

# ATLAS SCT End-cap

#### Stephen Haywood

#### **Rutherford Appleton Lab**



# This Talk

- ATLAS
- Inner Detector
- SCT
- End-cap
  - Silicon Modules
  - Disks
  - Support Structures & Thermal Enclosures
  - Assembling the End-caps
  - Integration at CERN
  - Status
- Tracking in the Inner Detector
- Status
- Conclusions

Focus is Engineering for the SCT End-cap – JINST 3 P05002 (2008)

## ATLAS

Goals

- Understand the "mass mechanism": Higgs, Technicolour ...
- Investigate physics beyond the Standard Model: SUSY, Extra Dimensions (Black Holes), Additional Symmatries, etc.
- Investigate the Standard Model at 14 TeV: QCD, etc.
- Improve measurements of the Standard Model parameters: M<sub>W</sub>, m<sub>top</sub>, B-sector, etc.

Measurements at LHC:

- √s = 14 TeV
- Design Lumi =  $10^{34}$  cm<sup>-2</sup>s<sup>-1</sup>

## **ATLAS Detector**



## Inner Detector (ID)





## TRT

Straw Tracker – continuous tracking Transition Radiation detected by Xe – distinguish electrons and pions

#### Barrel:

- Effectively 36 layers of straws
- Embedded in "mats" of polypropylene fibres

#### End-cap:

 Stacks of 16 15 μm polypropylene foils, each separated by 200 μm

Total num straws = 400,000







#### **Pixels**

- 50 μm × 400 μm Pixels
- Bump-bonded chips
- 1744 Modules
- 82M channels

#### Barrel:

3 barrels at R = 5, 9, 12 cm

#### End-cap:

■ 2 × 3 disks





## SemiConductor Tracker (SCT)

#### Barrel:

- 4 cylinders
- 2112 Modules

#### End-cap:

- 2 × 9 disks
- 1976 Modules

#### Typical Module:

- 2 × 6 cm × 6 cm axial strips
- 2 × 6 cm × 6 cm stereo strips (40 mrad)
- Strips ~80 μm wide
- 6M channels



# SCT End-cap



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### Requirements

- Provide 4 space-points within  $|\eta| < 2.5$
- Modules placed on Disks to 70 (Inner) or 220 (Outer) μm
- Disks placed in Cylinder to 100 (x-y) and 1000 (z) μm
- Aligned to O(1)  $\mu$ m; stable to O(1)  $\mu$ m/hour
- Modules kept at -7°C each End-cap generates 10 kW heat
- End-cap to be kept dry; dew-point O(-30)°C
- Withstand hadron fluences of 2×10<sup>14</sup> cm<sup>-2</sup> 1 MeV neutron equiv
- Minimise magnetic materials (Fe,Ni)
- Minimise potential activation (Ag)
- Minimise electrical noise pick-up from ext sources and emission
- Comply with fire-safety requirements
- Reduce mass (radiation & interaction lengths)
- Tolerate Solenoid quench

## **End-cap Modules**





- 6 × 128 channels on each side
- Themal pyrolytic graphite (TPG) spine provides rigidity & cooling path
- Cooling at hybrid and "second point" (opposite end)
- Build precision: O(10) μm; 5 μm in most important params; measured to O(2) μm
- Expected measurement precision: 17 μm × 580 μm confirmed in Test Beam

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- Modules are "key-stone" phi-strips are radial
- 3 different radii: Outer, Middle, Inner (shorter)
- Disk 8 has "Short Middles"
- Total of 4 different types
- Stereo alternates orientation (same in Barrel): uφ, φv, uφ, φv, …
- Achieved by rotating Modules by ±20 mrad





- Support Modules
- Support Module Services

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#### **CFRP** Disks





- 180 μm CFRP facesheets: 3 plies at 0, ±60°
- 8.3 mm Korex® honeycombe core Korex: aramid fibres with phenolic resin; low moisture absorption
- 1<sup>st</sup> natural frequency: 22 Hz
- Out of plane distortions expected to be less than 40 μm

#### Services on Front of Disk



#### Services on Rear of Disk



## **On-Disk Cooling Circuits**

- Modules generate upto 10 W  $\rightarrow$  10 kW per End-cap
- Must be kept at -7°C to reduce radiation damage to silicon
- Stable temperature essential to reduce thermal motion
- Use evaporative cooling (latent heat): C<sub>3</sub>F<sub>8</sub>

- Tried Al pipes corrosion problems
- Use CuNi (70:30) 3.7 mm OD, 70 μm wall-thickness Good corrosion properties; easy to solder
- "Wiggly" design for stress relief
- Difficult to bend with bend radius of 4×diameter
- Watch holes in wall (from inclusions)  $\rightarrow$  careful QA

- Modules bolted to Pin on Cooling Blocks
- Cooling Blocks made of Carbon-Carbon: 100 W/m/K in good direction
- PEEK insulation between detector and hybrid portion of Block
- Gold-plated to avoid grease absorption







### **Power Tapes**

Supply

- LV digital & analogue power for detector
- HV for detector
- Power for Opto-electronics
- Control lines



- LV power (higher current): copper-clad aluminium twisted pair
- Rest: Cu traces on Polyimide tape (Aluminium too fragile)
- Due to complex design (modularity) 21 flavours of tape required

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#### **OptoHarnesses**

Optical fibres for

Data from Modules

Timing/Trigger/Control info to Modules
250 µm fibres clad in 0.9 mm OD furcation tubing
Contained in 12-way ribbons for upto 6 Modules



## FSI

#### Frequency Scanning Interferometry provides real-time alignment info

(Interfere light from measured length with light from reference length; scanning frequency allows absolute determination of length)

Precision O(1)  $\mu$ m in length Installed only in SCT



Measures movement ... due to thermal & humidity effects, gravitational sag, etc

Delicate emitter/receiver fibres in holders & reflectors on Disks





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Front

# Took 2 years to assemble 9 UK Disks



Rear

## **Module Mounting**

- By hand, with tooling (Barrel used robot)
- Thermal grease applied in controlled amount to Cooling Blocks
- Modules held to Block by washer & nut



## Testing

#### Extensive testing of

- Disks with Services
- Modules
- Modules on Disks

#### Metrology:

- Require Cooling Block Pins' position to 37, 60, 190 µm for Inner, Middle, Outer Modules (for sufficient overlap)
- Measure with CMM to 10 μm
- Global rotations, but Pin-Pin position in spec for all but one Pin



- 8000 "problematic" strips 0.26% of total, cf spec of 1%
- Mean of 4 out 1536 strips per module
- 80% of these are "dead"; 20% noisy or unbonded

#### **Finished Disks**





## **Support Structures**



- CFRP composites similar design to Disks: Faceskins & Korex honeycomb
- Cost 2/3 M\$ and consumed several years

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## Support Cylinder



- 9 mm thick
- Inserts accurate to 250 μm to position Disks

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#### Front & Rear Supports



Front: 9 mm thick



Rear: 25 mm thick

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### Support





Front & Rear Supports rest on **TRT** rails

Supported kinematically by "Mechanisms"



Front support



ATLAS SCT End-cap 31

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#### FEA & Tests



37,000 element model

#### Effect of gravity & CTE

- CTE  $1.4 \times 10^{-6}$  /°C ... 30 °C over 2 m  $\rightarrow$  80 µm CME =  $1.0 \times 10^{-4}$
- 1<sup>st</sup> mode 6 Hz; 2<sup>nd</sup> mode 24 Hz

∆humidity=50% @RT

- Taking a conservative vibration spectrum, expect deviations of 3 (40) μm perpendicular (parallel) to axis
- Test sample panels to > 2.5 MPa
- Load structure to ×1.5 working load; measure deflections of 0.74 and 0.87 mm, cf predictions of 0.63 mm

## **Thermal Enclosures**

- Contain SCT environmental gas (N<sub>2</sub>) ... external gas is CO<sub>2</sub>
- Moisture barrier
- Thermal barrier ... TRT is at 22.5 °C
- Prevent formation of condensation/ice on outside of SCT
- Faraday shield



#### **Outer Thermal Enclosure**



- Sandwich: aluminised polyimide / 8 mm foam / Cu-polyimide
- Use Araldite 2011
- Much prototyping

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#### **Inner Thermal Enclosure**



- CFRP laminate cylinder / 5 mm foam / Cu-polyimide
- Includes gas channels with 0.3 mm holes for gas purge

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### **Membranes**



- Gas seals
- Complete Faraday shield
- Connect all Cu-Kapton foils with solder



## Heater Pads

- Critical component: To ensure keep outside above dew-point and in case moist air gets into Inner Detector, cover outside with heater pads
- 8 μm thick Cu tracks sandwiched between polyimide
- 150 W/m<sup>2</sup> or 300 W/m<sup>2</sup>
- Double tracks for redundancy
- Integral thermocouples
- Switch power (rise/fall time O(1) ms)





Above ignores Module heating / cooling ~ O(10) kW per End-cap Net inward flux ~ 200 W – so small perturbation for Module cooling

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# Assembling the End-caps

#### Took place in

- Liverpool
- Nikhef



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## **Disk Insertion**



# Disk "grabbed" at inner radius



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- Located with microscopes longitudinally to 200 μm
- Located with telescopes transversely to ~100 μm
- Cylinder pre-loaded to compensate for subsequent additions of Disks & Services
- Each Disk held by 12 pins around circumference

## **Services**



+ Optical Fibres

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- Significant heat load
- In absence of cooling, expect temp rise of 50°C at worst
- Wrap in 150 μm AI foil, including dedicated (spare) cooling pipes
- Compress with Cu-Be spring

## **Transportation from Liverpool**



- Temperature & humidity controlled, air-sprung lorry
- Serious test run with dummy load
- Carefully monitored
- Insured for 9 MSF

# Integration at CERN

- Lots of checks and re-laying of Services at CERN
- Add Rear Support and mount on cantilever beam





## Add Thermal Enclosures



## Insert into TRT



- TRT on rails and slid over the SCT on cantilever beam
- Add Front Support
- Align on TRT Rails
- Seal Thermal Enclosure & dry out
- Electrical tests



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## Insert into ATLAS



- Use Contact Sensor to work out when in place (up to Barrel)
- Both End-caps make contact 5 mm before nominal one End-cap
  3 mm too long

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### Services & Patch Panels (PPF1)







Service lengths carefully calculated to avoid deficit/excess

Panels

• Cable trays added

front of TRT

Patch panels at end of Ecal Cryostat

SCT Services in ID Patch

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SCT

## **Cooling System**

- C<sub>3</sub>F<sub>8</sub> liquid enters SCT at room temp
- Leaves as vapour/liquid at around –20 °C
- Heat Exchanger to heat/cooling entering/leaving C<sub>3</sub>F<sub>8</sub>
- Heat Exchangers occupy space foreseen for cancelled TRT C-wheels
- Must boil off excess fluid, else will cause condensation on un-insulated pipes
- Heaters have been cause of many problems



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# Grounding & Shielding

- Can be a make-or-break factor
- Careful consideration: Module, Disk, End-cap, Services, Cable Trays, ID, ATLAS
- Solid connections (<< 0.2  $\Omega$ ) & insulation (> 1 M $\Omega$ )
- Avoid apertures > 1 cm × 10 cm where possible
- Use Alochrome 1200 & Fingerstock
- Make measurements before/after insertion into TRT & ATLAS









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# Timelines

**Inner Detector** 





Apr 97 ID TDR Nov 01 EC FDR

- Nov 01 Start Tendering process for Support Structures
- Sep 02 Work Starts on Support Structures
- Jul 04 Support Structures completed
- Sep 03 Disc Services PRR
- Jan 05 TE PRR



- Feb 06 EC-C Transported from Liverpool to CERN
- Jun 07 EC-A Inserted in ATLAS
- Feb 08 End-caps signed off; first Cosmic



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# Mass

- Target for error is 1% more critical at lower radii (tracking volume)
- Very careful bottom-up estimates of component masses
- Disk (without Modules) correct to 1.4%
- Mass supported by Front & Rear Supports is 178 kg, cf initial design estimate of 168 kg
- Difference between two End-caps (some +'s and -'s) is < 1 kg
- Attempt to weigh SCT (inside TRT) was inconclusive

Component	Mass (kg)
Modules	24
Disks	33
Support Cylinder, Services & OTE	57
Other Support Structures	23
Rest of Services	88
PPF1 Patch Panels	35
Total	259

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### **Radiation Lengths**



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# **RAL Contributions**

#### Barrel:

- 600 Modules
- On-Cylinder Cooling Circuits
- Services  $\rightarrow$  Cylinders
- Thermal Enclosure design & manufacture
- Mass

#### End-cap:

- On-Disk Cooling Circuits
- Services → 9 Disks
- Support Structures design & procurement
- Thermal Enclosure design
- UK Transportation & Insurance
- Off-Disk Cooling Circuits, Services routing & Patch Panels
- Mass

# **Tracking Performance**

![](_page_56_Picture_1.jpeg)

### **Track Parameter Resolutions**

![](_page_57_Figure_1.jpeg)

### **Reconstruction Efficiency**

![](_page_58_Figure_1.jpeg)

5 GeV  $\mu$ ,  $\pi$ , e

![](_page_58_Figure_3.jpeg)

Tracks in b-jet from ttbar events, as function of distance from core of jet

![](_page_58_Figure_5.jpeg)

### **Vertex Reconstruction**

![](_page_59_Figure_1.jpeg)

### **Electrons & Conversion Photons**

![](_page_60_Figure_1.jpeg)

 $J/\psi$  mass resolution in Barrel (left) & End-cap (right)

![](_page_60_Figure_3.jpeg)

#### Corresponding plots for Muons

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### **Electrons & Conversion Photons**

![](_page_61_Figure_1.jpeg)

 $J/\psi$  mass resolution in Barrel (left) & End-cap (right)

#### Pion rejection in TRT

Conversion identification

![](_page_61_Figure_5.jpeg)

![](_page_61_Figure_6.jpeg)

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**Status** 

![](_page_62_Picture_1.jpeg)

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- ATLAS was ready for Collisions on Sunday 21 Sep, but ... ⊗
- I am not aware of any serious problems with any of subdetectors
  Inner Detector

TRT

Some dead HV cards

**Pixels** 

- 11 (4) dead (problematic) modules
- 36 modules unusable on Disks due to problems with 3 cooling loops
- Currently, can operate 95% of Pixels hope to recover even more SCT
- Leak rates of N<sub>2</sub> exceed spec, but are tolerable: dry-out achieved, so operate with lower overpressure (Barrel SCT has large leak rate)
- One Module Cooling Circuit on Disk 9 has large leak of C<sub>3</sub>F<sub>8</sub> and is inoperable loss of 13 Modules ... not terrible
- One Module Cooling Circuit on Disk 1 has Heater problem and is currently inoperable – loss of 23 Modules ... not great – now fixed

# SCT Summary

	Barrel	End-cap A	End-cap C
Total Num Modules	2112	988	988
Modules not functional	3	0	1
Modules not cooled (2008)	0	0	36→13
Dead Strips (%)	0.2	0.3	0.3
Chips lost	13	0	0
Functional channels (%)	99.6	99.7	96.0→98.3

## **Inner Detector Commissioning**

Huge amount of testing, more recently with:

- Cosmics very useful for Alignment
- Beam "splash" very useful for timing: "Unique opportunity to time whole the detector at once in one event! This saves may be months of work."
- Beam-gas nice for Software Reconstruction, but not many events

![](_page_65_Figure_5.jpeg)

## Cosmics

![](_page_66_Figure_1.jpeg)

TRT, SCT & Pix Barrels

![](_page_66_Figure_3.jpeg)

### **Cosmics: Measurements**

![](_page_67_Figure_1.jpeg)

## Single Beam

![](_page_68_Figure_1.jpeg)

# Conclusions

- Two ATLAS SCT End-caps have been constructed, meeting almost all of the specs.
- The project has taken ~15 years, with ~6 years required for Construction and Assembly.
- ~200 people have worked on the End-caps.
- The insured value of the hardware was 9 MCHF for each End-cap.
- Apart from one two Cooling Circuit problems, the End-caps are close to fully functional.

![](_page_69_Picture_6.jpeg)

The ATLAS Inner Detector is ready to receive LHC collisions and the Software is in place to reconstruct the First Data.

![](_page_69_Picture_8.jpeg)

# Acknowledgments

- Work of many institutes: Modules, Services
- Engineering especially: Liverpool, Nikhef, RAL, CERN
- Underlying Paper benefitted from input from:

![](_page_70_Picture_4.jpeg)

Pepe Bernabeu Pawel Bruckman Craig Buttar Sandra Ciocio Paul Dervan Didier Ferrere Peter Ford Martin Gibson Jennifer Haywood Nigel Hessey Tim Jones Maaike Limper Caroline Magrath Chris Nelson Brian Smith Luis Sospedra Jason Tarrant Tony Weidberg Pippa Wells Patrick Werneke Ian Wilmut

![](_page_70_Picture_7.jpeg)

# **RAL Contributors**

#### <u>PPD</u>

Jeff Bizzell Stephen Haywood Richard Holt Bruce Gallop Martin Gibson Peter Phillips Mike Tyndel

<u>ED - PED</u> Paul Barclay Steve Butterworth Peter Ford Debbie Greenfield Tony Jones Chris Nelson John Noviss Brian Smith Jason Tarrant Ian Wilmut

![](_page_71_Picture_4.jpeg)

<u>ID</u> Craig McWaters John Matheson

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