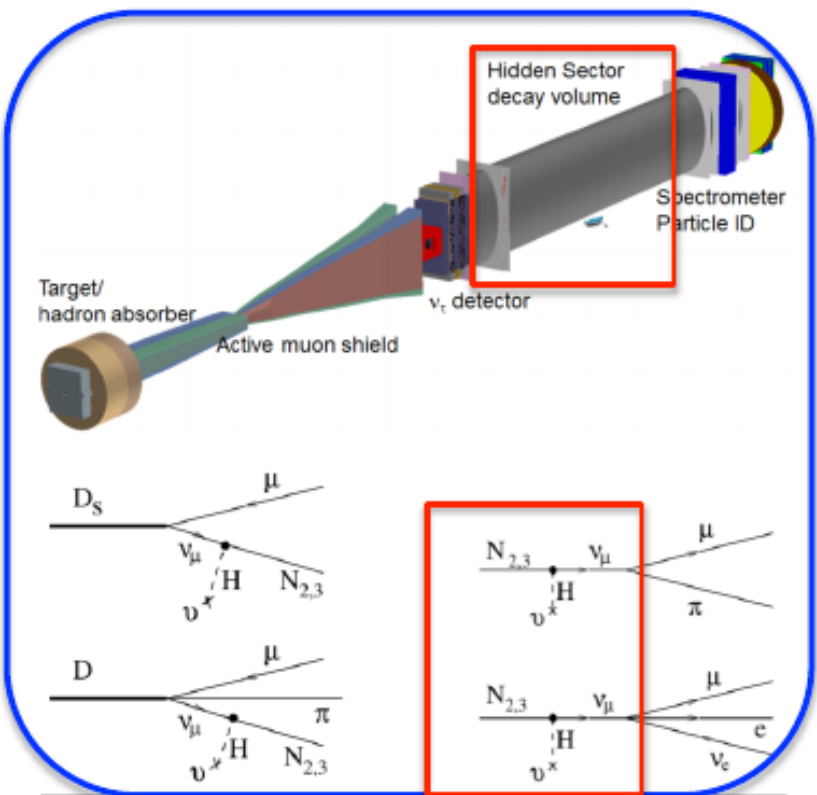


OPERA-type modules (emulsion cloud chambers (ECC))

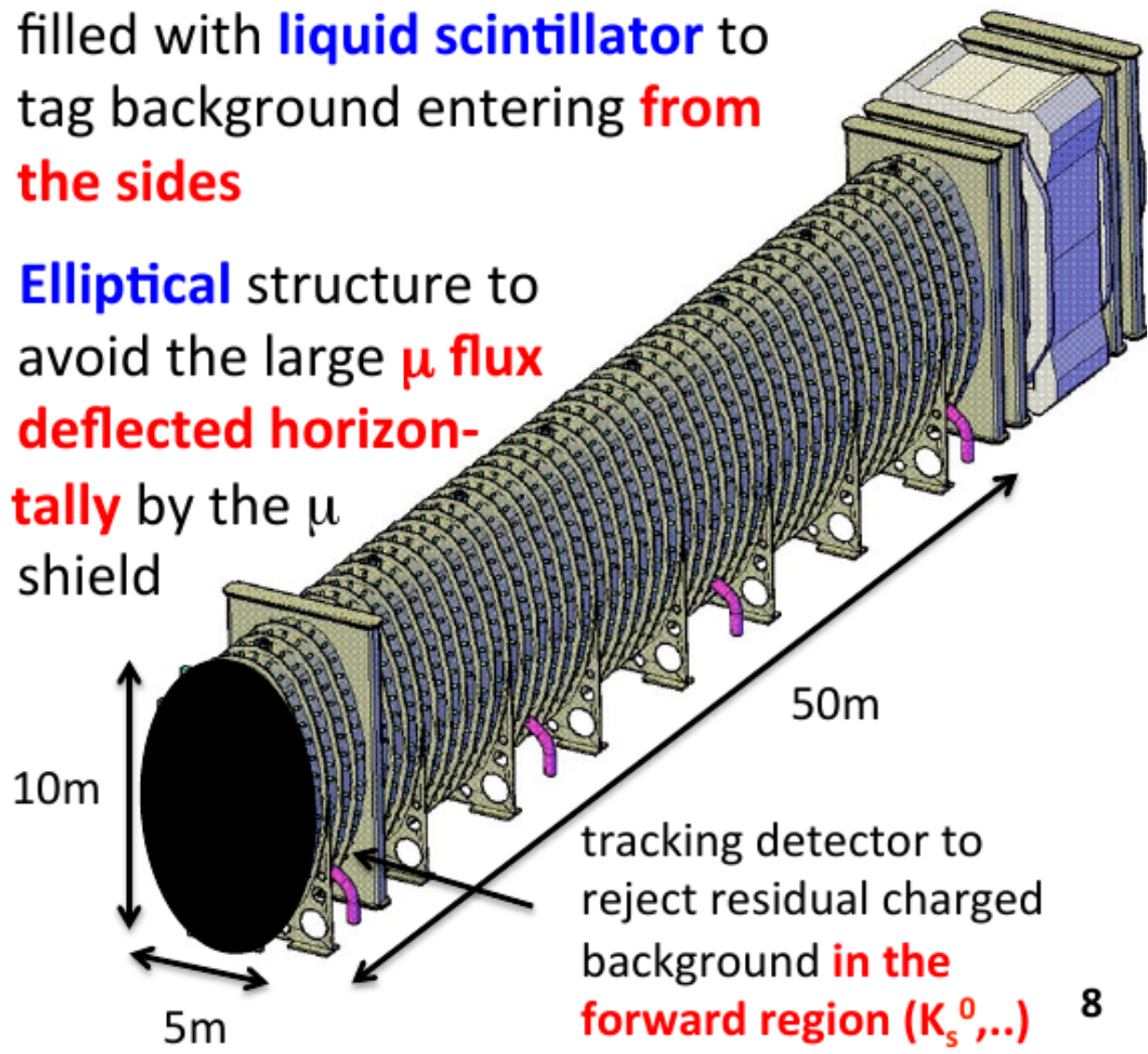
- 1x1m, close to the beam line
- 2 parts : ECC brick and compact spectrometer

1. ν interaction : identified through detection of τ production and decay (contained within a brick due to short τ lifetime)
2. distinguish ν_τ anti- ν_τ : Look at the charge of the τ decay products

The vacuum vessel

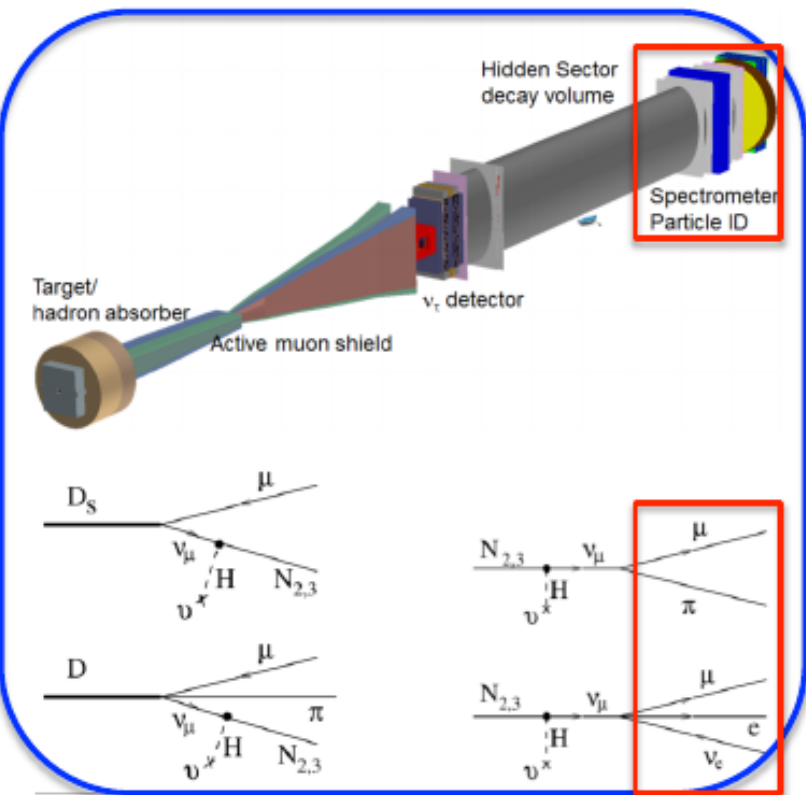


- **10^{-6} mbar vacuum** needed suppress the **neutrino interactions**
- Double-wall structure, space filled with **liquid scintillator** to tag background entering **from the sides**
- **Elliptical** structure to avoid the large **μ flux deflected horizontally** by the μ shield



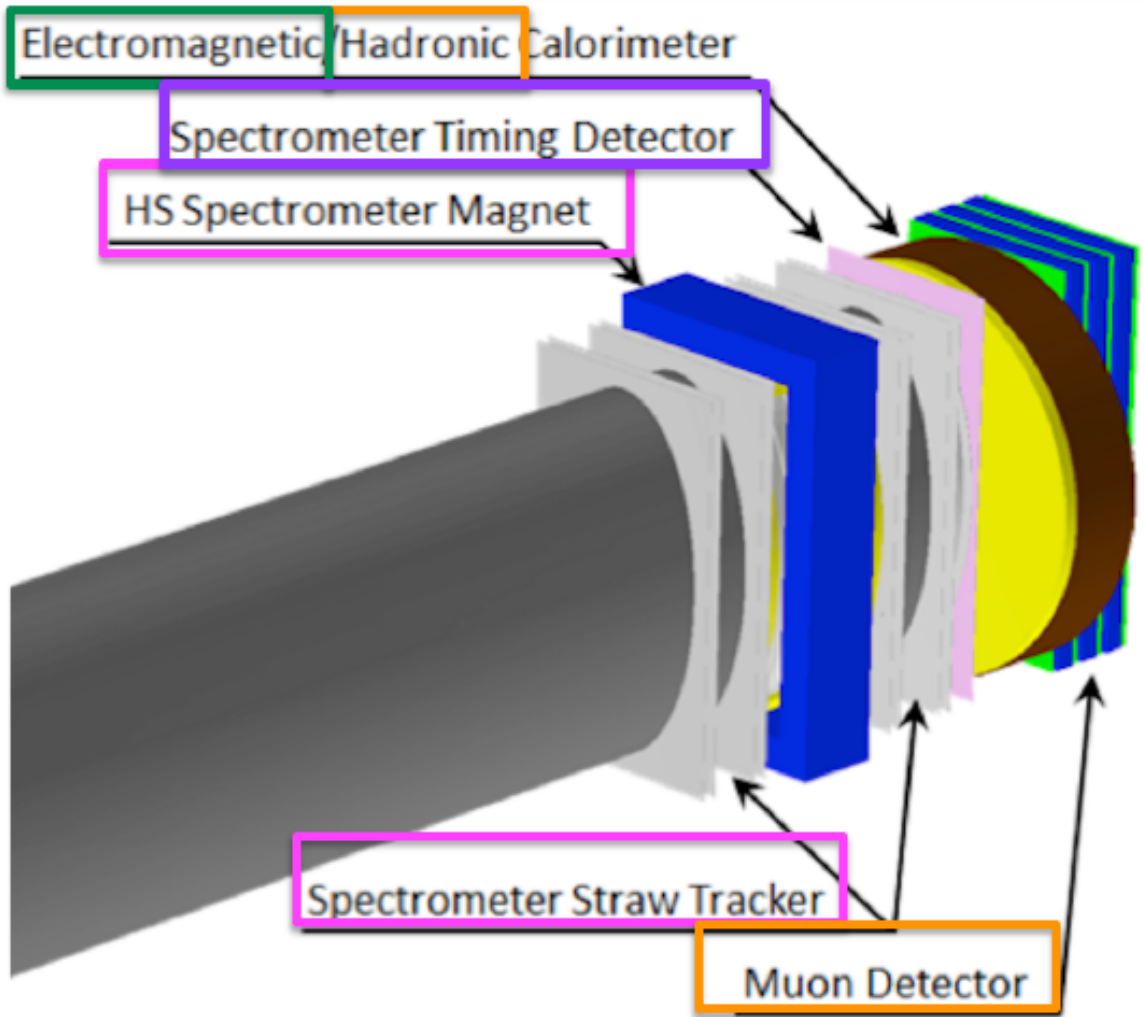
- **Veto tagger** located just after ν_τ detector to tag indirectly **neutral K** produced by ν and μ interactions in the passive material of the ν_τ detector and **μ entering the vessel from the front**

The spectrometer



Signal reconstruction and background rejection: warm magnet (LHCb) with 0.65Tm bending power; tracker (NA62) with horizontal straws and stereo angle

Veto anti-coincidence from combinatorial : timing detector (50ps resolution)

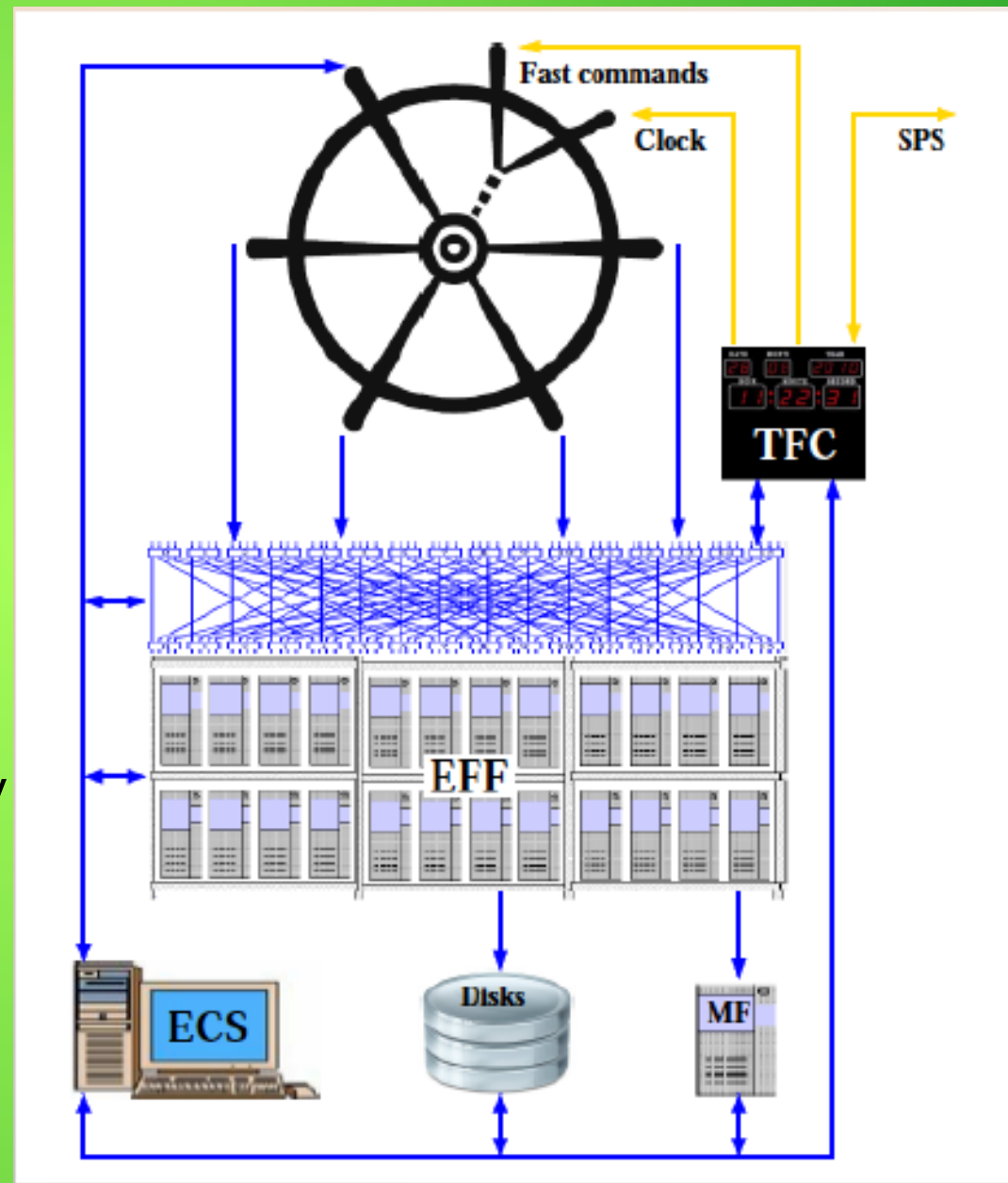


e/γ identification, π^0 and η reconstruction: ECAL (Shashlik technique, LHCb)

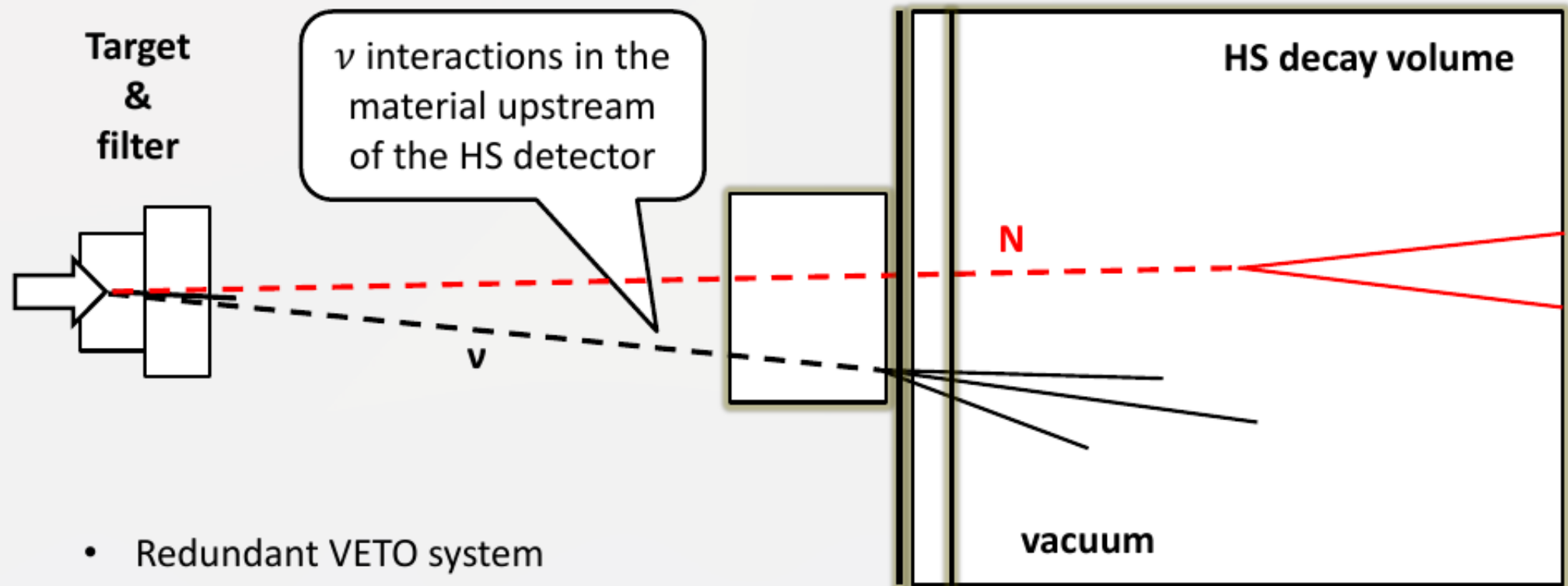
π/μ separation : hadronic calorimeter (similar technology as ECAL), muon detector (WLS fiber bars, MINOS)

Trigger and DAQ

- Trigger and Event building on all data and trigger decision at EF
- TFC system generates the clock
- All sub-systems send data through ethernet links (no need for radiation hardness) to Event Filter Farm via a switch
- Fraction of data sent to Monitoring
- Farm to evaluate performance
- Smallest time slice that could potentially contain all data from one pot (100 ns)
- Since some events spread over more than one frame, 100 frames are combined into a “package”, with 1 overlap



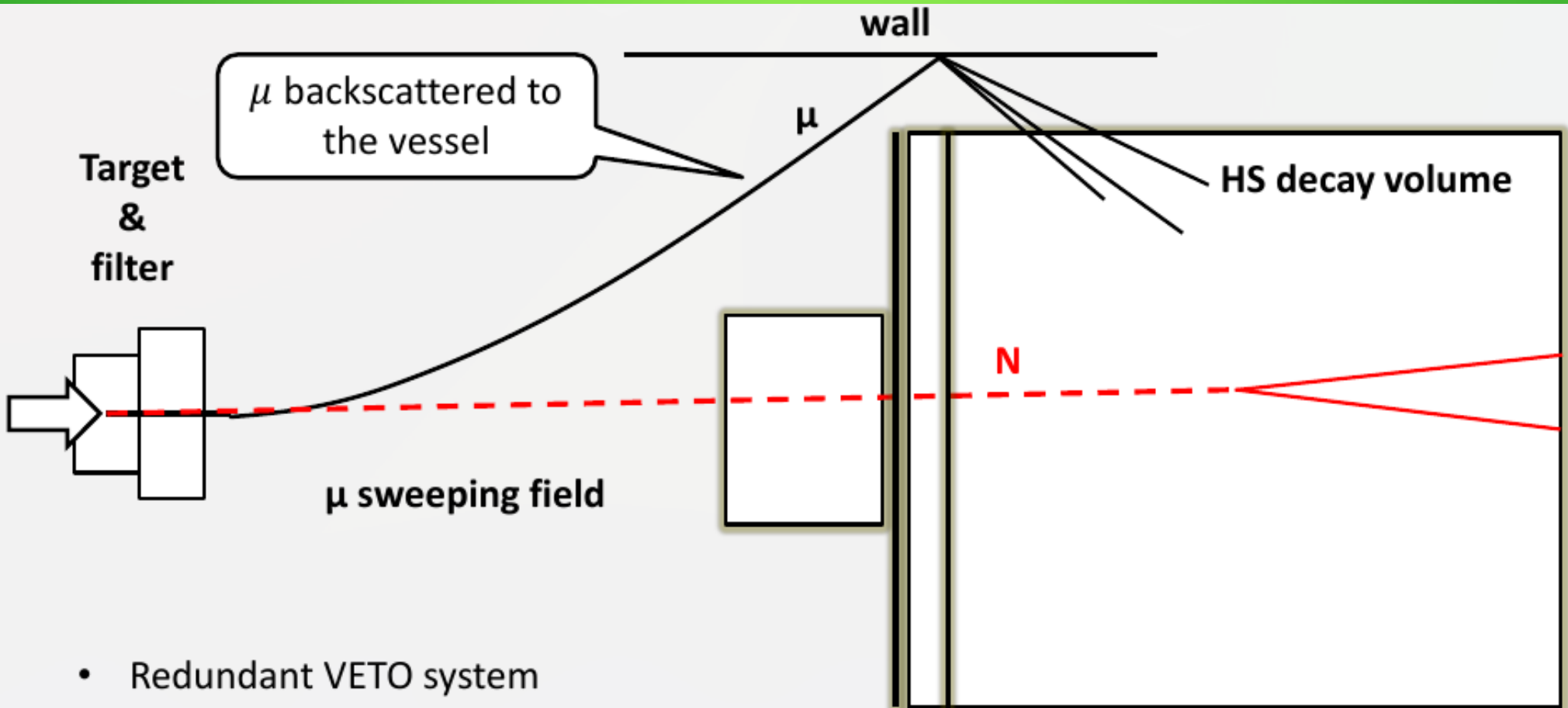
Background rejection: upstream neutrino interactions



- Redundant VETO system
- Combinatorial rejected by timing detector
- Impact parameter to the target
- 75% selection efficiency for signal

After selections:
 $\leq 0.1 \text{ bkg} / 5 \text{ y}$

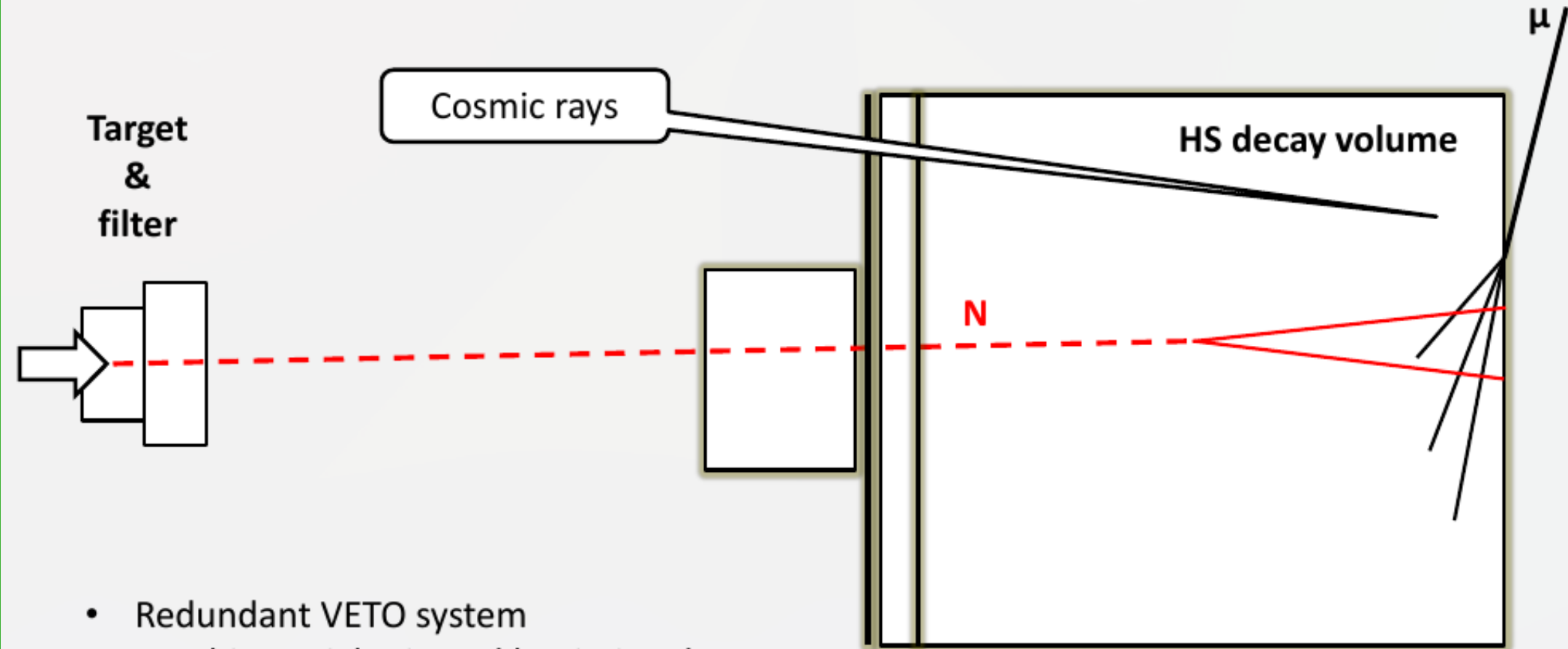
Background rejection: interactions with experimental hall



- Redundant VETO system
- Combinatorial rejected by timing detector
- Impact parameter to the target
- 75% selection efficiency for signal

After selections:
 $\leq 0.1 \text{ bkg} / 5 \text{ y}$

Background rejection: cosmic

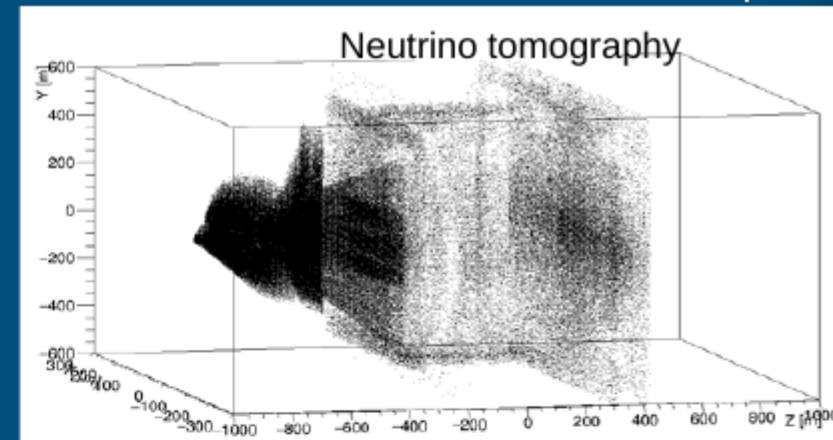


- Redundant VETO system
- Combinatorial rejected by timing detector
- Impact parameter to the target
- 75% selection efficiency for signal

After selections:
 $\leq 0.1 \text{ bkg} / 5 \text{ y}$

Backgrounds: summary

Background source	Decay modes
ν or μ + nucleon $\rightarrow X + K_L$	$K_L \rightarrow \pi e \nu, \pi \mu \nu, \pi^+ \pi^-, \pi^+ \pi^- \pi^0$
ν or μ + nucleon $\rightarrow X + K_S$	$K_S \rightarrow \pi^0 \pi^0, \pi^+ \pi^-$
ν or μ + nucleon $\rightarrow X + \Lambda$	$\Lambda \rightarrow p \pi^-$
n or p + nucleon $\rightarrow X + K_L$, etc	as above



Background summary: no evidence for any irreducible background

- No events selected in MC \rightarrow Expected background UL @ 90% CL

Background source	Stat. weight	Expected background (UL 90% CL)
ν-induced		
$2.0 < p < 4.0$ GeV/c	1.4	1.6
$4.0 < p < 10.0$ GeV/c	2.5	0.9
$p > 10$ GeV/c	3.0	0.8
$\bar{\nu}$-induced		
$2.0 < p < 4.0$ GeV/c	2.4	1.0
$4.0 < p < 10.0$ GeV/c	2.8	0.8
$p > 10$ GeV/c	6.8	0.3
Muon inelastic	0.5	4.6
Muon combinatorial	–	<0.1
Cosmics		
$p < 100$ GeV/c	2.0	1.2
$p > 100$ GeV/c	1600	0.002

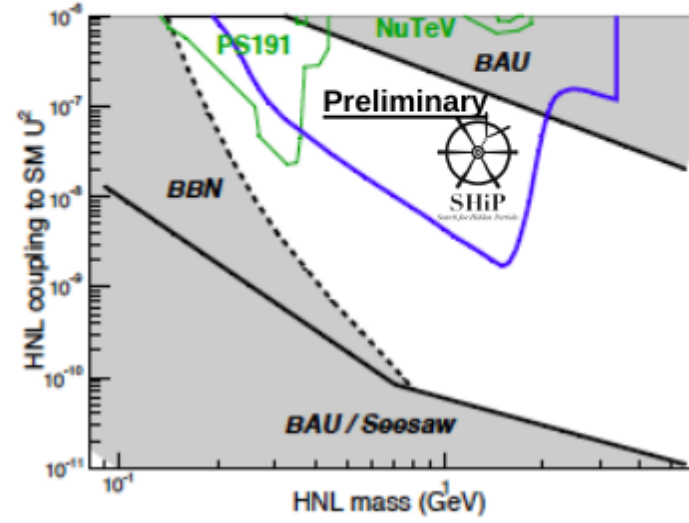
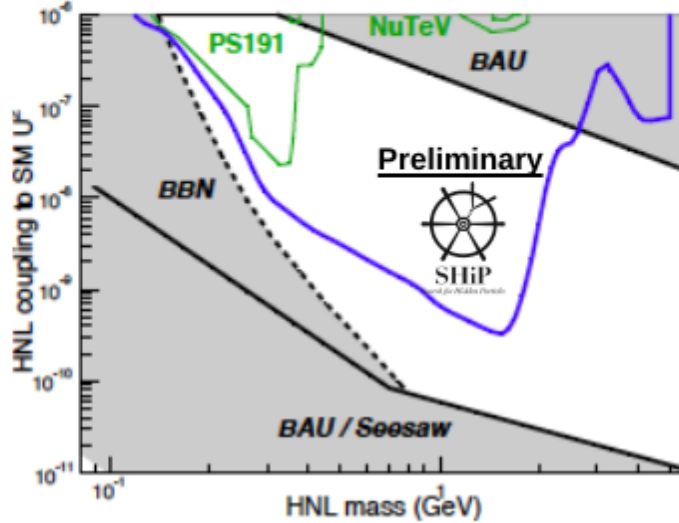
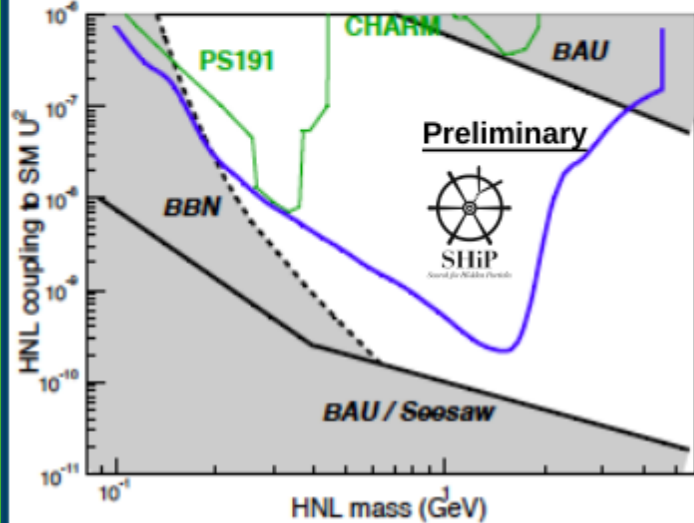
Sensitivity to HNL

- Visible decays = At least two tracks crossing the spectrometer
 - Ex. For $m_N = 1$ GeV with $U^2 = 10^{-8}$ and $BR(N \rightarrow \mu\pi) = 20\%$, expect ~ 330 signal events

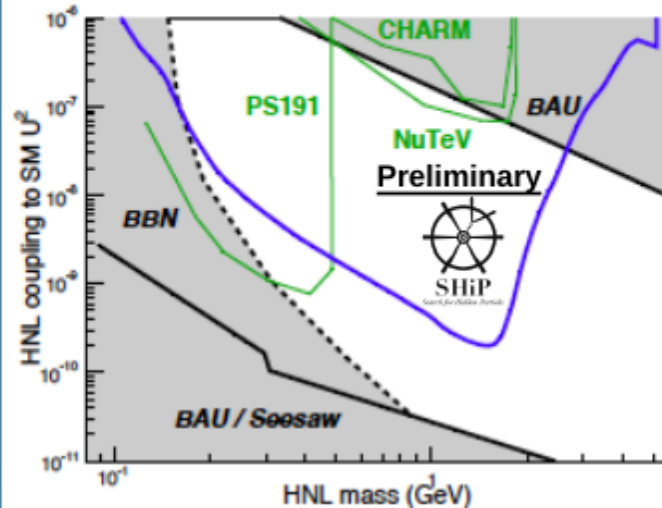
$U_e^2:U_\mu^2:U_\tau^2 \sim 52:1:1$, inverted hierarchy

$U_e^2:U_\mu^2:U_\tau^2 \sim 1:16:3.8$, normal hierarchy

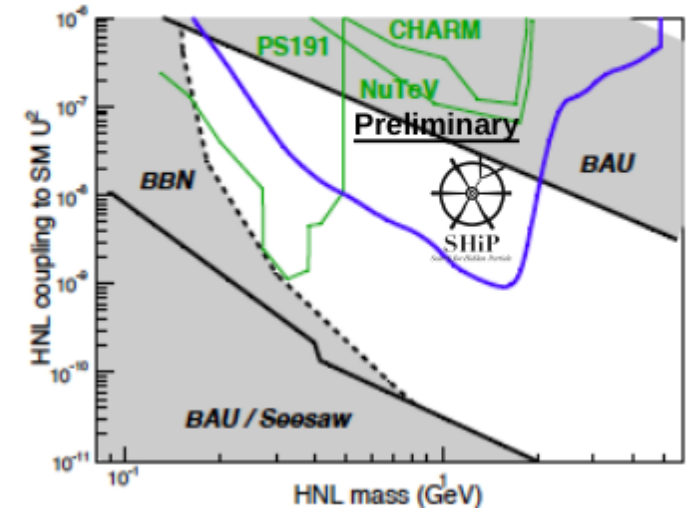
$U_e^2:U_\mu^2:U_\tau^2 \sim 0.061:1:4.3$, normal hierarchy



$U_e^2:U_\mu^2:U_\tau^2 \sim 48:1:1$, inverted hierarchy



$U_e^2:U_\mu^2:U_\tau^2 \sim 1:11:11$, normal hierarchy

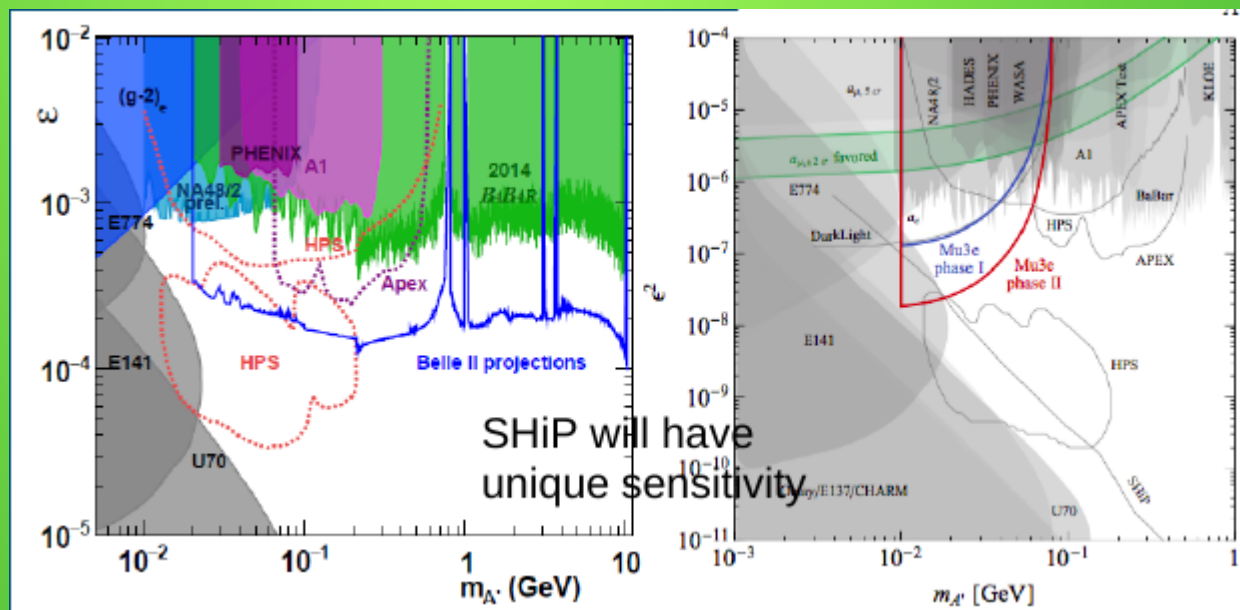
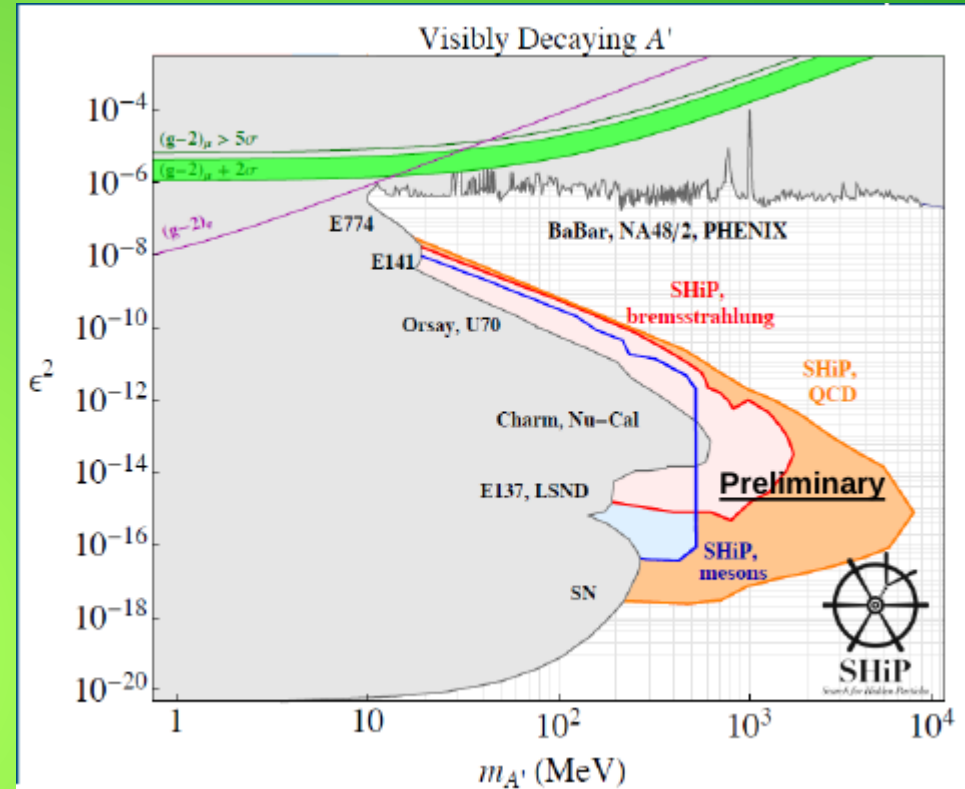


Scenarios for which baryogenesis was numerically proven

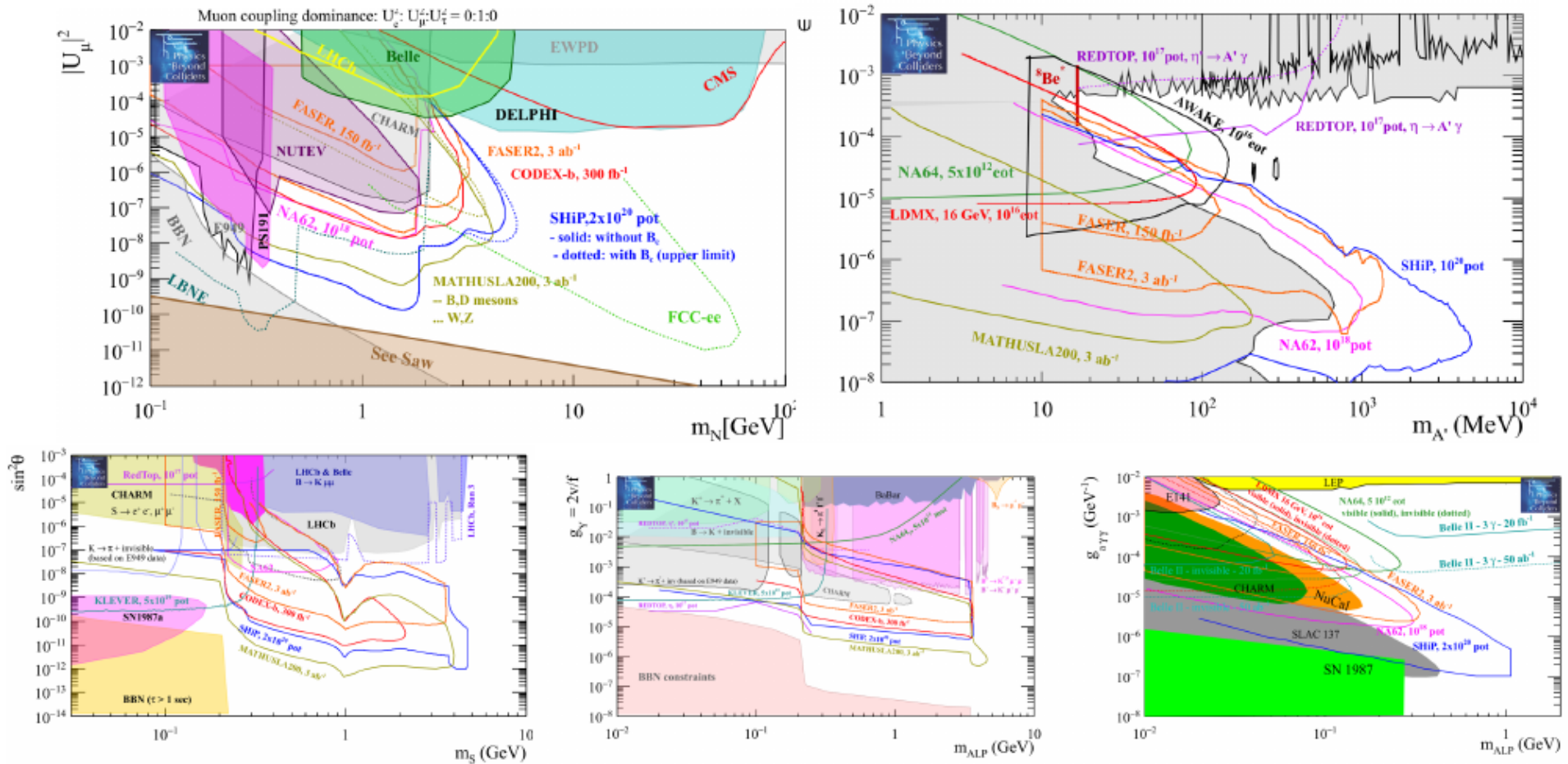
Sensitivity to dark photons

Production

- Decays of $\pi^0 \rightarrow V\gamma$, $\eta \rightarrow V\gamma$, $\omega \rightarrow V\pi^0$
- Proton bremsstrahlung and parton bremsstrahlung above Λ QCD
- Decay into pair of SM particles



Updated physics reach



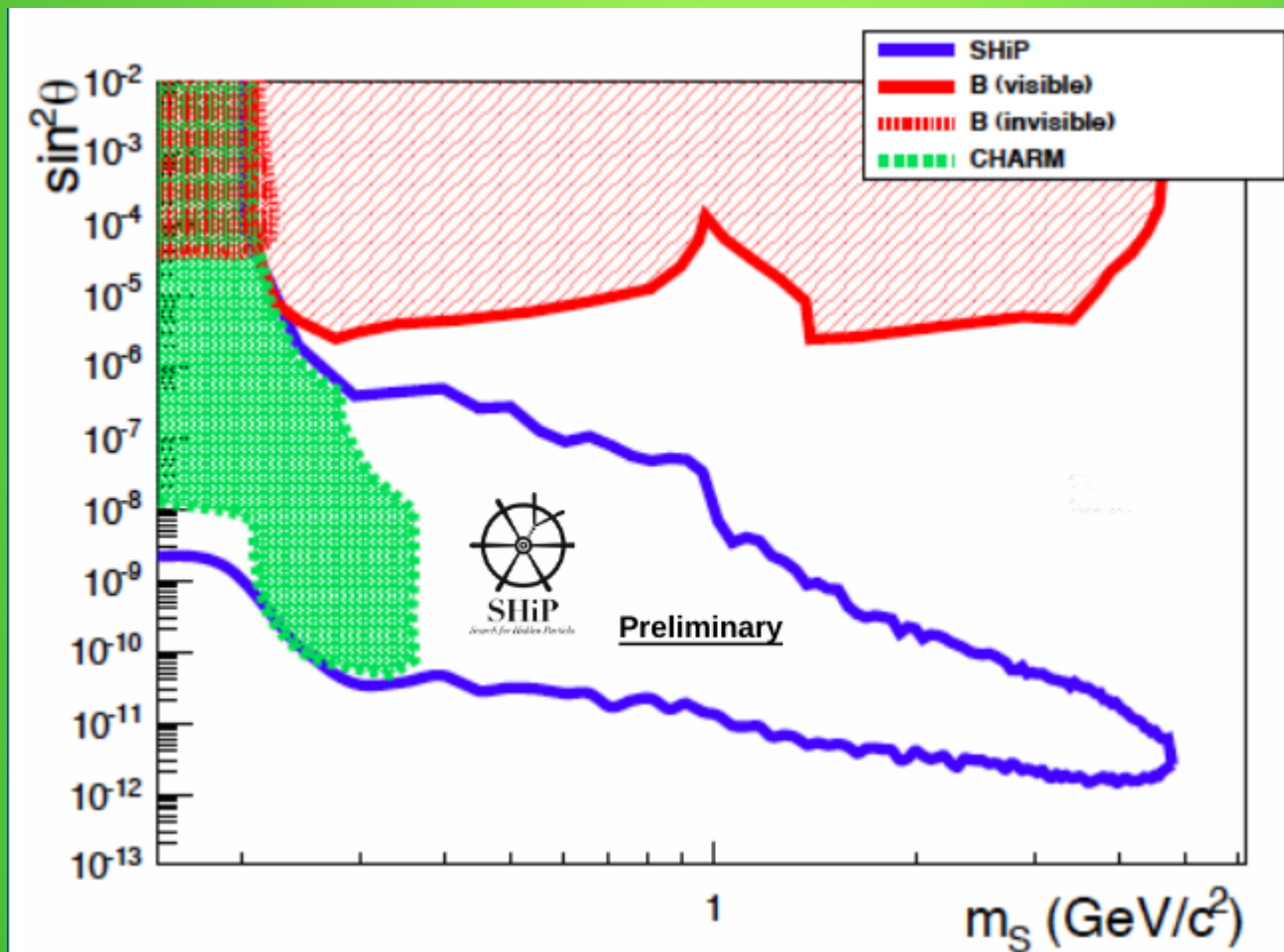
[1504.04956, 1504.04855, 1811.00930, 1901.09966]

- ▶ from top left: **HNL** (heavy meson decays), **dark photon** (decays + bremsstrahlung + QCD), **scalar** (K and B decays), **ALPs** coupled to fermions, **ALPs** coupled to photons

Hidden scalars

Production from B and K decays

Decay into fermion or meson pairs



Axion portal

○ Axion Like Particles, pseudo-scalars pNGB, axial vectors a

- Appear in extended Higgs, SUSY breaking, motivated by coupling with dark sector, possibility of inflaton, etc
- Generically light pseudo-scalars arise in spontaneous breaking of approximate symmetries at a high mass scale F
 - Couplings suppressed by the breaking scale F and masses are light $\sim \Lambda/F^2$
- SM portal through mixing with gauge bosons and fermions

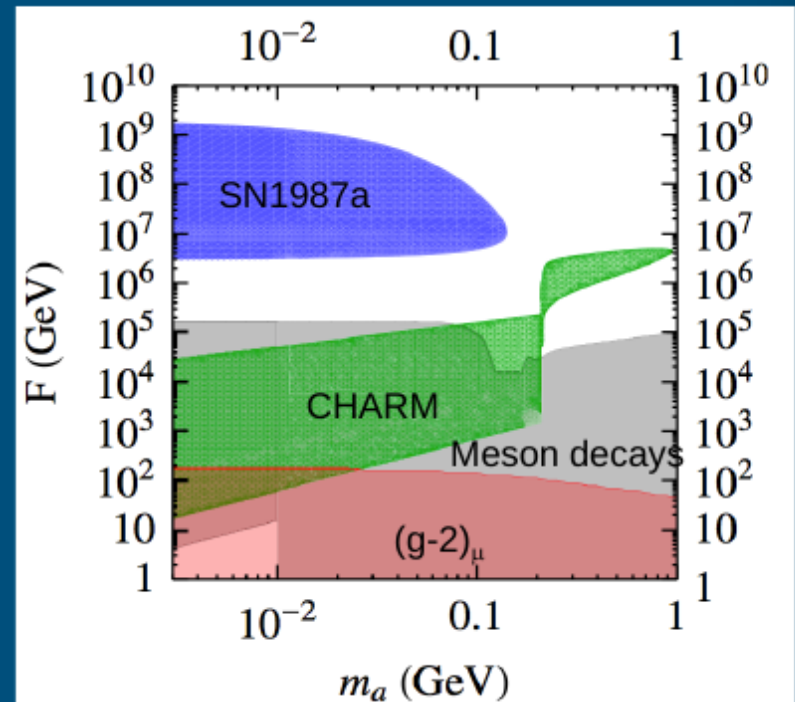
$$\mathcal{L} = \frac{a}{F} G_{\mu\nu} \tilde{G}^{\mu\nu}, \frac{\partial_\mu a}{F} \bar{\psi} \gamma_\mu \gamma_5 \psi, \text{ etc}$$

○ Production

- Resonant production from Drell-Yan photons
- Production from mixing with pions and heavy meson decays

○ Decays

- Decays to e^+e^- , $\mu^+\mu^-$, hadrons above 1 GeV
- Decays to photon pair



Tau neutrino physics

Charged current neutrino nucleon scattering

neutrino scattering

anti-neutrino scattering

$$\frac{d^2\sigma}{dx dy} = \frac{G_F^2 M_N E_\nu}{\pi} \left(\frac{M_W^2}{Q^2 + M_W^2} \right)^2 \left[\left(xy^2 + \frac{m_l^2 y}{2E_\nu M_N} \right) F_1 + \left(1 - y - \frac{M_N xy}{2E_\nu} - \frac{m_l^2}{4E_\nu^2} \right) F_2 \right. \\ \left. \pm \left(xy \left(1 - \frac{y}{2} \right) - \frac{m_l^2 y}{4E_\nu M_N} \right) F_3 + \frac{m_l^2 (m_l^2 + Q^2)}{4E_\nu^2 M_N^2 x} F_4 - \frac{m_l^2}{E_\nu M_N} F_5 \right]$$

Structure functions

- ▶ F_1 |
- ▶ F_2 | \longrightarrow More precise estimation from other experiments
- ▶ F_3 | \longrightarrow Opposite sign for ν and **anti- ν**
- ▶ F_4 |
- ▶ F_5 | \longrightarrow Dependent on the lepton mass. Suppressed in case of ν_μ interactions, becomes relevant for ν_τ interactions

- ▶ Evaluation of F_3
- ▶ First evaluation of F_4 and F_5 , not accessible with lighter neutrinos

Some tau neutrino numbers

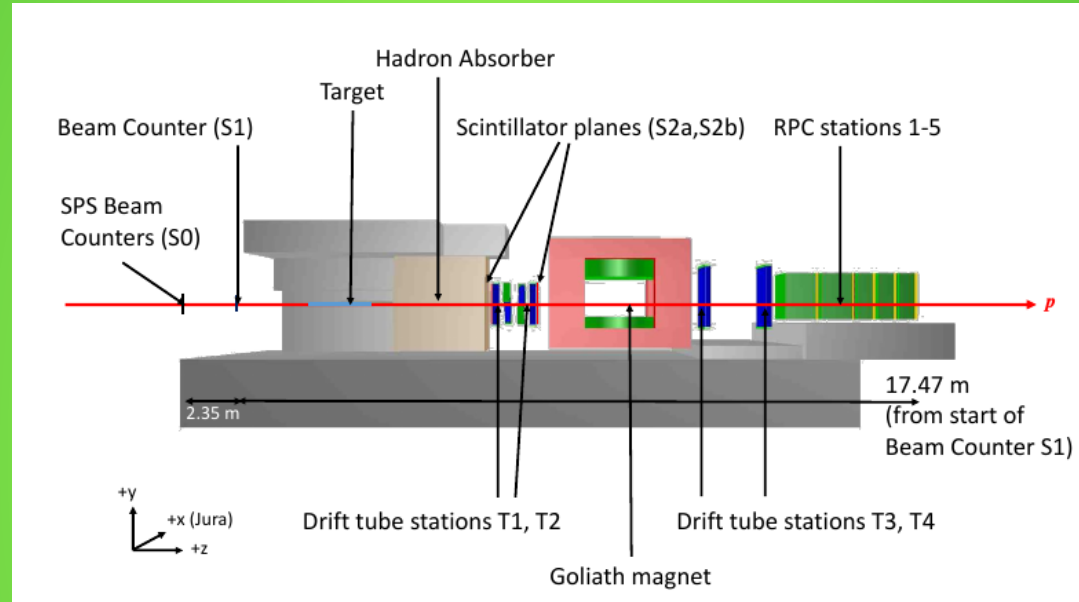
Current status of tau neutrino observations:

- DONUT observed 9 events (from charm) with a background of 1.5
- OPERA observed 4 events (from oscillations)
- No tau antineutrino has been even observed
- Ship can increase by 200 the current tau neutrino sample, and discover tau antineutrinos
- Measurement of tau neutrino differential cross-section in CC interactions
- Measurement of charm production for muon neutrinos and antineutrino (factor of 100 increase wrt CHORUS)
- A good fraction of the old OPERA collaborators are joining SHiP to build the neutrino sub-detector and analyse its data.

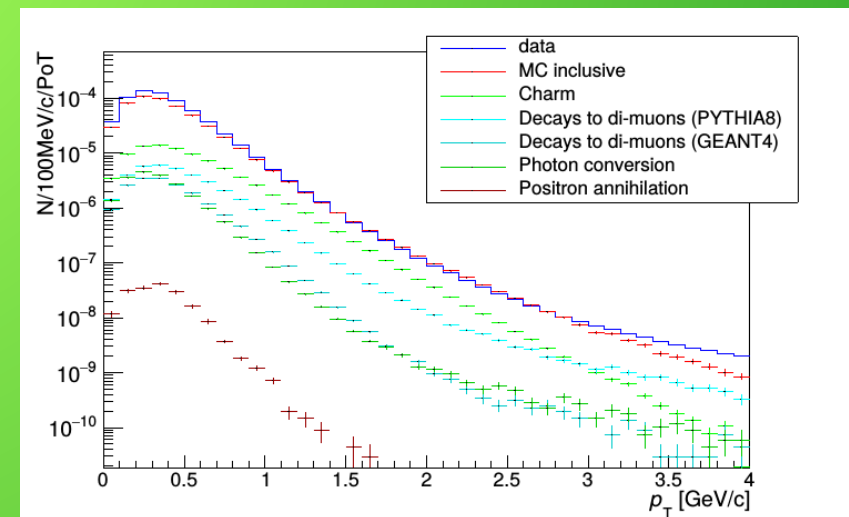
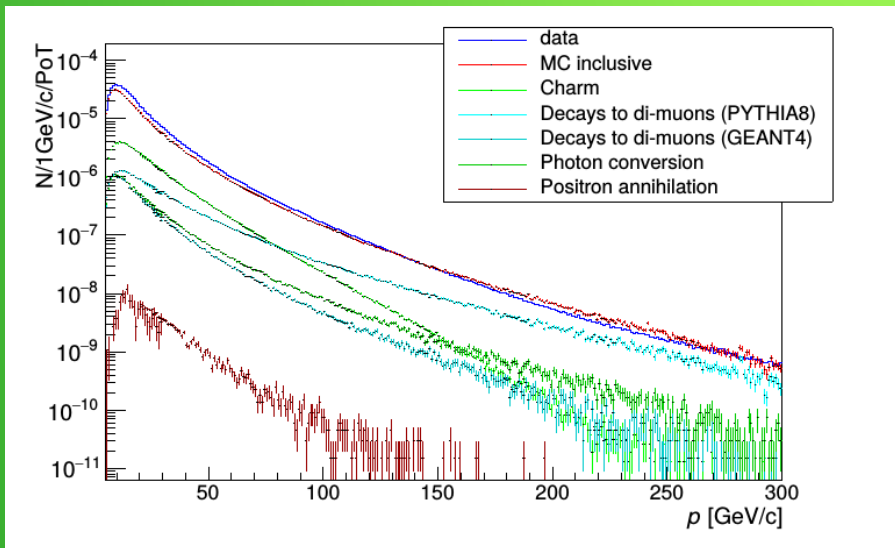
Muon flux measurement in 2018

- To validate simulations for the fundamental muon BG, a prototype target and hadron absorber have been exposed to the SPS beam

JHEP 2001.04784



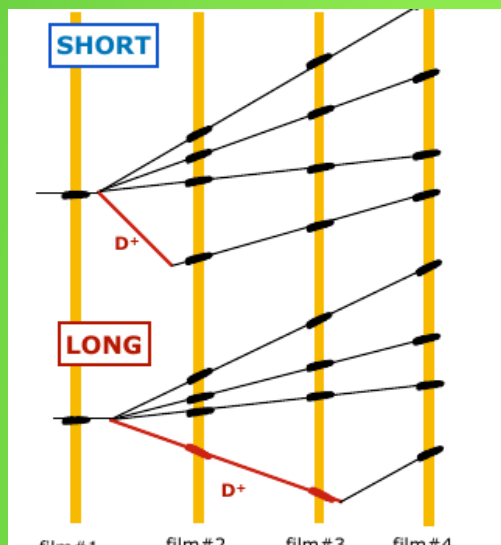
Reasonable agreement with simulation, but tails not well modeled (on 1% of SHiP spill)



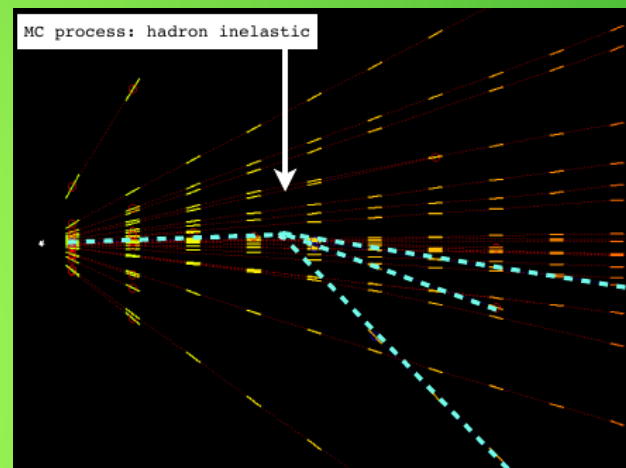
Double Charm production (preliminary)

- Used Emulsion Cloud Chambers to identify charm decay topology
- Pixel + SciFi + drift tubes to measure momentum, RPCs to identify muons

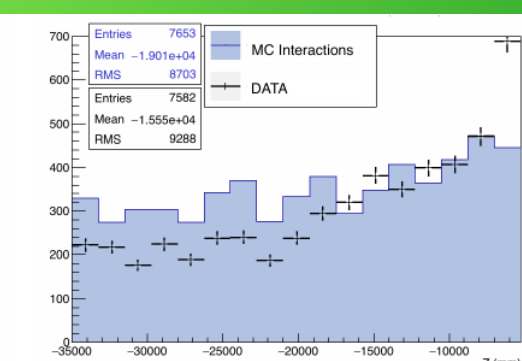
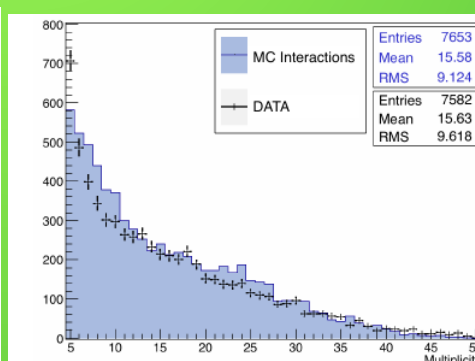
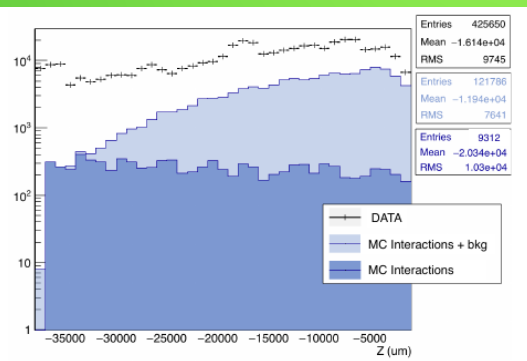
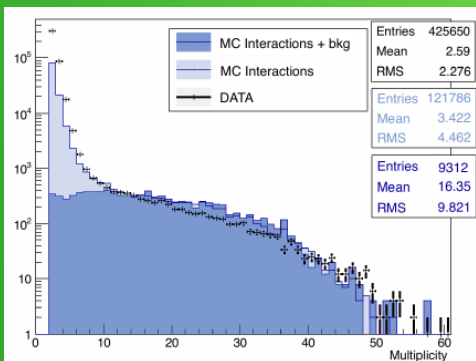
Charm interactions



BG

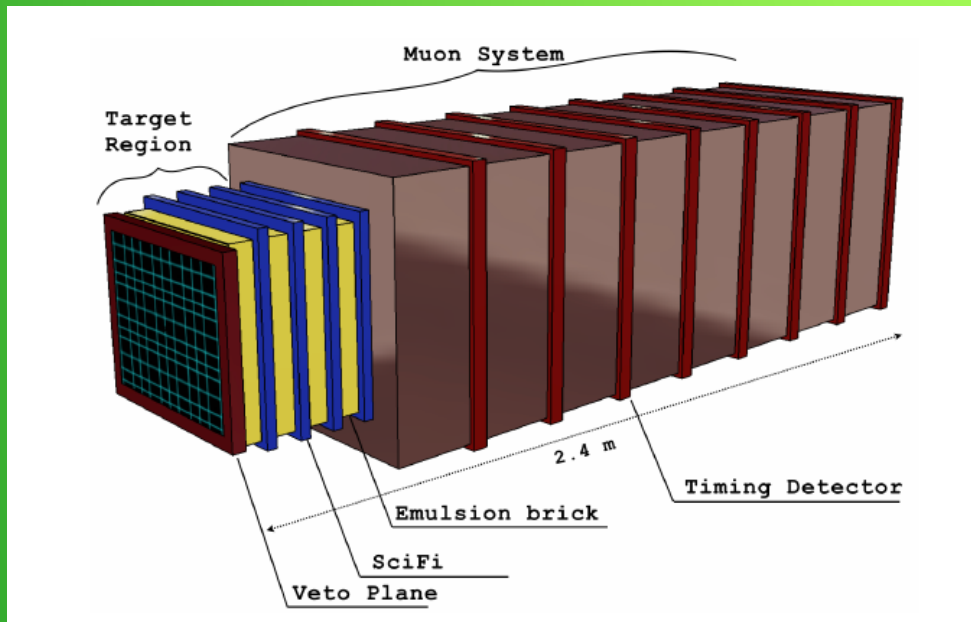
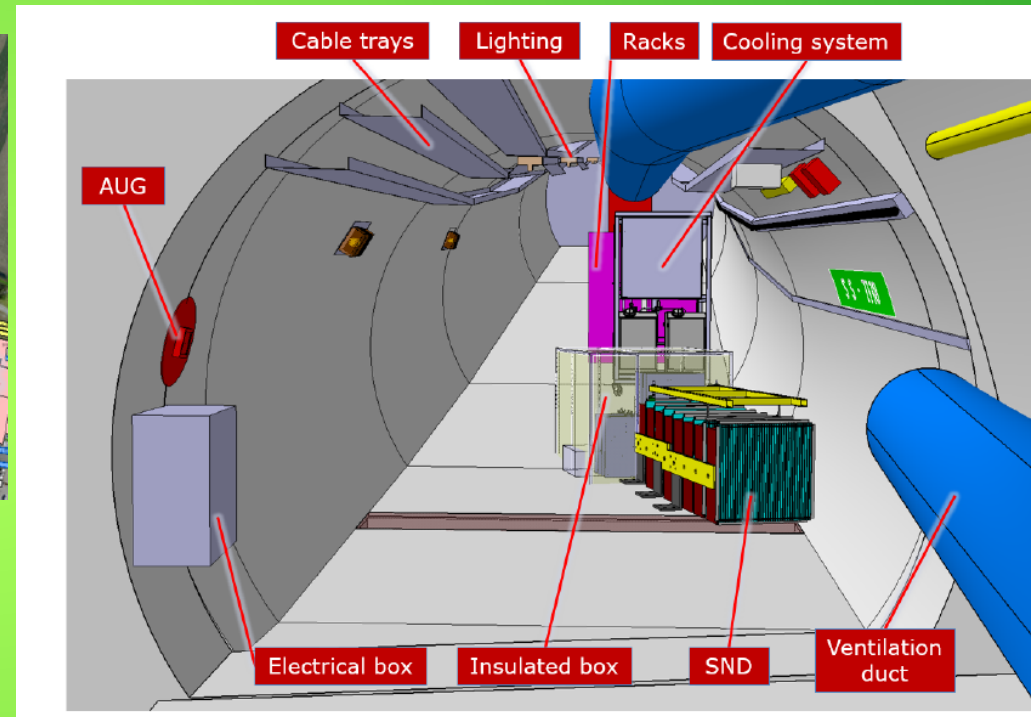


Multivariate techniques used to suppress background in vertex identification; charm analysis in emulsions still ongoing.



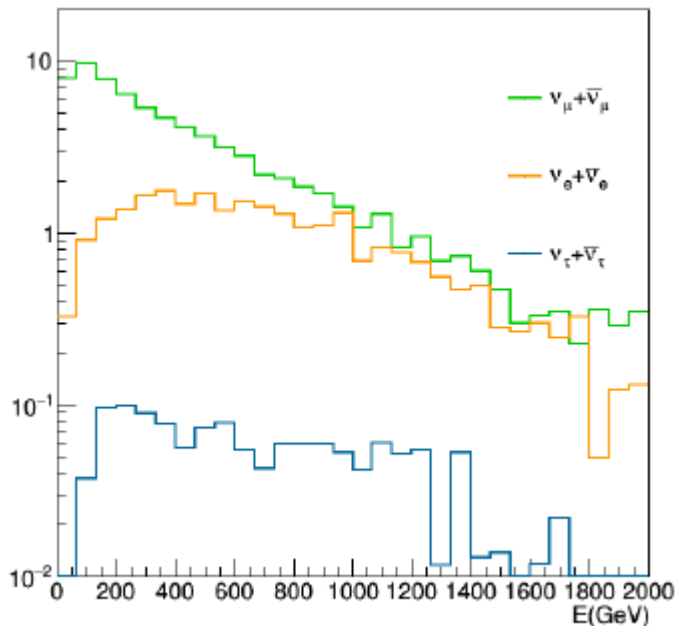
SND@LHC: arXiv 2002.08722

- A Scattering and Neutrino Detector to measure $pp \rightarrow \nu X$ at the LHC, to search for feebly interactive particles in an unexplored domain, using a prototype of the SHiP neutrino system in a LHC service tunnel covering $7.2 < \eta < 8.7$.

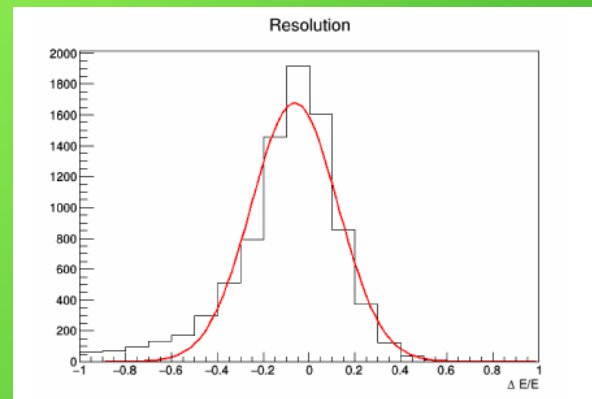


SND@LHC Neutrino physics

Neutrino flavour	$\langle E \rangle$ GeV (incident)	Neutrino Flux	$\langle E \rangle$ GeV (interacting)	CC Interactions Initial config	CC Interactions Updated config
ν_μ	150	4.6×10^{11}	460	62	975
ν_e	390	5.9×10^{10}	710	21	332
ν_τ	420	3.0×10^9	720	1	18
$\bar{\nu}_\mu$	150	4.0×10^{11}	480	27	429
$\bar{\nu}_e$	390	6.2×10^{10}	740	11	174
$\bar{\nu}_\tau$	360	2.9×10^9	720	0	7
TOT		9.87×10^{11}		122	1935

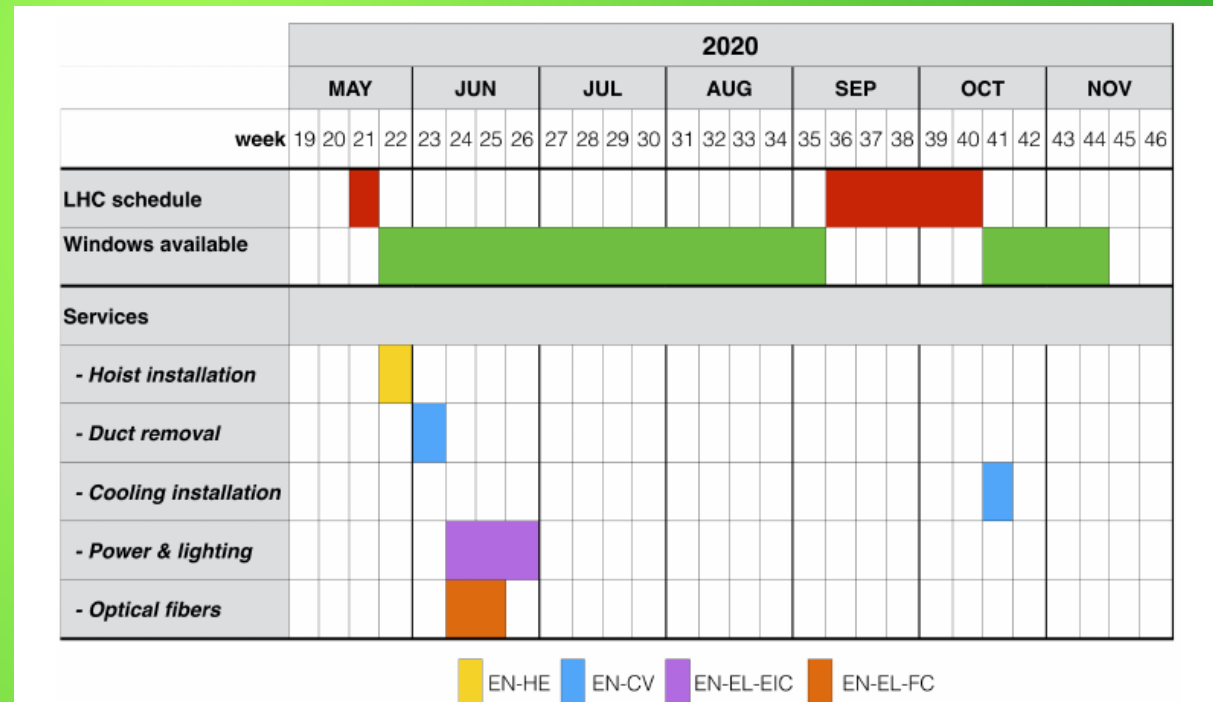
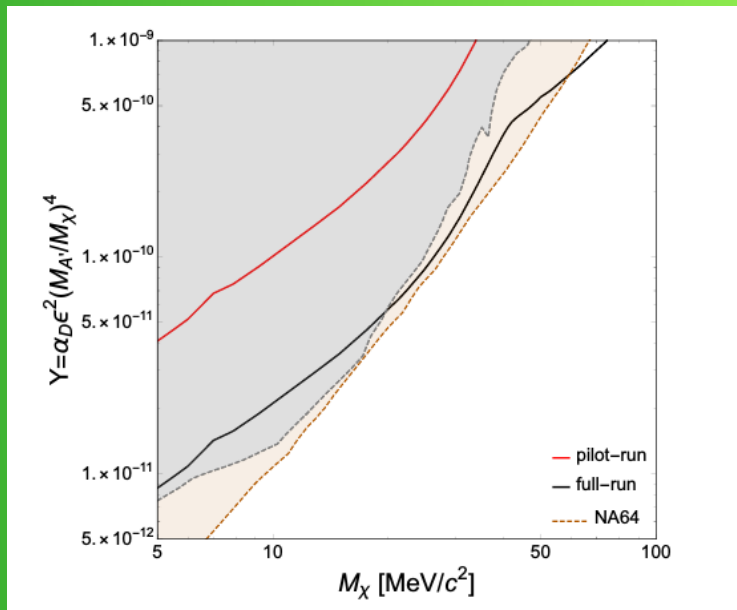


Tau identification efficiency about 50%, with BG (about 3 events) coming from charmed hadrons produced by other neutrino species. Neutrino energies can be reconstructed with 20% resolution.



Light dark matter: Dark Photon

- Dark photons of ~ 1 GeV mass could be produced by meson decays or photon bremsstrahlung
- Decay mode into a pair of LDM candidates $A' \rightarrow \chi\chi'$ followed by scattering in the emulsion target $\chi e^- \rightarrow \chi e^-$



Limits in a 0-BG scenario, with $m(A') = 3 m(\chi)$ and coupling $\alpha(D) = 0.1$

Very aggressive schedule, with installation of services already this summer!

Conclusions

- LHC Run 2 results gave no positive evidence for new physics
- We need an alternative approach to a next big brute-force, general-purpose high-energy collider
- Particle physics could reinvent itself in becoming smaller and smarter, designing experiments that target specific problems (dark matter, neutrinos, etc.)
- A detector like ShiP perfectly fits this philosophy
- Very positive feedback so far from CERN, we have just submitted a CDR and were a major player in the Physics Without Colliders initiative as well as in the European Strategy.
- Waiting for experiment approval, already interesting results from muon flux measurement, charm production, and tau neutrino physics from the proposed SND installation at the LHC