The FASER experiment



The Standard Model and its missing bits

- The Standard Model of particle physics is certainly a successful effective theory that describes well our (ewk-scale) world
 - The discovery of the Higgs boson and the precision measurements performed so far have cast a milestone. E.g.
 - Couplings with quarks and leptons
 - (very recently) Evidence of off-shell Higgs production and measurement of total Higgs width $\Gamma_{\rm H} = 4.6 + 2.6 2.5$ MeV
- However, we know that the SM is incomplete
 - No gravity embedded in it
 - No unification of forces at high energies
 - No explanation on Higgs (and other particles) mass(es)
 - No explanation on why we are here as 'matter' (matter-antimatter asymmetry) dark energy
 - No explanation of the Universe mass \rightarrow dark matter!
- The nature of Dark Matter (DM) is unknown, so understanding it it is one of the biggest challenges of particle physics nowadays





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The quest for a hidden, dark sector

- DM could be just one of the many new particles belonging to a 'hidden' dark sector (DS)
- The mechanism of portals as the lowest canonicaldimension operators that mix new dark-sector states with gauge-invariant combinations of SM fields is often considered, with 4 notable examples (benchmarks):

Portal	Coupling
Vector (Dark Photon, A_{μ})	$-\frac{\varepsilon}{2\cos\theta_W}F'_{\mu\nu}B^{\mu\nu}$
Scalar (Dark Higgs, S)	$(\mu S + \lambda_{HS}S^2)H^{\dagger}H$
Fermion (Sterile Neutrino, <i>N</i>)	<i>y_NLHN</i>
Pseudo-scalar (Axion, a)	$\frac{a}{f_a}F_{\mu\nu}\tilde{F}^{\mu\nu}, \frac{a}{f_a}G_{i,\mu\nu}\tilde{G}_i^{\mu\nu}, \frac{\partial_{\mu}a}{f_a}\overline{\psi}\gamma^{\mu}\gamma^{5}\psi$

- The resulting new particles could be Long-Lived (LLP)
 - Lifetime related to SM-DS mixing coupling $\boldsymbol{\epsilon}$
- Targeted searches for these BSM models have been identified as a priority by the European Strategy Updated and the Snowmass process



Portal

solid lines show were we are...

The LHC (main) experiments

- Dark sectors are certainly part of the LHC experiments programmes
- The Run 3 has just started, with exciting opportunities offered by the upgrade of the accelerator, leading to an increased centre-of-mass energy (13.6 TeV), renewed detectors and novel triggers



LHCb's Vertex Locator (VELO)

ALICE time-projection chamber





ATLAS New Small Wheel

FASER and the DS: Motivation

- So, why dedicated experiments? → might complement LHC main experiments in terms of targeted phase space and mitigate issues notably, large background rates and difficulties in triggering
- Idea of FASER: a Forward Detector for low-mass LLPs
 - Background mitigated by rock/shielding

Gwilliam

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Credits to



FASER and the DS: Motivation

- So, why dedicated experiments? → might complement LHC main experiments in terms of targeted phase space and mitigate issues notably, large background rates and difficulties in triggering
- Idea of FASER: a Forward Detector for low-mass LLPs
 - Background mitigated by rock/shielding
 - Simpler / no triggering needed







Detector is far from IP \rightarrow target **long lifetimes**

LLPs produced in forward-peaked light hadron decays \rightarrow e.g. O(10¹⁴) pions within FASER angular acceptance

$$\theta \simeq \tan \theta = \frac{p_T}{p} \sim \frac{m}{E} \ll 1$$

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Credits to C. Gwilliam

FASER location



- FASER is located at ~ 480 m downstream of the ATLAS interaction point (IP) in the TI12 - an unused SPS maintenance tunnel intersecting collision axis
- The beam is highly collimated (mrad diameter) \rightarrow only a small detector needed, with a magnet aperture of 20cm diameter

Insertion

D2 Q4

beam

150

L[m]

Arc

RI12

LLP

TI12

Infrastructure & rock catches most collision products



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beam collision axis

FASER

UJ12

10m

FASER

FASER Physics reach

Designed for events of the kind



- Probes large range of BSM models in regions favoured according to muon g-2, DM hypotheses and anomalies
 - Dark photons models:



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Dark photon production

A'

10⁻³

a LOO

FASER Physics reach (2)

a

Also sensitive to axion-like particles (ALPS), new particles that can mediate the interactions between the SM and the hidden sector by coupling to photons, gluons, W and Z bosons, and fermions. Still produced and decaying as:

pp \rightarrow LLP, LLP travels ~480m, LLP \rightarrow ee, $\gamma\gamma$, $\mu\mu$, ...



Other targeted models: Dark Higgses, Heavy Neutral leptons, etc.

FASER v Physics reach

- FASER is also sensitive to high-energy neutrinos produced along beamline !
 - A dedicated component, FASERv, has been added in 2020 to study an uncovered range of energies for v



Expected number of charged current neutrino interaction events (for 250 fb⁻¹) based on "F. Kling and L.J. Nevay, Forward Neutrino Fluxes at the LHC, Phys. Rev. D 104, 113008"

Gen	erators	FASER		SND@LHC			
light hadrons	heavy hadrons	$\nu_e + \bar{\nu}_e$	$ u_{\mu} + ar{ u}_{\mu} $	$\nu_{\tau} + \bar{\nu}_{\tau}$	$\nu_e + \bar{\nu}_e$	$ u_{\mu} + ar{ u}_{\mu}$	$ u_{ au} + ar{ u}_{ au} $
SIBYLL	SIBYLL	1501	7971	24.5	223	1316	12.6
DPMJET	DPMJET	5761	11813	161	658	1723	31
EPOSLHC	Pythia8 (Hard)	2521	9841	57	445	1871	19.2
QGSJET	Pythia8 (Soft)	1616	8918	26.8	308	1691	12
Combin	ation (all)	2850^{+2910}_{-1348}	9636^{+2176}_{-1663}	67.5^{+94}_{-43}	408^{+248}_{-185}	1651^{+220}_{-333}	$18.8^{+12}_{-6.6}$
Combination	(w/o DPMJET)	1880^{+641}_{-378}	8910^{+930}_{-938}	$36^{+20.8}_{-11.5}$	325^{+118}_{-101}	1626^{+243}_{-308}	$14.6^{+4.5}_{-2.5}$

Expected number of charged current neutrino interaction events occurring in FASERv and SND@LHC assuming 250 fb-1 Run 3 lumi.

Predictions from different MC generators differ substantially

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FASER v Physics reach

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Overview of the FASER detector



Angular acceptance $|\theta| < 0.21$ mrad region, $\eta > 9.2$ FASERv extends $|\theta|$ to 0.41 mrad and $\eta \sim 8.5$ FASER Trigger rate: **650 Hz** expected (*dominated by muons produced close to the IP*) 2021 JINST 16 P12028

Very low radiation levels

Detector paper: <u>https://arxiv.org/abs/2207.11427</u>

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Overview of the FASER detector

• 10 cm radius







Scintillator veto

1/22

Overview of the FASER detector

• 10 cm radius



FASER Tracking: components and layout

Composed of two distinct parts: the tracking spectrometer (3 tracking stations) and the Interface Tracker (1 tracking station), placed after the FASERv emulsion detector



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3 Tracker planes per station (12 total)

NIMA 166825 (2022)

Low material central region: 2.1% radiation length

Component	Material	Number	X ₀ (%)	
		/ station	Central region	Edge region
Silicon sensor	Si	6	1.8%	1.8%
Station Covers	CFRP	2	0.3%	0.3%
SCT module support	TPG	3	-	0.6%
C-C Hybrid	C (based)	3	-	2.2%
ABCD chips	Si	3	-	6.5%
Layer frame	Al	3	-	10.1%
Total / station	-	-	2.1%	21.5%

FASER Tracking: tests and performa

Extensive tests have been carried out to evaluate performance, standalone and after installation:

- long-term stability and control checks (temperature, humidity, electronics)
- Quantification of noisy/dead strips
- Alignment and metrology of tracker planes

Commissioning with cosmic rays and LHC pilot run



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NIMA 166825 (2022)

Injected charge [fC]

100



purple line: combined track fit to the hits in the tracking stations during 900 GeV pilot beam 22/11/22

FASER calorimeter, pre-shower and scintillator systems

- Four scintillator stations with multiple scintillator layers in each station
 - (a) FASER ν Veto, (b) Interface Veto, (c) Timing, & (d) Preshower
 - >99.98% efficiency, sufficient to veto all incoming muons
 - photo-multiplier tubes to detect the scintillation signals.

Note: Preshower scintillator to be replaced by silicon pixel detector (<u>tech.</u> <u>proposal</u>) in 2023/2024 To detect 2-photon axion-like particle signals

- Electromagnetic calorimeter made of spare LHCb modules
 - 66 layers of lead-scintillator plates read by 2x2 array of PMTs
 - calorimeter readout optimised to measure multi-TeV deposits w/o saturation



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(3x3 version)

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 w/o saturation

$$\frac{\sigma_E}{E} = \frac{9.2\%}{\sqrt{E}} \oplus 0.2\% + \text{expected 1\% constant term}$$

At 1 TeV, about 1.6% of electrons are expected to leak more than 3% of their energy



FASERv detector

- 700 layers of an emulsion film and 1.1 mm tungsten plate: 25 cm×30 cm×1.1 m, 1.1 tons, 220 X₀
- Pilot detector (30 kg) exposed in FASER location for 1 month
 - Observed (2.7σ) first collider v candidates!
- FASERv will be exchanged frequently during Run 3
 - First full detector (TS1): 26th July 13th Sept
 - Second detector (TS2): 13th Sept 8th Nov
- Frequently exchanged (~ every 3 months) to keep a manageable detector occupancy. Procedure:

Phys. ReV. D 1004, L091101







Trigger and DAQ

2021 JINST 16 P12028

FASER Trigger rate: 650 Hz expected (dominated by muons)

- PMTs from scintillators and calorimeter provide trigger signals
- Trigger system run synchronously to the 40.08 MHz LHC clock
- Data Acquisition (DAQ): Configuration & readout
- Monitoring: checking data flow, detector conditions, and data quality to spot/resolve problems

Data Control & Safety (DCS): powers detector and protects it from unusual conditions



FASER Trigger/DAQ Overview



12

14

16

10

Preshower ch7 (in MIPs) / Preshower ch6 (in MIPs)

18

10

 10^{2}

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2

10

0.2

Local X residual [mm]

¹⁰³Beam Energy [GeV] ∠∠/ I I / 22

FASERv detector commissioning

- ~30% of the full emulsion for commissioning
- MIP efficiency of the veto system was also measured in the test beam
- Better performance than the requirement (>99.98%) obtained.





Full detector installed

-

After further commissioning tests, the detector was ready to see first data on July 5th!

2t



FASERnu

rom

First data !

- Thousands of events were already collected with charged particle tracks traversing the detector even prior to official start on 5th of July
 - Great for performance studies, optimizing operation procedures, & commissioning reconstruction software.
- With 13.6 TeV beams, good events seen in the detector consistent with coming from collisions.



One of first event displays from collisions



Data-taking

- Instantaneous luminosity measured at IP1 and the FASER total trigger recorded rate for 8 LHC fills between August 19th to 23rd, 2022:
 - The trigger rate trend generally follows the luminosity trend but falls off more strongly at the beginning of fills → higher beam losses at the beginning of fills. Overall very good performance and small issues well understood



Recorded rate very similar to triggered rate, efficient data taking (< 2% physics deadtime) Monica DOnofrio, Seminar @ Birmingham

Rates and trigger efficiency



High rate at the beginning of the fill

Higher than expected trigger rate in general ~1.3 kHz at L=2×10³⁴ cm⁻²s⁻¹ Dominated by IP1 muon flux, not problematic for physics

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FASER has 4 physics trigger items (partially overlapping, not prescaled):

- calorimeter (lowest rate, ~ 50 Hz)
- Veto-scintillator or preshower scintillator
- Time scintillator station (highest rate, ~ 800 Hz here)
- Coincidence trigger

Trigger efficiency of the calorimeter



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Run 3 events

FJSER

Run 8336

Event 1477982

Collision event with a muon traversing FASER. The measured track momentum is 21.9 GeV. The waveforms are shown for signals in scintillator counters and calorimeter modules and are fitted using a Crystal Ball function.

2022-08-23 01:46:15 To ATLAS IP Tracking spectrometer Trigger Magnets Interface Veto FASERv stations Pre-shower station FASERv Decay volume Tracker (IFT) station veto station station emulsion detector Calorimeter AL FASERv veto station, layer 1 Trigger station, top layer, PMT left Trigger station, top layer, PMT right Veto station, layer 2 Veto station, layer 3 Pre-shower station, layer 1 Calorimeter, top row, left module Calorimeter, top row, right modul Mean: 827.3 ns Peak: 187.0 mV Mean: 828.0 ns Mean: 848.8 ns Peak: 17.5 mV Mean: 796.1 ns Peak: 378.9 mV Integral: 186.4 pC £ 500 J 350 300 250 E 50 Peak: 272.5 mV 2.5 0 15.0 ę Integral: 71.9 pC Integral: 92.2 pC 용 100 Integral: 6.8 p itude 12.5 ₹ 10.0 200 150 100 775 825 850 875 900 825 875 825 850 875 900 825 850 875 900 775 800 825 850 875 900 775 850 875 900 850 875 900 875 900 Time [ns] S 120 Trigger station, bottom layer, PMT left Trigger station, bottom layer, PMT right FASERv veto station, layer 2 Veto station, layer 4 Pre-shower station, layer 2 Calorimeter, bottom row, left module Calorimeter, bottom row, right module 400 350 300 250 ≥ 17.5 Hean: 850.0 ns Peak: 13.5 mV ⊕ 15.0 - Integral: 4.4 pC ž Mean: 828.6 ns Peak: 525.2 mV ntegral: 185.8 pC Mean: 828.1 ns Peak: 100.3 mV Integral: 27.6 pC Mean: 829.0 ns Mean: 797.8 ns Mean: 825.9 ns E 500 Peak: 346.2 mV Integral: 184.7 pC Peak: 96.0 mV Integral: 31.3 pC Peak: 2.3 mV Integral: 0.7 pC 400 12.5 10.0 300 200 150 200 100 875 375 900 Time [ns] 825 875 900 875 900 800 825 875 90 800 825 850 875 900 800 800 Time Ins Time [ns] Time (ns Time [ns] Time [ns] Time (ns)

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More event displays showing tracking system



Run 8336 Event 1477982 2022-08-23 01:46:15



- event triggered by modules in the scintillator systems
- magnets are shown as light blue cylinders
- tracking stations as grey cuboids
- In each tracking station, the silicon sensors facing forward are shown in dark blue and the readout electronics in orange.
- The detected hits in the semiconductor tracker modules are shown with blue lines
- The reconstructed track is shown with a red line.

More from the Run 3

- Reconstructed tracks (above ~1 GeV) in 1 mm ~ 1 mm ~ 20 emulsion films from the 2022 first module of the FASERv detector, which collected 0.5 fb⁻¹ of data.
- The yellow line segments show the trajectories of charged particles in the emulsion films, others are interpolations, with colours changing depending on the depth in the detector.



Tracks from the first FASERv emulsion films

The track density is around $2.3 \sim 10^4$ /cm²/fb⁻¹.

Distributions of position and angular distributions

Distributions of the position deviation between the track hits and the straight-line fits to reconstructed tracks. measured in the 2022 first module of FASERnu

00001 N tracks 0006 N 1000

7000

6000



Data analysis readiness

- On-going tests on full production chain from generation all the way through to analysis
- Representative background and signal processes have been produced
 - Full FASER detector geometry implemented and validated in offline software
 - Calypso software package based on ATLAS framework (Gaudi and Athena)
 - Genie & FLUKA used for neutrinos studies and muon-induced background





Good tests for track reconstruction methods, momentum resolution and calorimeter deposits measurements

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Neutrino events simulation also fully ready

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The forward future: FASER(v)2

- We must think ahead!
- It could be that we do not see LLPs or NP in Run 3:
 - Extended coverage needs a bigger detector!
- Envisaging a scaled-up version of FASER with ~100 x active area
 - Magnets: Superconducting w/ B = 1 T
 - Tracker: much larger using e.g. SiFI/SiPM
 - Calo/Muon: enhanced PID & position resol.



	FASER	FASER2
R [m]	0.1	1
DV [m}	1.5	10
TS [m]	2.6	10

FASERv2: Focus on v_{τ}



Latest updates: https://indico.cern.ch/event/1196506/

The forward future: FASER(v)2

Substantial increase in sensitivity for LLPs from B, D hadrons decays (e.g. Dark Higgs) thanks to larger radius, broader scope including QCD physics

Benchmark Model	FASER	FASER 2
Dark Photons	\checkmark	√
B - L Gauge Bosons	V	V I
$L_i - L_j$ Gauge Bosons		<u> </u>
Dark Higgs Bosons		√
Dark Higgs Bosons with hSS	—	\checkmark
HNLs with e	—	\checkmark
HNLs with μ	—	√
HNLs with τ	\checkmark	\checkmark
ALPs with Photon	√	√
ALPs with Fermion	_	✓
ALPs with Gluon	\checkmark	\checkmark
Dark Pseudoscalars		√

Systematic studies on-going to establish the layout compromising costs and new physics sensitivity



- FASER2

100

ALP mass m_a [GeV]

 10^{0}

 ${}^{10^{-2}}_{10^{-1}}$

₩ 10-1

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 10^{-1}

|^{דא}ט| פר

.≍ ∑ 10⁻³

10-4

100

 $10^{-2} \underset{10^{-1}}{\downarrow}$

₩ 10-1

FASER2

100

HNL mass m_N [GeV]

10

 10^{-2}

 10^{-1}

₩ 10-1

101

10

SHIP

100

Dark Higgs Mass m_{ϕ} [GeV]

Where? The Forward Physics Facility (FPF)

- Proposal to build a new dedicated forward physics facility
 - Hosting a suite of far-forward experiments at the HL-LHC



Current planned detectors

- FASER2
 - FASER scaled to r=1m
 - Light dark sector parts.
- FASERv2
 - ~20t emulsion + tungsten detector
 - Mainly v_{τ}
- AdvSND
 - Off axis v detector
 - Fwd charm + low-x gluon
- FORMOSA
 - Scintillating bars
 - Millicharged particles
- FLArE
 - ~10t LAr TPC
 - DM + v physics

Detailed (429pp) paper submitted as part of Snowmass: <u>https://arxiv.org/abs/2203.05090</u> Monica DOnofrio, Seminar @ Birmingham

Summary

- FASER gives access to light, weakly-interacting particles with significant lifetime, providing sensitivity to a wide range of BSM physics models (dark γ, ALPS and more) complementary to GPDs; FASERv can measure high energy neutrinos in a previously unconstrained region of phase space
- FASER and FASER ν are now fully operational and taking data!
 - Test beam results show excellent tracker cluster efficiency and uniform calorimeter response within a few percent across different beam positions
 - Data collection has started with Run 3! More than 35/fb of data collected so far...
 - detector working beautifully, no real operational issues to date
- Development of analysis and software tools ongoing
 - First results expected for Spring 2023 stay tuned!
- A forward look: proposal for FPF, a dedicated forward physics facility @ CERN, to take advantage of HL-LHC and build a FASER2
 - Would give a rich and broad physics programme