## THE ROYAL SOCIETY

## FASER: First Results **Birmingham Seminar** Josh McFayden (University of Sussex) 18/10/2023

## UNIVERSITY **OF SUSSEX**

# EASER

JoshMcFayden aser.web.cern.ch









Background & Motivation

FASER construction & operation

 First direct observation of collider neutrinos

First observation of v<sub>e</sub> at collider

First FASER Dark Photon search

Looking ahead to HL-LHC







# **FASER Philosophy**

## LHC searches/experiments focus on heavy, strongly interacting particles Produced ~isotropically and at relatively low rates, especially in high $p_T$ regions ► $\sigma \sim \text{fb to pb} \rightarrow \text{In Run-3 N} \sim 10^2 - 10^5$









# **FASER Philosophy**

- Could be misguided Need to target light and weakly interacting particles Lack of new physics in "traditional" searches.

  - Scenarios that e.g. satisfy Dark Matter relic density.
  - Exploit huge inelastic pp cross section  $\rightarrow 10^{16}$  collisions in Run 3  $\rightarrow 10^{17} \pi$ ,  $10^{13}$  B
  - Light meson: low  $p_T \sim \Lambda_{QCD} \rightarrow$  particles are collimated:
    - $\theta \sim \Lambda_{QCD}/E \sim mrad$





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    - $\theta \sim \Lambda_{QCD}/E \sim mrad$
  - From only 10<sup>-8</sup> of solid angle 1% of  $\pi_0$ s are in acceptance.
  - Gain sensitivity to **long-lived particles** with very weak couplings and neutrinos







## **FASER** is a new experiment designed to cover this scenario at the LHC.







# **FASER Overview**





## Detector is 480m from ATLAS IP1

## Neutrinos produced copiously in decays of forward hadrons Very weakly interacting LLPs could be produced in significant numbers

## In line with beam collision axis. Transverse size of 10cm $\rightarrow$ mrad regime ( $\eta$ >9.1)





# **Example 1 Concept and Location**



## $\blacktriangleright \text{Old SPS} \rightarrow L$

- On line-of-si
- Shielded by
- Low beam b



18/10/2023

# **FASER Location** Wider setting at the LHC





# **FASER Location** In relation to ATLAS at Point 1

# ATLAS

forward jets

## FASER layout $\equiv$ high energy neutrino beamline

charged particles (P<7 TeV)



neutrino, dark photon

LHC tunnel

TI12 tunnel

 $FASER\nu$ 





# **FASER Location**

## A closer look at the LHC infrastructure on the line-of-sight:





#### FASER: ForwArd Search ExpeRiment at the LHC

Jonathan L. Feng,<sup>1,\*</sup> Iftah Galon,<sup>1,†</sup> Felix Kling,<sup>1,‡</sup> and Sebastian Trojanowski<sup>1,2,§</sup>

<sup>1</sup>Department of Physics and Astronomy, University of California, Irvine, CA 92697-4575 USA

> <sup>2</sup>National Centre for Nuclear Research, Hoża 69, 00-681 Warsaw, Poland

#### Abstract

New physics has traditionally been expected in the high- $p_T$  region at high-energy collider experiments. If new particles are light and weakly-coupled, however, this focus may be completely misguided: light particles are typically highly concentrated within a few mrad of the beam line, allowing sensitive searches with small detectors, and even extremely weakly-coupled particles may be produced in large numbers there. We propose a new experiment, ForwArd Search ExpeRiment, or FASER, which would be placed downstream of the ATLAS or CMS interaction point (IP) in the very forward region and operated concurrently there. Two representative on-axis locations are studied: a far location, 400 m from the IP and just off the beam tunnel, and a near location, just 150 m from the IP and right behind the TAN neutral particle absorber. For each location, we examine leading neutrino- and beam-induced backgrounds. As a concrete example of light,

13-12-

forward jets

IHC magnets

## 14th June 2017





18/10/2023

#### FASER: ForwArd Search ExpeRiment at the LHC

Jonathan L. Feng,<sup>1,\*</sup> Iftah Galon,<sup>1,†</sup> Felix Kling

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Submitted to the LHCC, 18 July 2018

CERN-LHCC-2018-030, LHCC-I-032 UCI-TR-2018-18, KYUSHU-RCAPP-2018-05

#### LETTER OF INTENT

## FASER FORWARD SEARCH EXPERIMENT AT THE LHC

Akitaka Ariga,<sup>1</sup> Tomoko Ariga,<sup>1,2</sup> Jamie Boyd,<sup>3,\*</sup> David W. Casper,<sup>4</sup> Jonathan L. Feng,<sup>4,†</sup> Iftah Galon,<sup>5</sup> Shih-Chieh Hsu,<sup>6</sup> Felix Kling,<sup>4</sup> Hidetoshi Otono,<sup>2</sup> Brian Petersen,<sup>3</sup> Osamu Sato,<sup>7</sup> Aaron M. Soffa,<sup>4</sup> Jeffrey R. Swaney,<sup>4</sup> and Sebastian Trojanowski<sup>8</sup>

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IHC magnets

## 14th June 2017

## 18th July 2018



18/10/2023

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## FASER layout $\equiv l$

17-17-

IHC magnets

forward jets

## 14th June 2017

## 18th July 2018

neutrino, dark photon

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CERN-LHCC-2018-036, LHCC-P-013 UCI-TR-2018-22, KYUSHU-RCAPP-2018-07

## 7th Nov 2018

nline

112 tur

FASERv

#### **TECHNICAL PROPOSAL**

## FASER

## FORWARD SEARCH EXPERIMENT AT THE LHC

Akitaka Ariga,<sup>1</sup> Tomoko Ariga,<sup>1,2</sup> Jamie Boyd,<sup>3,\*</sup> Franck Cadoux,<sup>4</sup> David W. Casper,<sup>5</sup> Francesco Cerutti,<sup>3</sup> Salvatore Danzeca,<sup>3</sup> Liam Dougherty,<sup>3</sup> Yannick Favre,<sup>4</sup> Jonathan L. Feng,<sup>5,†</sup> Didier Ferrere,<sup>4</sup> Jonathan Gall,<sup>3</sup> Iftah Galon,<sup>6</sup> Sergio Gonzalez-Sevilla,<sup>4</sup> Shih-Chieh Hsu,<sup>7</sup> Giuseppe Iacobucci,<sup>4</sup> Enrique Kajomovitz,<sup>8</sup> Felix Kling,<sup>5</sup> Susanne Kuehn,<sup>3</sup> Mike Lamont,<sup>3</sup> Lorne Levinson,<sup>9</sup> Hidetoshi Otono,<sup>2</sup> John Osborne,<sup>3</sup> Brian Petersen,<sup>3</sup> Osamu Sato,<sup>10</sup> Marta Sabaté-Gilarte,<sup>3,11</sup> Matthias Schott,<sup>12</sup> Anna Sfyrla,<sup>4</sup> Jordan Smolinsky,<sup>5</sup> Aaron M. Soffa,<sup>5</sup> Yosuke Takubo,<sup>13</sup> Pierre Thonet,<sup>3</sup> Eric Torrence,<sup>14</sup> Sebastian Trojanowski,<sup>15,16</sup> and Gang Zhang<sup>17</sup>

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charged particles (P<7 TeV)







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## [arXiv:1708.09389]

17

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## FASE

[arXiv:

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## 29th Nov 2018

## 18th July 2018

CERN-LHCC-2018-036, LHCC-P-013 UCL-TR-2018-22 KYUSHU-RCAPP-2018-07

## 7th Nov 2018

UCI-TR-2018-19, KYUSHU-RCAPP-2018-06



#### **FASER's Physics Reach for Long-Lived Particles**

#### **FASER** Collaboration

Akitaka Ariga,<sup>1</sup> Tomoko Ariga,<sup>1,2</sup> Jamie Boyd,<sup>3</sup> Franck Cadoux,<sup>4</sup> David W. Casper,<sup>5</sup> Pierre Thonet,<sup>3</sup> Eric Torrence,<sup>14</sup> Sebastian Yannick Favre,<sup>4</sup> Jonathan L. Feng,<sup>5</sup> Didier Ferrere,<sup>4</sup> Iftah Galon,<sup>6</sup> Sergio Gonzalez-Sevilla,<sup>4</sup> Shih-Chieh Hsu,<sup>7</sup> Giuseppe Iacobucci,<sup>4</sup> Enrique Kajomovitz,<sup>8</sup> Felix Kling,<sup>5</sup> Susanne Kuehn,<sup>3</sup> Lorne Levinson,<sup>9</sup> Hidetoshi Otono,<sup>2</sup> Brian Petersen,<sup>3</sup> Osamu Sato,<sup>10</sup> Matthias Schott,<sup>11</sup> Anna Sfyrla,<sup>4</sup> Jordan Smolinsky,<sup>5</sup> Aaron M. Soffa,<sup>5</sup> Yosuke Takubo,<sup>12</sup> Eric Torrence,<sup>13</sup> Sebastian Trojanowski,<sup>14,15</sup> and Gang Zhang<sup>16</sup>

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## Approved by CERN in March 2019~18 months from theory paper tocharged particles (P<7 TeV) forward jets start of construction! 29th Nov 2018

#### IHC magnets

## 14th June 2017

Submitted to the LHCC, 7 November 2018

#### **TECHNICAL F**

## FASE

## FORWARD SEARCH EXPE

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#### neutrino, dark photon









## **Ready to take data for start of LHC Run3** Detector Paper last year

# Approved by CERN in March 2019 high energy ~18 months from theory paper to<sup>charged</sup> particles (P<7 Te

start of construction. forward jets

IHC magnets

## 23rd July 2022

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#### The FASER Detector

#### **FASER Collaboration**

Henso Abreu<sup>1</sup>, Elham Amin Mansour<sup>2</sup>, Claire Antel<sup>2</sup>, Akitaka Ariga<sup>3,4</sup>, Tomoko Ariga<sup>5</sup>, Florian Bernlochner<sup>6</sup>, Tobias Boeckh<sup>6</sup>, Jamie Boyd<sup>7</sup>, Lydia Brenner<sup>8</sup>, Franck Cadoux<sup>2</sup>, David W. Casper<sup>9</sup>, Charlotte Cavanagh<sup>10</sup>, Xin Chen<sup>11</sup>, Andrea Coccaro<sup>12</sup>, Olivier Crespo-Lopez<sup>7</sup>, Stéphane Débieux<sup>2</sup>, Monica D'Onofrio<sup>10</sup>, Liam Dougherty<sup>7</sup>, Candan Dozen<sup>13</sup>, Abdallah Ezzat<sup>14</sup>, Yannick Favre<sup>2</sup>, Deion Fellers<sup>15</sup>, Jonathan L. Feng<sup>9</sup>, Didier Ferrere<sup>2</sup>, Edward Karl Galantay<sup>2</sup>, Jonathan Gall<sup>16</sup>, Enrico Gamberini<sup>7</sup>, Stephen Gibson<sup>17</sup>, Sergio Gonzalez-Sevilla<sup>2</sup>, Carl Gwilliam<sup>10</sup>, Daiki Hayakawa<sup>4</sup>, Shih-Chieh Hsu<sup>18</sup>, Zhen Hu<sup>11</sup> , Giuseppe Iacobucci<sup>2</sup>, Tomohiro Inada<sup>11</sup>, Sune Jakobsen<sup>7</sup>, Eliott Johnson<sup>2</sup>, Enrique Kajomovitz<sup>1</sup>, Hiroaki Kawahara<sup>5</sup>, Felix Kling<sup>19</sup>, Umut Kose<sup>7</sup>, Rafaella Kotitsa<sup>7</sup>, Jesse Krusse<sup>20,21</sup>, Susanne Kuehn<sup>7</sup>, Helena Lefebvre<sup>17</sup>, Lorne Levinson<sup>22</sup>, Ke Li<sup>18</sup>, Jinfeng Liu<sup>11</sup> , Chiara Magliocca<sup>2</sup>, Fulvio Martinelli<sup>2</sup>, Josh McFayden<sup>23</sup>, Sam Meehan<sup>7,24</sup>, Matteo Milanesio<sup>2</sup>, Manato Miura<sup>4</sup>, Dimitar Mladenov<sup>7</sup>, Théo Moretti<sup>2</sup>, Magdalena Munker<sup>2</sup>,

neutrino, dark photon

FASERv







## FASER: ForwArd Search ExpeRiment

"The acronym recalls another marvelous instrument that harnessed highly collimated particles and was used to explore strange new worlds."











**Example 7 Physics Motivation** The LHC experiments are producing incredible results, searching in measurements.

But the lack of any observation of BSM physics motivates **looking elsewhere** too.





# previously unexplored phase spaces and performing increasingly precise





## **Example 2 Physics Motivation** The indirect observations of dark matter offers one of the most tangible indictions of BSM physics and strongly motivates closer attention.











# **Example 2 Physics Motivation**

- indictions of BSM physics and strongly motivates closer attention.
- Main region of interest is for new particles that satisfy DM relic density requirements.

SM

SM





# The indirect observations of dark matter offers one of the most tangible



# **Example 7 Physics Motivation**

- One of the defining characteristics of weakly interacting light particles is their long lifetime.
- Distinct signatures
- Opportunity for exploration!











# **Detector Construction and Operation**



# **F** Philosophy

- limited budget we have focused on:
  - Detector that can be constructed and installed quickly & cheaply
  - Have tried to re-use existing detector components where possible
  - Aimed for a simple, robust detector (access difficult)
  - Tried to minimize the services to simplify the installation and operations
- Rough dimensions
  - 10 cm radius, 1.5 m decay volume, ~7 m total length
- Many challenges of the large LHC experiments not there for FASER: trigger rate ~500Hz (mostly single muon events)

  - Iow radiation
  - Iow occupancy / event size

Given the very tight timeline between experiment approval and installation & the



# **Example 1 Example 1 Constallation**





8/2019





3/2021

8/2018

4/2020











# **Example 7 Detector design**

## Small inexpensive design [2207.11427]

## **Tracking spectrometer stations**

3 x 3 layers of ATLAS SCT strip modules



## **Trigger / pre-shower** scintillator system

Magnets

0.57 T Dipoles 1.5 m decay volume

## **Scintillator** veto system

2 x 20 mm thick 30 x 30 cm area

Decay volume

## **Front Scintillator**

## veto system

2 x 20 mm thick 35 x 30 cm area

TO ATLAS IP

## Interface Tracker (IFT)

## **FASER**v emulsion detector

730 layers of 1.1 mm tungsten + emulsion (8 interaction lengths)

## **Trigger / timing** scintillator station

10mm thick + dual PMT readout ( $\sigma = 400 \text{ ps}$ )











# **E**Installation







## FASER layout $\equiv$ high energy neutrino beamline

IHC magnets

forward jets

Martin Martin



charged particles (P<? TeV)



Decay

## FASER layout $\equiv$ high energy neutrino beamline

forward jets



charged particles (P<? TeV)

# **E** Operations

- Successfully operated throughout 2022
  - Continuous data taking
  - Largely automated
  - Up to 1.3 kHz
- Recorded 96.1% of delivered lumi.
  - DAQ dead-time of 1.3%
  - couple of DAQ crashes
- Emulsion detector exchanged twice
  - Needed to manage occupancy
  - First box only partially filled
- Calorimeter gain optimised for:
  - Low E (<300 GeV) before 2nd exchange</p>
  - High E (up to 3 TeV) after this exchange



Analyses presented use 27.0 fb<sup>-1</sup> or 35.4 fb<sup>-1</sup>












## File Analysis

- Neutrinos produced copiously in decays of forward hadrons
  - ► Highly energetic (TeV scale) → high interaction cross section
- Extends FASER physics program into SM measurements
  - Targets measurement of highest energy human-made neutrinos
  - Energy range complementary to existing neutrino experiments







## **E** Neutrino analysis

- Neutrinos produced copiously in decays of forward hadrons
  - Highly energetic (TeV scale)  $\rightarrow$  high interaction cross section
- Extends FASER physics program into SM measurements
  - Targets measurement of highest energy human-made neutrinos
  - Energy range complementary to existing neutrino experiments





## F Neutrino analysis

- Neutrinos produced copiously in decays of forward hadrons
  - $\blacktriangleright$  Highly energetic (TeV scale)  $\rightarrow$  high interaction cross section
- Extends FASER physics program into SM measurements
  - Targets measurement of highest energy man-made neutrinos
  - Energy range complementary to existing neutrino experiments

For 35 fb <sup>-1</sup>	Ve	Vμ	ντ
Main source	Kaons	Pions	Charm
# traversing FASERv	~1010	~1011	~108
# interacting in FASERv	≈200	≈1200	≈4

[PRD 104, 113008]

## Study at colliders originally proposed by Rújula and Rückl in 1984! Josh McFayden | Birmingham | 18/10/2023



## **F** Neutrino analysis

- 1. Collision event with good data quality
- 2. No signal (<40 pc) in 2 front vetos</p>
- 3. Signal (>40 pC) in other 3 vetos



### Selection

### [2303.14185

- ▶ 4. Timing and preshower consistent with  $\geq$ 1 MIP
- 5. Exactly 1 good fiducial (r < 95 mm) track</p>
  - ▶ p > 100 GeV and  $\theta$  < 25 mrad
  - Extrapolating to r < 120 mm in front veto</p>



## File Neutrino analysis Selection

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Time [ns]



eto station, laver 3





### [2303.14185]

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  - ▶ p > 100 GeV and  $\theta$  < 25 mrad
  - Extrapolating to r < 120 mm in front veto</p>

Josh McFayden | Birmingham |





## Neutrino analysis | Backgrounds

### Neutral hadrons

- Estimated from 2-step simulation
- Expect ~300 neutral hadrons with E>100 GeV reaching FASERv
- Most accompanied by µ but conservatively assume missed
- Estimate fraction of these passing event selection
- Most are absorbed in tungsten with no highmomentum track
- Predict N =  $0.11 \pm 0.06$  events



## F Neutrino analysis | Backgrounds

### Scattered muons

- Estimated from data sideband
- Take events w/o front veto radius requirement and single track segment in first tracker station with 90 < r < 95 mm
  - Fit to extrapolate to higher momentum
- Scale by # events with front veto cut
  - Use MC to extrapolate to signal region
- Predict N =  $0.08 \pm 1.83$  events
- Uncertainty from varying selection



[2303.14185]











## **Neutrino analysis** | Backgrounds

### Veto inefficiency

- Estimated from final fit
- Fit events with 0 (SR) and also 1 (1st or 2nd) or 2 front veto layers firing
- Final negligible background due to very high veto efficiency

Category	Events	Expectation
Signal		$n_{ u} + n_b \cdot p_1 \cdot p_2 + n_{ m had} + n_{ m geo} \cdot f_{ m geo}$
$n_{10}$	4	$n_b \cdot (1-p_1) \cdot p_2$
$n_{01}$	6	$n_b \cdot p_1 \cdot (1-p_2)$
$n_2$	64014695	$n_b \cdot (1-p_1) \cdot (1-p_2)$





**F** Neutrino analysis | Results Unblinded to find 153 events with no veto signal

Just 10 events with one veto signal

### First direct detection of collider neutrinos!

- With signal significance of 16σ
- Expected  $151 \pm 41$ events from GENIE simulation





[2303.14185]

# For the second second

- Candidate neutrino events match expectation from signal
  - High occupancy in front tracker station
  - Most events have high µ momentum
  - More  $v_{\mu}$  than anti-  $v_{\mu}$
  - High occupancy in front tracker station
  - Large angle θ with respect to LOS
- NB: no acceptance corrections nor any systematic uncertainties in these plots



# Emulsion analysis

- Emulsion detector with 1.1T tungsten target
  - 730 1.1mm thick tungsten plates interleaved with emulsion film
    - Well understood neutrino detector technology
  - Replace every 20-50 fb<sup>-1</sup> to maintain track density low

### Challenges:

- Logistics to transport and replace the 1-ton-scale detector every technical stop (3 times/year)
- Procedure well developed for production and offline analysis
- New analysis!

ν





### Josh McFayden Birmingham I



## **Emulsion analysis** | Dataset & Bkgds

- First analysis includes 150 of 730 plates
  - 68kg target mass for this analysis (24 x 9 x 16.5 cm)
  - 9.5 fb<sup>-1</sup> of LHC proton collision data
- Expect 29.4  $\pm$  5.0 ( $v_{\mu}$ ) and 11.8  $\pm$  7.5 ( $v_{e}$ ) charged current (CC) neutrino interactions before selection
  - Select vertices with associated lepton candidate (e or  $\mu$ ) and E>200 GeV

### Backgrounds

- Neutral hadron background; low-momentum signal
- Neutral-current neutrino interactions

Scanned volume (255 films) (this time) **FASER***v* 730 films and  $\nu \longrightarrow$ tungsten plates

> Target volume for the first analysis (150 tungsten plates)











**FASER** Preliminary



500 μm











### First direct observation of collider electron neutrinos!





### **Emulsion analysis** | Results Preliminary results: <u>CERN-FASER-CONF-2023-002</u> Expected 0.6–5.2 ( $v_e$ CC) and 3.0–8.6 ( $v_\mu$ CC) passing selection • Observed 3 $v_e$ vertices (5 $\sigma$ ), and 4 $v_\mu$ vertices (2.5 $\sigma$ ) - **Candidates with E ~ 1TeV!**

**FASER Preliminary** 



1000 μm











### First direct observation of collider electron neutrinos!

# • Observed 3 $v_e$ vertices (5 $\sigma$ ), and 4 $v_\mu$ vertices (2.5 $\sigma$ ) - **Candidates with E ~ 1TeV!**







**E** Dark Photon Search

- Dark photon a common feature of hidden sector models
  - Weakly coupling to SM via kinetic mixing (ε) with SM photon

$$\mathcal{L} \supset \frac{1}{2} m_{A'}^2 A'^2 - \epsilon e \sum_{f} q_f \bar{f} A' f$$

MeV A's produced mainly in meson decays at LHC

$$\pi^{0} - \cdots - \left( \int_{A'}^{\gamma} B(\pi^{0} \to A'\gamma) = 2\epsilon^{2} \left( 1 - \frac{m_{A'}^{2}}{m_{\pi^{0}}^{2}} \right)^{3} B$$

FASER targets small ε, where A' has long decay length

$$L = c\beta\tau\gamma \approx (80 \text{ m}) \left[\frac{10^{-5}}{\epsilon}\right]^2 \left[\frac{E_{A'}}{\text{TeV}}\right] \left[\frac{100 \text{ M}}{m_A}\right]$$

Below 2m<sub>u</sub>, A' has 100% decay to e<sup>+</sup>e<sup>-</sup> pair



- $S(\pi^0 \to \gamma \gamma)$
- [Val

\* arXiv:2105.07077



- $A' \rightarrow e+e-simulated with FORESEE*$ 
  - $\pi_0$  and  $\eta$  via EPOS-LHC generator
  - Subdominant dark brem. via FWW
- Generator uncertainty dominates
  - Difference to QGSJET/SIBYLL
  - Parameterised based on A' energy Josh McFayden | Birmingham | 18/10/2023





### **Dark Photon Search** | Selection

- Simple and robust  $A' \rightarrow e^+e^-$  selection
  - Blind events with no veto signal and E(calo) > 100 GeV

### Selection

- 1. Collision event with good data quality
- 2. No signal (< 40 pc) in any veto scintillator</p>
- ▶ 3. Timing and preshower consistent with  $\geq$ 2 MIPs
- 4. Exactly 2 good fiducial tracks
  - p > 20 GeV and r < 95 mm && Extrapolating to r < 95 mm at vetos</p>
- 5. Calo E > 500 GeV





Efficiency of ~40% across sensitive region



### **Dark Photon Search** | Selection

- Simple and robust  $A' \rightarrow e^+e^-$  selection
  - Blind events with no veto signal and E(calo) > 100 GeV

### Selection

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- 4. Exactly 2 good fiducial tracks
  - p > 20 GeV and r < 95 mm && Extrapolating to r < 95 mm at vetos</p>
- ► 5. Calo E > 500 GeV





sensitive region











# *F* Dark Photon | Backgrounds

### Veto inefficiency

- Measured layer-by-layer via muons with tracks pointing back to vetos
- Layer efficiency > 99.998%
- ▶ 5 layers reduce exp. 10<sup>8</sup> muons to negligible level (even before cuts) (<10<sup>-20</sup> inefficiency)



- Non-collision backgrounds
  - Cosmics measured in runs with no beam
  - Near-by beam debris measured in noncolliding bunches
  - No events observed with  $\geq 1$  track or E(calo) > 500 GeV individually







## **Dark Photon** | Backgrounds

- Main background is from Neutrino interactions
  - Primarily coming from vicinity of timing detector
  - Estimated from GENIE simulation (300 ab-1)
  - Uncertainties from neutrino flux & mismodelling
  - Predicted events with E(calo) > 500 GeV

### $N = (1.5 \pm 2.0) \times 10^{-3}$

- Neutral hadrons (e.g. Ks) from upstream muons interacting in rock in front of FASER
  - Heavily suppressed since:
    - muon nearly always continues after interaction
    - has to pass through 8 interaction lengths (FASERv)
    - decay products have to leave E(calo) > 500 GeV
  - Estimated from lower energy events with 2/3 tracks and different veto conditions

### $N = (0.8 \pm 1.2) \times 10^{-3}$



[arXiv:2308.05587]





- No events in unblinded
  - with  $\geq 1$  fiducial track

10' 10<sup>6</sup>





## **Dark Photon Search** | Limits

- unexplored parameter space.
- First incursion (with NA62) into thermal relic region from low ε since 1990's.
- Background-free analysis bodes well for future sensitivity.
- Expect ~10fb<sup>-1</sup> luminosity in Run 3 from 2023-25.



After unblinding, no events seen in signal region, FASER sets limits on previously

Kinetic Mixing. -010-3  $L = 27.0 \text{ fb}^{-1}$ Expected Limit ( $\pm 1 \sigma_{exp}$ , 90% CL) Observed Limit (90% CL) NA62 (µµ) Limi BaBar Limit KLOE Limit LHCb Limit NA48 Limit  $10^{-5}$ NA64 Limit E141 Limit Orsay Limit NuCal Limit 10<sup>-6</sup> E137 Limit CHARM Limi Relic Target m<sub>2</sub>=0.6m<sub>4</sub>,  $\alpha_{p}$ =0.1  $10^{-7}$ . . . . . 10<sup>2</sup> 10  $10^{\circ}$ m<sub>A'</sub> [MeV] Josh McFayden



Birmingham | 18/10/2023

# **F** B-L Gauge Boson Search | Limits

- The region probed is cosmologically relevant
  - limit includes region favoured by thermal freeze-out
  - mechanism, resulting relic density may be significant in excluded region.



Assuming a dark matter particle with mass  $\frac{1}{2}m(A'_{B-L}) - m(A'_{B-L})$  & very large Q(B-L),

► As B-L model includes 3 sterile neutrinos, that could be produced through freeze-in









## **Dark Photon Search** | Limits

### Preshower Upgrade

- More transverse information
- Beneficial for 2-photon signals, such as ALP searches
- Installation planned for YETS 24/25
  - Technical proposal

### Run 3 and Beyond

- Expect to collect 10x more data in Run 3
- Excellent performance so far therefore...
- Begun request process to continue operations after LS3
- during the HL-LHC era



Potentially succeeded by FASER2/FASERv2 in the planned Forward Physics Facility (FPF)







# **Example 7** Forward Physics Facility

- FASER, FASERv, and other proposed detectors are currently highly support experiments.
- QCD, dark matter, dark sector, cosmic rays, and cosmic neutrinos.
- forward experiments

constrained by tunnels and infrastructure that was never designed to

• At the same time, it is becoming clear that there is a **rich physics program** in the far-forward region, spanning long-lived particle searches, neutrinos,

Strongly motivates to create a dedicated facility to house several far-







## **Example 7** Forward Physics Facility

- in Europe and abroad:
  - 2020 European Strategy Update:
    - "The full physics potential of the LHC and the HL- LHC...should be exploited"
    - *laboratories in Europe should be supported"*

### Snowmass 2021 Energy Frontier Report:

upgrades"



### The FPF is well aligned with the recommendations of recent community studies

• "The quest for dark matter and the exploration of flavour and fundamental symmetries are crucial components of the search for new physics. This search can be done in many ways, for example through ... searches for axions, dark sector candidates and feebly interacting particles. ...A diverse programme that is complementary to the energy frontier is an essential part of the European particle physics Strategy. Experiments in such diverse areas that offer potential high-impact particle physics programmes at

• "Our highest immediate priority accelerator and project is the HL-LHC,...including the construc- tion of auxiliary experiments that extend the reach of HL-LHC in kinematic regions uncovered by the detector







# **Forward Physics Facility**



### FPF Papers: FPF White Paper: J. Phys. G (2022) milliquirks axions dark matter charged particles $\boldsymbol{a}$ $\gamma$



FPF core sample to study site geology, refine cost estimates

Josh McFayden | Birmingham |









- Proof of principle now exists from FASER
- Slightly different design philosophy limited by large aperture magnet technologies
- Program for BSM and SM physics (main spectrometer to neutrino exps.)
- Currently considering, SciFi tracker and dual-readout calorimetry.







- charge ID





Increasing detector radius to 1m would allow sensitivity to new physics produced in heavy meson (B, D) decays increasing the physics case beyond just the increased luminosity.











- Increasing detector radius to 1m would allow sensitivity to new physics produced in heavy meson (B, D) decays increasing the physics case beyond just the increased luminosity.
- FASER2 therefore becomes very strong compared to low energy experiments for certain models (dark Higgs), due to large B/D production rates at LHC:
- N<sub>B</sub>/N<sub> $\pi$ </sub>~10<sup>-2</sup> (~10<sup>-7</sup> at beam dump expts)






# **FASER2**

### FASER2 (R = 1 m, L = 5-20 m) can discover

- All candidates with renormalizable couplings (dark photon, dark Higgs, HNL)
- ALPs with all types of couplings  $(\chi, f, g)$
- and many other particles.

Among the PBC benchmark scenarios, FASER2's discovery potential extends to all benchmark scenarios

Except BC2 and BC3.

Benchmark Model	FASER	FASEF
BC1: Dark Photon	$\checkmark$	$\checkmark$
BC1': U(1) <sub>B-L</sub> Gauge Boson	$\checkmark$	
BC2: Invisible Dark Photon	_	-
BC3: Milli-Charged Particle	_	_
BC4: Dark Higgs Boson	_	$\checkmark$
BC5: Dark Higgs with hSS	_	$\checkmark$
BC6: HNL with e	_	$\checkmark$
BC7: HNL with $\mu$	_	$\checkmark$
BC8: HNL with $\tau$	$\checkmark$	$\checkmark$
BC9: ALP with photon	$\checkmark$	$\checkmark$
BC10: ALP with fermion	$\checkmark$	$\checkmark$
BC11: ALP with gluon	$\checkmark$	$\checkmark$





# **E Neutrino physics**

- Study neutrino interactions at high energy
  - Search for BSM physics in neutrino production, propagation and interaction
  - Study PDFs by DIS of neutrino in the target (fixed target 75 GeV CoM E)
  - Study forward hadron production measureme FASER layou







### **Forward Physics Facility** | Neutrinos The FPF experiments will see $10^5 v_e$ , $10^6 v_\mu$ and $10^4 v_\tau$ interactions at E~TeV where there is currently no data



Birmingham | 18/10/2023 Josh McFayden



# **Forward Physics Facility** | Neutrinos

- spectra will inform:
- Astroparticle physics: muon puzzle, ...
- QCD: pdfs at  $x \sim 10^{-1}$ ,  $x \sim 10^{-7}$ , intrinsic charm, small-x gluon saturation, ...
- Neutrino properties: short-baseline neutrino experiment. Sensitive to sterile neutrinos

Neutrinos are produced by forward hadron production: π, K , D.... Energy







### **Forward Physics Facility** | Neutrinos Neutrinos are produced by forward hadron production: π, K , D.... Energy

- spectra will inform:
- Astroparticle physics: muon puzzle, ...
- QCD: pdfs at  $x \sim 10^{-1}$ ,  $x \sim 10^{-7}$ , intrinsic charm, small-x gluon saturation, ...
- Neutrino properties: short-baseline neutrino experiment. Sensitive to sterile neutrinos
- Fully differential neutrino DIS scattering XS will improve constraints on pdfs by up to ~2x













- First direct observation of collider neutrinos
  - Opens a new field: neutrino physics at the LHC
  - Published in PRL [2303.14185]
- First Dark Photon search
  - First limit in thermal relic region from low coupling for 30 yrs
  - Submitted to [arXiv:2308.05587]
- High-energy  $v_e$  interactions in emulsion detector
  - [CERN-FASER-CONF-2023-002]
- More neutrino studies and BSM searches to come
  - Including searches for ALPs, light gauge bosons, ...
- Strongly motivates FPF & FASER2 for HL-LHC era





### Josh McFayden | FPF UK | 11/10/2023





 $10^{2}$ 





First dire

Opens a Published

First Dark

First limit

Submitte

High-ene ► [CERN-FA

### More neu

Including



### Josh McFayden | FPF UK | 11/10/2023





















### And would additionally like to thank

- LHC for the excellent performance in 2022
- ATLAS for providing luminosity information
- ATLAS for use of ATHENA s/w framework
- ATLAS SCT for spare tracker modules
- LHCb for spare ECAL modules
- CERN FLUKA team for background sim
- CERN PBC and technical infrastructure groups for excellent support during design construction and installation



## **F** Collaboration

- 87 members
- 24 institutions
- 10 countries



BERN

**b UNIVERSITÄT** 















7.





### UNIVERSITÉ DE GENÈVE





UNIVERSITY of WASHINGTON



CHIBA UNIVERSITY

International laboratory

covered by a cooperation

agreement with CERN

MAGOYA UNIVERSITY









Tsinghua University



The University of Manchester

















### Detector Performance: Trigger + DAQ DAQ running smoothly up to 1.3 kHz with deadtime only 1.3% Total trigger rate falls off faster than luminosity during run (higher beam-

induced backgrounds) but coincidence trigger rate flat with respect to

luminosity





# **FASERV** Physics case

- The energy spectrum expected at FASERv is rather complementary to existing neutrino experiments
- Expected cross section sensitivity significantly extends current measurements during Run 3 (150 fb<sup>-1</sup>)
- Being located on line-of-sight FASERv is able to observe a maximum rate of all neutrino flavours:





## Neutrino analysis



### [arXiv:1908.02310]



	Particles	Main Decays	E	Q	S	
	$\pi^+$	$\pi^+ \to \mu \nu$	$\checkmark$	$\checkmark$		
	$\mathrm{K}^+,K_S,K_L$	$K^+ \to \mu \nu,  K \to \pi \ell \nu$	$\checkmark$	$\checkmark$		
$\mathbf{S}$	$\Lambda, \Sigma^+, \Sigma^-, \Xi^0, \Xi^-, \Omega^-$	$\Lambda \to p \ell \nu$	$\checkmark$	$\checkmark$		
	$D^+, D^0, D_s, \Lambda_c, \Xi_c^0, \Xi_c^+$	$D \to K \ell \nu, D_s \to \tau \nu, \Lambda_c \to \Lambda \ell \nu$			$\checkmark$	
	$B^+, B^0, B_s, \Lambda_b, \ldots$	$B \to D\ell\nu, \Lambda_b \to \Lambda_c\ell\nu$				







# **Example 7 Neutrino analysis**

[PRD 104, 113008]









# **Emulsion analysis** | Results





## Emulsion analysis | Results

Selection	Quantity	$K_L$	n	Λ	Total
$\nu_{\mu}$ selection $E > 50 \text{ GeV}$	$_{\iota}$ selection $E > 50 \text{ GeV}$ Raw MC events		71	70	-
$\mu$ solution $E > 200 \text{ CoV}$	Raw MC events	1	3	2	-
$\nu_{\mu}$ selection $E > 200 \text{ GeV}$	Scaled to analysis dataset	0.07	0.16	0.10	0.32
$\nu_e$ selection $E > 50$ GeV	Raw MC events	0	1	0	-
$\mu$ soluction $E > 200 \text{ GeV}$	Raw MC events	0	0	0	-
$\nu_e$ selection $E > 200 \text{ GeV}$	Scaled to analysis dataset	0	0	0	0

Background	$\nu_{\mu} \ \mathrm{CC}$	$\nu_e \ \mathrm{CC}$
Neutral-hadron interactions	$0.32 \pm 0.15 \text{ (stat.)} \pm 0.16 \text{ (syst.)}$	$0.002 \pm 0.002$ (stat.) $\pm 0.002$ (syst.)
NC neutrino interactions	$0.19\pm0.15$	_
Total	$0.51 \pm 0.27$	$0.002 \pm 0.003$





## **Emulsion analysis** | HTS

- HTS: https://arxiv.org/abs/1704.06814
- The track reconstruction algorithm is based on that of the NA65/DsTau experiment and is described in <u>http://arxiv.org/abs/1906.03487</u>.
- Vertex: Convergence of >4 tracks, >3 tracks with  $tan(\theta) < 0.1$

Charged vertex: Looser track selection, tracks within 10 films before vertex, with 3 track hits,  $d0 <= 5 < -\mu m$ , min. distance to 3 vertex tracks of  $<= 3\mu m$ .







# **FASERv** | Rich neutrino physics program

### BSM physics

- New light weakly coupled gauge boson ( $\rightarrow v_{\tau}$ ) could enhance  $v_{\tau}$  flux.
- Sterile neutrinos with mass ~40 eV can cause oscillations at FASER

### ► QCD

- FASER's neutrino flux measurements will provide novel complimentary constraints that can be used to validate/improve MC generator very forward particle production.
- Neutrinos from charm decay could allow to test transition to small-x factorisation, constrain low-x gluon PDF and probe intrinsic charm

### Cosmic rays and neutrinos

- IceCube needs measurements of high energy and large rapidity charm for precise measurements of cosmic neutrino flux.
- Direct measurement of prompt neutrino production at FASER would provide important data for current & future neutrino telescopes



 $10^{-1}$ 

10-2 -

10-

10-

 $10^{-5}$ 

10-6 -

 $10^{-3}$ 

CDF

LESB

Birmingham | 18/10/2023 Josn Mic⊢ayden

 $B - 3L_{\tau}$  Gauge Boson

 $\pi^0 \to V\gamma$ ,

FixedTarget

Tevatron

HERA

LHC

 $10^{-1}$ 

 $m_{A'}$  [GeV]

LHCv<sub>e</sub> with  $7 < \eta_v < 8$ LHCv<sub>e</sub> with  $8 < \eta_v < 9$ 

 $10^{0}$ 

 $10^{-2}$ 





 $10^{-1}$ 

# **Example 7 Dark Photon** | Signal

- Acceptance 10<sup>-6</sup>
- Decay volume 10-8 solid angle
- $P(decay in FASER) = 10^{-3}$



$$L = c\beta\tau\gamma \approx (80 \text{ m}) \left[\frac{10^{-5}}{\epsilon}\right]^2 \left[\frac{E_{A'}}{\text{TeV}}\right] \left[\frac{100 \text{ MeV}}{m_{A'}}\right]$$









## **Example 7 Dark Photon** | Selection

Description	Value	
Pre-selecton		
Time consistent with a colliding bunch identifier		
Timing scintillator trigger		
Scintillator		
Timing station:		
Top or Bottom Scintillator charge	$> 70 \ {\rm pC}$	
OR Top and Bottom charge	$> 30 \ \mathrm{pC}$	
Each Preshower scintillator charge	>2.5  pC	
Each Veto scintillator charge	$<\!40~{ m pC}$	
Tracking		
Exactly 2 Good Tracks		
Momentum	$> 20 { m ~GeV}$	
$\chi^2/\mathrm{NDF}$	< 25	
Number of tracker layers on track	>=7	
Number of tracker hits on track	>= 12	
Fiducial selection		
Track extrapolated to all scintillators		
and tracking stations	$< 95 \mathrm{mm}$	
Calorimeter		
Calorimeter energy (sum of four channels)	$> 500 { m ~GeV}$	

TABLE I. Summary of selection requirements.



Selection Criteria	Efficiency
Good collision event	99.7%
No Veto Signal	98.4%
Timing/Preshower Signal	97.3%
$\geq 1 \text{ good track}$	89.2%
= 2  good tracks	44.5% *
Track radius $<95~\mathrm{mm}$	42.3% *
Calo energy $> 500~{\rm GeV}$	41.6% *

 $\epsilon = 3 \times 10^{-5} \text{ m}_{A'} = 25.1 \text{ MeV}$ 



## **E** Dark Photon | ABCD method

Selection	Nevents E<100 GeV	Ne
3 tracks (VetoNu signal)	544.7	
2 tracks (No VetoNu signal)	1	

events E > 500 GeV11.0

Predicted: 0.02



## **Example 7** Dark Photon | Scintillator efficiencies



Scintillator	Efficiency
NuVeto-0	0.9999805(5)
NuVet0-1	0.9999810(5)
Veto-0	0.9999985(1)
Veto-1	0.9999984(1)
Veto-2	0.9999986(1)





## **E** Dark Photon | Performance





Source	Value	Effect on signal yield
Signal Generator	$\frac{0.15{+}(E_{A'}/4{\rm TeV})^3}{1{+}(E_{A'}/4{\rm TeV})^3}$	$15-65\% \ (15-45\%)$
Luminosity	2.2%	2.2%
MC Statistics	$\sqrt{\sum W^2}$	1-3%~(1-2%)
Track Momentum Scale	5%	< 0.5%
<b>Frack Momentum Resolution</b>	5%	< 0.5%
Single Track Efficiency	3%	3%
Two-track Efficiency	7%	7%
Calo E scale	6%	0-8%~(<1%)



### **Forward Physics Facility** BSM FPF Papers: FPF "Short" Paper: Phys. Rept. 968, 1 (2022) $10^{-3}$







### **Forward Physics Facility** | Neutrinos FPF Papers: 0.9 accelerator data



## **FASER Location**

### A closer look at the LHC infrastructure on the line-of-sight:





# **Example 5** Magnets | Overview

- The FASER magnets are 0.55T permanent dipole magnets based on the Halbach array design
  - Thin enough to allow the LOS to pass through the magnet centre with minimum digging to the floor in TI12
  - Minimize needed services (power, cooling etc..)
- Designed and constructed by magnet group at CERN



Pictured: longitudinal cut







### **Example 5** Magnets | Construction and testing Assembly at CERN of all 3 magnets completed, and all magnets measured at

- CERN
- Measured field quality well within specifications.







# **Example 5** Magnets Installation

- CERN
- Measured field quality well within specifications.
- All magnets now installed underground!



### Assembly at CERN of all 3 magnets completed, and all magnets measured at



## **Example 7 Fracker** | Modules

- Spare ATLAS SCT modules are used
  - 80µm strip pitch, 40mrad stereo angle (17µm / 580µm resolution)
    - precision measurement in bending (vertical) plane
  - Many thanks to the ATLAS SCT collaboration!











### **FTracker** Layers 8 SCT modules give a 24cm x 24cm tracking layer ▶ 9 layers (3/station, 3 stations) $\rightarrow$ 72 SCT modules needed for the full tracker













## **Figure 7 Tracker** | Stations

- Low radiation levels in TI12 allows silicon to be operated at room temp.
- used in Baby MIND neutrino experiment)





But the detector needs to be cooled to remove heat from the on-detector ASICs Tracker readout using FPGA based board from University of Geneva (already









# **E** Overground testing

- Have space at CERN Prevessin site (same building as Neutrino Platform)
- Used for dry run above ground
  - Assembly took place in Feb-April 2020
    - Test mechanical assembly
  - Commissioning from March 2020
    - Detector installation
    - Alignment procedures
    - Cabling
    - Cooling
    - TDAQ
    - Cosmics runs

Feb Mar Nov Dec Jan April Prepare ENH1 Install Det. Support Josh McFayden | Birmingham | 18/10/2023

- Install Calo/Scin & TDAQ
- (Partial) System Commissioning





# *i* Overground testing | Tracker

- Use full FASER TDAQ system to take data.
  - Operational experience
  - Tracker efficiency, resolution and alignment studies
  - Offline s/w debugging



# Cosmic data taking with station on its side, and a scintillator on top/btm.






### *v* **Overground testing** | Tracker Straight track candidate along with event display:







### **Example 2 Commissioning** | Overground Also have partial detector combined run All scintillators and calorimeters with one tracker station





# **Example 2** Commissioning | Overground Also have partial detector combined run

- All scintillators and calorimeters with one tracker station
- In just one week before disassembly started:
  - Common clock provided by Technion clock card (40.08 MHz)
  - Triggering on cosmic showers/random triggers
  - Reading out full detector
  - Tracker readout-timed in with respect to trigger signals
  - Ran with FASER DAQ system, run control GUI and monitoring
  - Data recorded to local disk and copied to EOS



# **FASERV** Physics case

- The energy spectrum expected at FASERv is rather complementary to existing neutrino experiments
- Expected cross section sensitivity significantly extends current measurements during Run 3 (150 fb<sup>-1</sup>):





# **FASERV** Physics case

- The energy spectrum expected at FASERv is rather complementary to existing neutrino experiments
- Expected cross section sensitivity significantly extends current measurements during Run 3 (150 fb<sup>-1</sup>)
- Being located on line-of-sight FASERv is able to observe a maximum rate of all neutrino flavours:





## **FASERv** | Detector design

### Emulsion detector with tungsten target

- 1000 1mm thick tungsten plates interleaved with emulsion film
  - Well understood neutrino detector technology
- Replace every 20-50 fb<sup>-1</sup> to maintain track density low

### Challenges:

- Logistics to transport and replace the 1-ton-scale detector every technical stop (3 times/year)
  - Benefit from transport infrastructure installed in UJ12 and TI12 to install FASER detector
- Procedure well developed for production and offline analysis:





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### **FASERV** | Pilot neutrino detector A 30 kg detector was installed in TI18 in 2018 ▶ 12.5 fb<sup>-1</sup> of data was collected

- ~30 neutrino interactions in the detector expected to have occurred
- Emulsion data developed, reconstructed and analysis ongoing
- Extremely valuable for validating the FASERnu, optimizing the detector & reconstruction Several neutral vertices identified, likely to be neutrino interactions, could also be neutral hadrons











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# **FASERv** | Rich neutrino physics program

### BSM physics

- New light weakly coupled gauge boson ( $\rightarrow v_{\tau}$ ) could enhance  $v_{\tau}$  flux.
- Sterile neutrinos with mass ~40 eV can cause oscillations at FASER

### ► QCD

- FASER's neutrino flux measurements will provide novel complimentary constraints that can be used to validate/improve MC generator very forward particle production.
- Neutrinos from charm decay could allow to test transition to small-x factorisation, constrain low-x gluon PDF and probe intrinsic charm

### Cosmic rays and neutrinos

- IceCube needs measurements of high energy and large rapidity charm for precise measurements of cosmic neutrino flux.
- Direct measurement of prompt neutrino production at FASER would provide important data for current & future neutrino telescopes



 $10^{-1}$ 

10-2 -

10-

10-

 $10^{-5}$ 

10-6 -

 $10^{-3}$ 

CDF

LESB

Birmingham | 18/10/2023 Josn Mic⊢ayden

 $B - 3L_{\tau}$  Gauge Boson

 $\pi^0 \to V\gamma$ ,

FixedTarget

Tevatron

HERA

LHC

 $10^{-1}$ 

 $m_{A'}$  [GeV]

LHCv<sub>e</sub> with  $7 < \eta_v < 8$ LHCv<sub>e</sub> with  $8 < \eta_v < 9$ 

 $10^{0}$ 

 $10^{-2}$ 





 $10^{-1}$ 



## Projections created with the FORSEE tool:

Phys. Rev. D 104, 035012

### https://github.com/KlingFelix/FORESEE

arXiv.org > hep-ph > arXiv:2105.07077

### **High Energy Physics – Phenomenology**

[Submitted on 14 May 2021]

### FORESEE: FORward Experiment SEnsitivity Estimator for the LHC and future hadron colliders

### Felix Kling, Sebastian Trojanowski

We introduce a numerical package FORward Experiment SEnsitivity Estimator, or FORESEE, that can be used to simulate the expected sensitivity reach of experiments placed in the far-forward direction from the proton-proton interaction point. The simulations can be performed for 14 TeV collision energy characteristic for the LHC, as well as for larger energies: 27 and 100 TeV. In the package, a comprehensive list of validated forward spectra of various SM species is also provided. The capabilities of FORESEE are illustrated for the popular dark photon and dark Higgs boson models, as well as for the search for light up-philic scalars. For the dark photon portal, we also comment on the complementarity between such searches and dark matter direct detection bounds. Additionally, for the first time, we discuss the prospects for the LLP searches in the proposed future hadron colliders: High-Energy LHC (HE-LHC), Super proton-proton Collider (SppC), and Future Circular Collider (FCC-hh).

Comments: 11 pages, 3 figures, FORESEE code available at this https URL High Energy Physics - Phenomenology (hep-ph) Subjects: arXiv:2105.07077 [hep-ph] Cite as: (or arXiv:2105.07077v1 [hep-ph] for this version)

Search. Help | Ad

### ∃ README.md

### FORESEE: FORward Experiment SEnsitivity Estimator

### By Felix Kling and Sebastian Trojanowski

arXiv 2105.07077

### Introduction

We present the numerical package FORward Experiment SEnsitivity Estimator, or FORESEE, that can be used to simulate the expected sensitivity reach of experiments placed in the far-forward direction from the proton-proton interaction point. We also provide a comprehensive list of validated forward spectra of various SM species.

### Paper

Our main publication FORESEE: FORward Experiment SEnsitivity Estimator for the LHC and future hadron colliders provides an overview over this package. We recommend reading it first before jumping into the code.

### Tutorials

In the main folder in this repository, we provide tutorials for different LLP models: the dark photon, the dark Higgs, the ALP with W couplings and the up-philic scalar.

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# *F* Physics Motivation | Dark photons

- Dark photons are particularly interesting for FASER as we have fast sensitivity to new regions of phase space
- There is a vast and largely unexplored parameter space
  - Bump hunts" exclude larger ε
  - Mostly fixed target experiments exclude the gray region
  - Astrophysics (supernova, BBN, CMB) exclude at very low ε
- Overall, light, weakly-interacting particles are much less constrained than ~TeV, strongly-interacting particles. Dark Sector models don't give us too much guidance on
- expected mass or coupling strengths.
- Some other intriguing observations...



# **Physics Motivation** | Intrigue...

- Focusing on the mass scale
  - Dark Sector Candidates
  - Anomalies
  - Search Techniques
- We see some interesting things in the ~MeV range



# **Example 2 Physics Motivation** | g-2 The 3.7 $\sigma$ discrepancy between the SM and experiment can be resolved by

- MeV-GeV particles with  $\varepsilon \sim 10^{-3}$ .
- The dark photon is no longer a viable solution
- But other particles with similar masses and couplings are.









# **E** Physics Motivation | He/Be nuclei

- a new particle with mass 17 MeV and couplings ~ 10-3 to 10-4.
- > 2019: A new  $7\sigma$  anomaly in the decays of excited <sup>4</sup>He nuclei can be explained by the same new particle...



Feng, Fornal, Galon, Gardner, Smolinsky, Tanedo, Tait (2016) Feng, Tait, Verhaaren (2020); Batell, Feng, Verhaaren (in progress) See also Zhang, Miller (2020)

2016: A 7σ anomaly in the decays of excited <sup>8</sup>Be nuclei can be explained by





# *F* Physics Motivation | Self-interacting DM

- strongly self-interacting.
- density) as predicted by standard cold dark matter.
- This can be explained by a characteristic dark sector mass scale of ~ 10-100 MeV.



There are indications from small-scale structure that dark matter may be

For example, there appear to be halo profiles that are not as cuspy (high central





## **F** Physics Motivation

**FASER** is probing a very interesting region of phase space

New sensitivity in this region will come even with only a small fraction of Run 3 data.





# *F* **Physics** | Dark portal

### Hidden sector physics:

- New mediating particles, couplings to SM via mixing with SM "portal" operator
- Related to nature of DM (mediator or candidate), baryogenesis, neutrino oscillations...
- Can possibly resolve low-energy experiment anomalies (muon g-2, proton size, Be8)

Typically long-lived particles (LLPs) that travel macroscopic distances before decaying to SM particles



Portal	Coupling
Dark Photon, $A_{\mu}$	$-rac{\epsilon}{2\cos heta_W}F'_{\mu u}B^{\mu u}$
Dark Higgs, $S$	$(\mu S + \lambda S^2) H^{\dagger} H$
Axion, $a$	$rac{a}{f_a}F_{\mu u} ilde{F}^{\mu u},\ rac{a}{f_a}G_{i,\mu u} ilde{G}_i^{\mu u},\ rac{\partial_{\mu}a}{f_a}\overline{\psi}\gamma^{\mu}$
Sterile Neutrino, $N$	$y_N LHN$







# **FASERV** | Emulsion detection

- Emulsion film made up of  $\sim 80 \mu m$  emulsion layer on either side of 200  $\mu m$  plastic
- Emulsion gel active unit silver bromide crystals (dia. 200nm)
- Charged particle ionization recorded and can be amplified and fixed by chemical development of film
- Track position resolution ~50nm, and angular resolution ~0.35mrad
  - But no time resolution!





Josh McFayden Birmingham I





## **FASERv** | Neutrino flux estimates

- Checking three simulations.
  - FLUKA (by F. Cerruti's group)
  - BDSIM (by H. Lefebvre, L. Nevay)
  - RIVET-module (by F. Kling)

### Differences between generators have been checked with the same propagation model (RIVET-module)



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# **FASERV** Interface to FASER

- To connect muon tracks from  $\nu\mu$  interactions for charge identification etc.
  - Interface tracker (IFT) with 3 layers of silicon strip detector. A copy of FASER tracker station.
  - Veto station consists of 2 scintillator layers with 2 cm thickness. >99.99% veto efficiency for a charged particle coming from upstream of FASER
  - Construction of the IFT will start in January 2021. Installation at FASER site is planned in fall 2021



Double Scintillator layer

Single Scintillator layer in front of absorber

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# **Example 2 Conditions** | Beam backgrounds

- expected backgrounds.
  - IP1 collisions (shielded by 100m rock)





# **Example 2** Conditions | Beam backgrounds

In situ measurements using emulsion detectors and TimePix BLM in TI12 in 2018 confirm expected particle flux, and correlation with IP1 luminosity.





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## **Example 2 Conditions** | Beam backgrounds

The FLUKA simulation tracks particle production, deflection, and energy loss with a detailed model of the geometry of the LHC tunnels, including the LHC material map and magnetic field layout. The simulation includes three potential sources of background at the FASER location:

- Particles produced in the pp collisions at the IP or by particles produced at the IP accompanying muons that interact further downstream, e.g., in the TAN neutral particle absorber.
- Particles from showers initiated by off-momentum (and therefore off-orbit) protons hitting the beam pipe in the dispersion suppressor region close to FASER.
- Particles produced in beam-gas interactions by the beam passing FASER in the ATLAS direction (for which there is no rock shielding).

Always co-linear with

 $-10^5 \rightarrow 10$  with veto

Minor



# **Example 7** Conditions | Radiation levels

- - less than 5 x 10<sup>-3</sup> Gy/year
  - less than 5 x 10<sup>7</sup> 1 MeV neutron equivalent fluence/year
- FASER detector does not need radiation hard electronics



Radiation level predicted to be very low in TI12 due to dispersion function of LHC at TI12. Measurements using BatMon radiation monitor in 2018 confirm FLUKA expectations:













- interaction models, grounded on LHC data
- production peaks at  $pT \sim \Lambda_{QCD}$
- enormous event rates  $N \sim 10^{15}$  per bin

- at  $pT \sim \Lambda_{QCD}$
- rates highly suppressed by  $\epsilon^2 \sim 10^{-10}$
- still rates N~10<sup>5</sup> per bin: LHC could be dark a photon factory

- rates suppressed by decay requirements
- still rates N~100 signal events within 20cm of beam collision axis

|9



## **E Overview** | LLP production modes









## *E* Overview | Dark photon reach







# *F* Overview | Dark photon reach

For lower lifetime the number of signal events becomes exponentially suppressed once the A' decay length drops below the distance to the detector

Combining dependence in both production rate and decay width, total number of signal events in the detector scales as  $\varepsilon^4$ 





### **Target scenarios** | Dark Photon





### **Target scenarios** | Dark Higgs



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### *Target scenarios* | Dark Higgs







### **Example 7 Target scenarios** | ALP



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### **Example 7 Target scenarios** | ALP







### *E* Civil Engineering in TI12 Trench

- To be aligned with the line-of-sight (LOS) in the vertical direction a shallow (<50cm deep) trench was needed in TI12
- Drain shallower than shown on historical drawings
  - Provided opportunity to increase trench depth parallel to LOS
- Plan area increased to allow more space for FASERv
- Trench strengthening
  - Improved rock characteristics enabled removal of steel frame
  - Less complex site works and better ground conditions enabled increased depth







80	[*10 <sup>3</sup> kN/m <sup>2</sup> ]
Lundin.	2.00
	0.00
X	-2.00
	-4.00
2	-6.00
$\geq$	-8.00
$\rightarrow$	-10.00
$\leftarrow$	-12.00
A.	-14.00
$\times$	-16.00
XI.	-18.00
Å.	-20.00
to.	-22.00
5	-24.00
7	-26.00
-	-28.00



### **F** Beam offset





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# **F** Modelling uncertainties







# Energy threshold







# **FASER2** | Magnet

Rectangular magnet 3 x 0.5 x 2m 2 T bending in horizontal direction



## Hide Otono





# Example 2 FASER2 | Magnet

- Circular magnet 1m radius, 0.5m high
  - More bending power in centre highest energy
  - Less stored energy
- 2 T bending in horizontal direction









## Hide Otono



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	-	- )	(=)	10	001
	-	- ,	(=)	150	oor
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				-	
		 		Τ	
		,		4,000	•



## **FASER2** Magnet **Cost estimation from TOSHIBA**

## **Roughly 10 MCHF** based on their experience on SAMURAI

ltem	Unit	Value	Remark
Magnet		Dipole magnet	
Magnetic field	Т	2	
Magnetic path length	T۰m	4.7	Rough estimati SAMURA
Stored energy	MJ	15	
Magnetic pole gap distance	mm	880	same a SAMUR
Magnetic pole radius	mm	2000	circular p
Coil		Solenoid	
Total weight	ton	400	











# **FASER2** | Tracker design and costing Based on SciFi detector installed in LHCb in LS2.

- SiPM+scintillating fibre design
- Fibres 250um diamater => 80um resolution.
- Each module consists of a mat of 4 fibres, with >99% efficiency.
- Costing done by scaling LHCb detector to the FASER2 design, and includes readout.
- Cost could be reduced by re-using tooling from LHCb if relevant institutes were involved.







# **FASER2** | SciFi Tracker

### The upstream tracker

- 6 vertical + 2 horizontal modules makes up a station.
- 3 stations.



- The stations should be relatively rotated e.g. 1 degree to maximize performance for multi tracks etc.
- Cost: ~3.8M CHF





## **FASER2** | Dual-readout calorimeter Existing dual-readout prototypes for Higgs factory detectors EM prototype exists, construction of hadronic-size prototype ongoing Costing based on HiDRa "hadronic size" prototype - INFN

- - 65x65x250 cm (presentation)
  - Aiming for 2023 construction and test heam











### Iacopo Vivarelli





# **FASER2** | Calorimeter design

## Fully segmented design

- Perpendicular crossing of EM layers
- Don't need dual readout no Cherenkov fibres

## Costing Option 3-5M Euros

Depending on readout and granularity













# **FASER2** | Calorimeter design

- Possibility to reuse old LHCb Preshower and Scintillating Pad Detector for FASER2 Calo
  - Active part is made up of scintillator pads with wavelength shifter embedded.
  - Pad size depends on the location and are 12 cm x 12 cm, 6 cm x 6 cm and 4 cm x 4 cm.
  - Pads are supported on "super modules" with an active area of about 1 m x 5.8 m





